

A Tour of Lightning With HAMMA



Phillip M. Bitzer



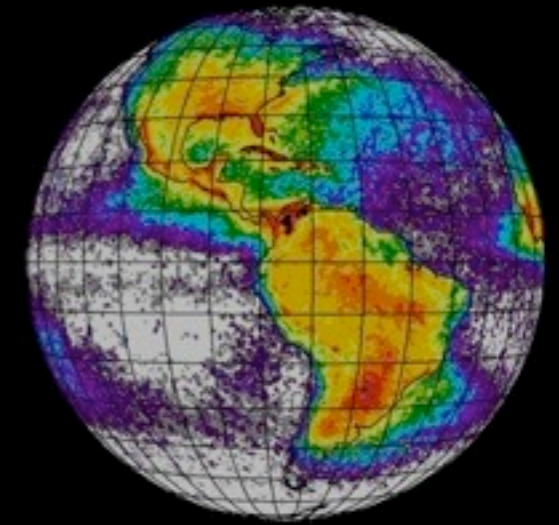
The HAMMA Team:

Hugh Christian, Mike Stewart, Jeff Burchfield, Scott Podgorny, David Corredor,
Veronica Franklin, John Hall, and Evgeny Kuznetsov

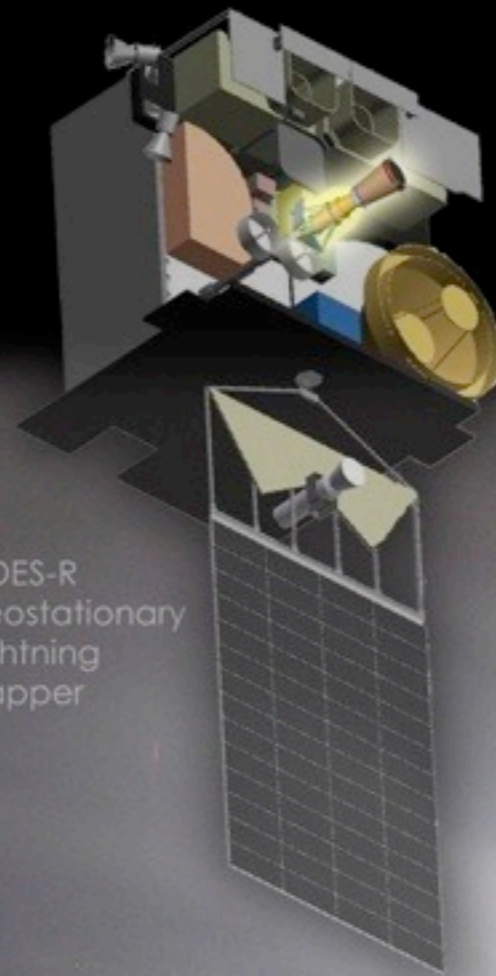


**Special thanks to:
Monte Bateman and Dennis Buechler**

**Dan Cecil, Henry Everitt and Martin Heimbeck, Chris and Elise Schultz,
Brandon and Anne Marie Strickland, and Winfred Thomas Agricultural Research Station**



Fundamental question for *scientific* validation:

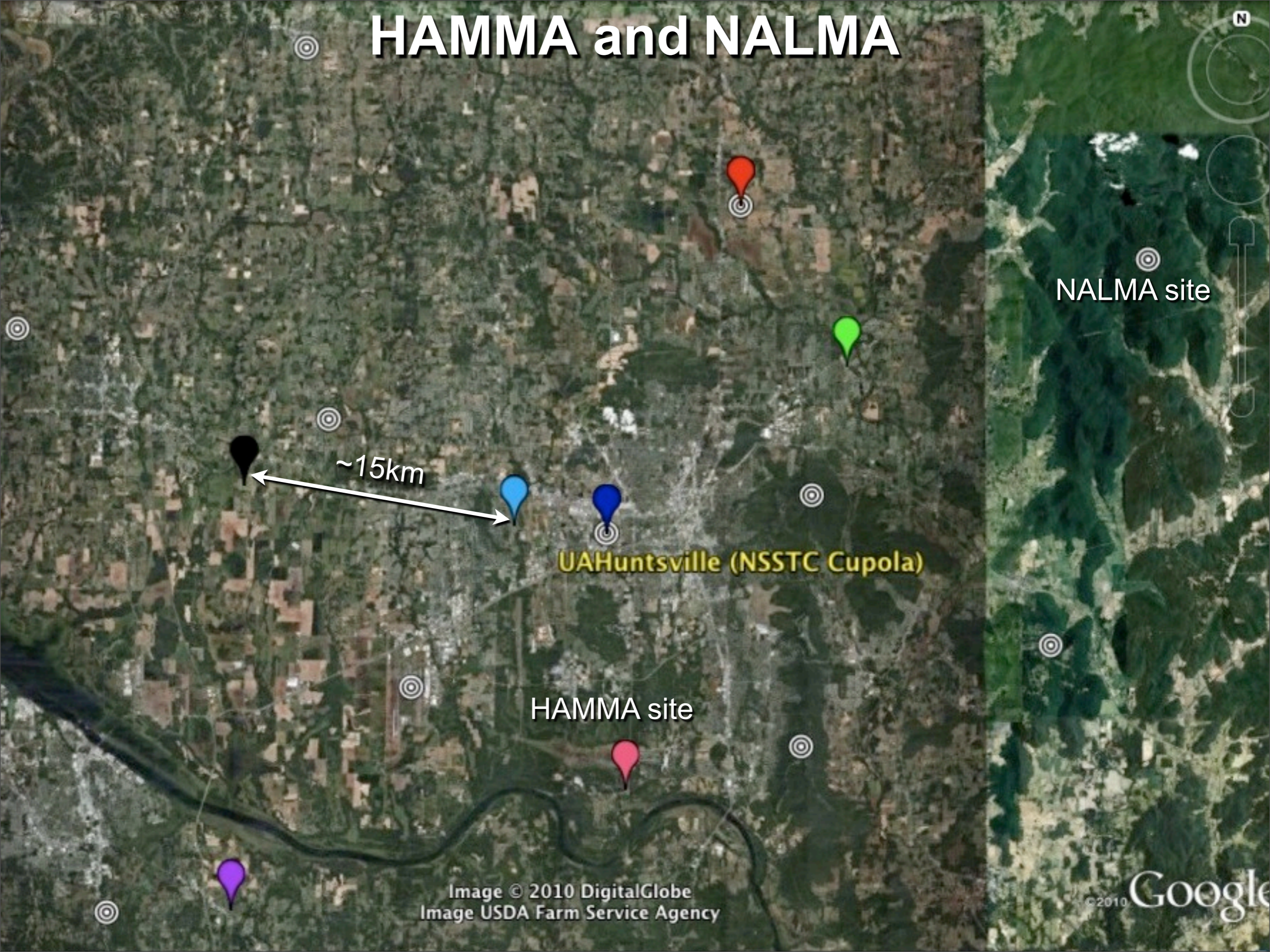


GOES-R
Geostationary
Lightning
Mapper

What is the best way from the ground to characterize what LIS “sees?”



HAMMA and NALMA



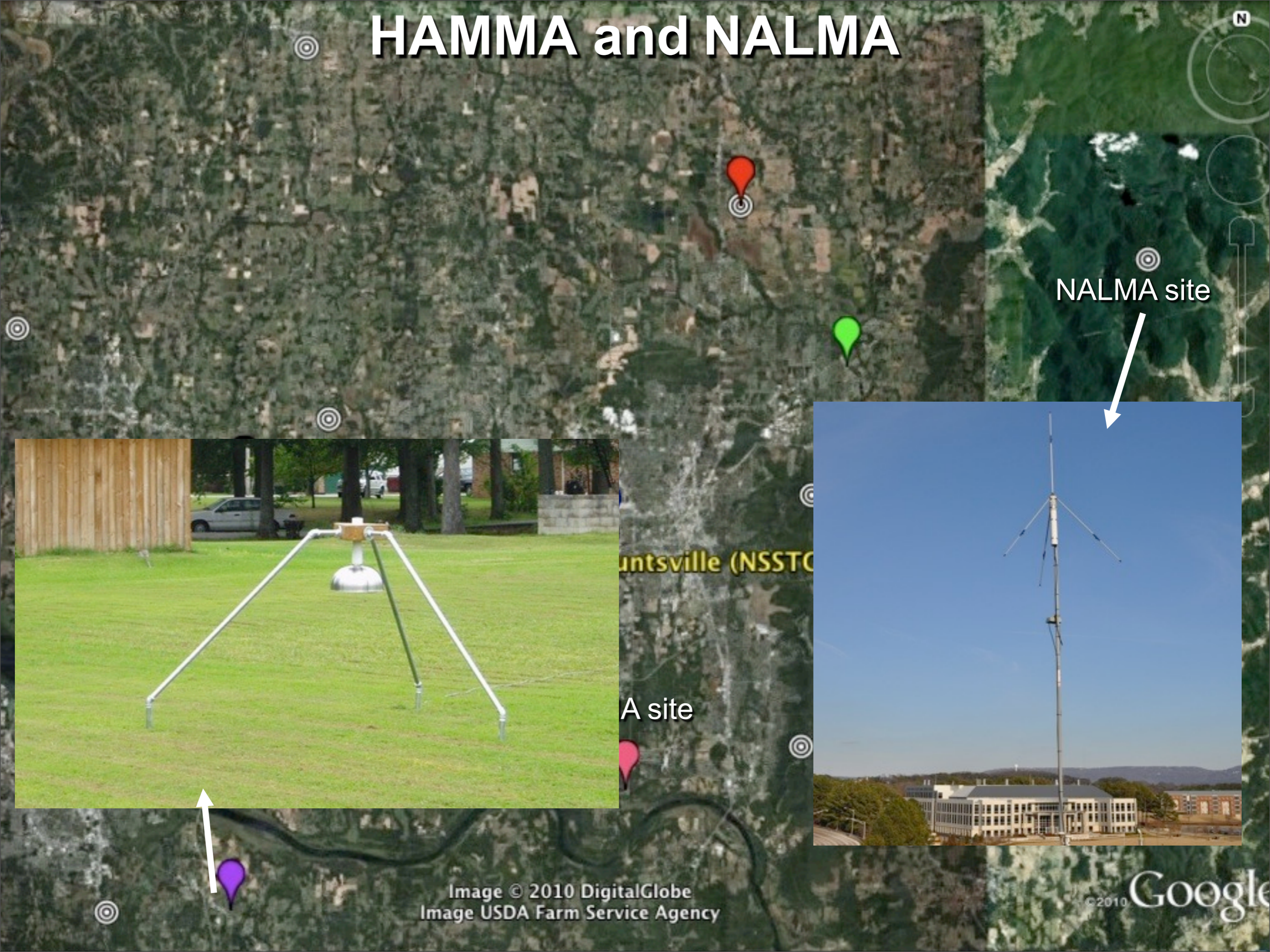
NALMA site

UAHuntsville (NSSTC Cupola)

HAMMA site

~15km

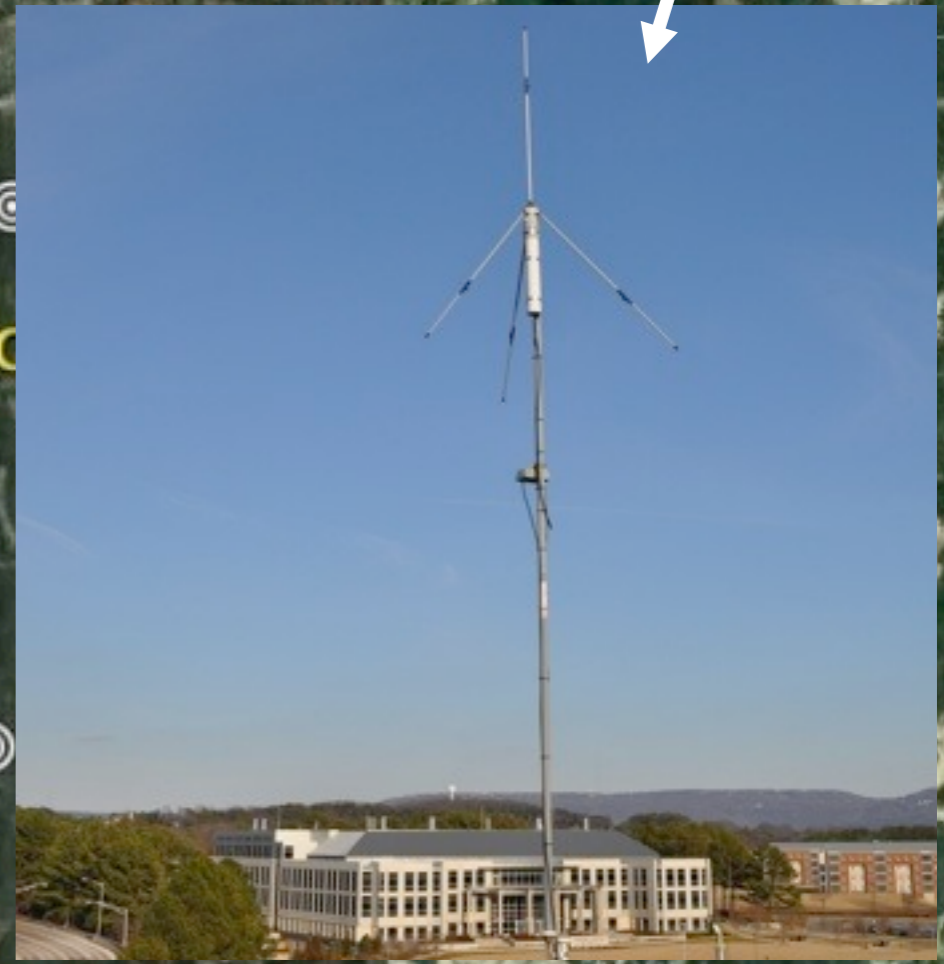
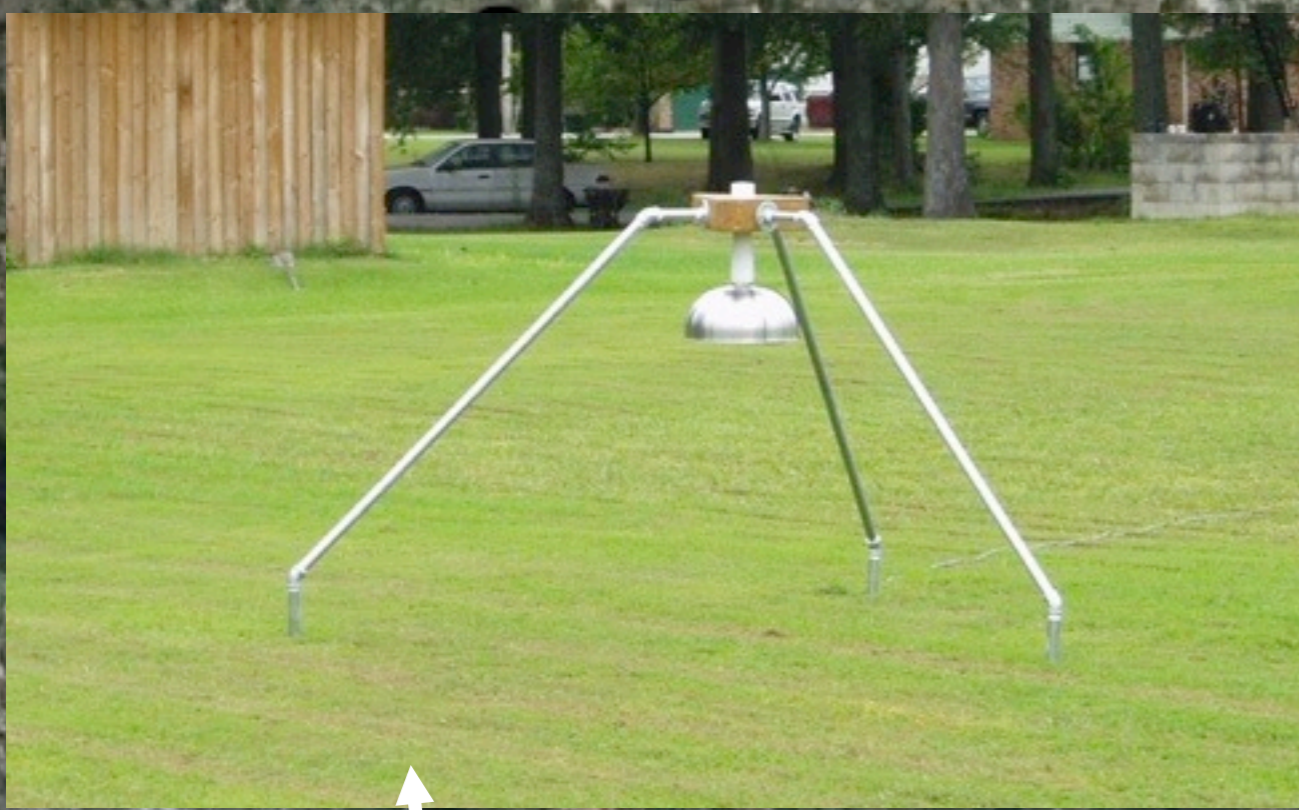
HAMMA and NALMA



NALMA site

untsville (NSSTC

A site



HAMMA

Huntsville Alabama Marx Meter Array

NALMA

North Alabama Lightning Mapping Array

HAMMA

Huntsville Alabama Marx Meter Array

Wideband ~1Hz-500kHz
Sample Rate: 1MHz

NALMA

North Alabama Lightning Mapping Array

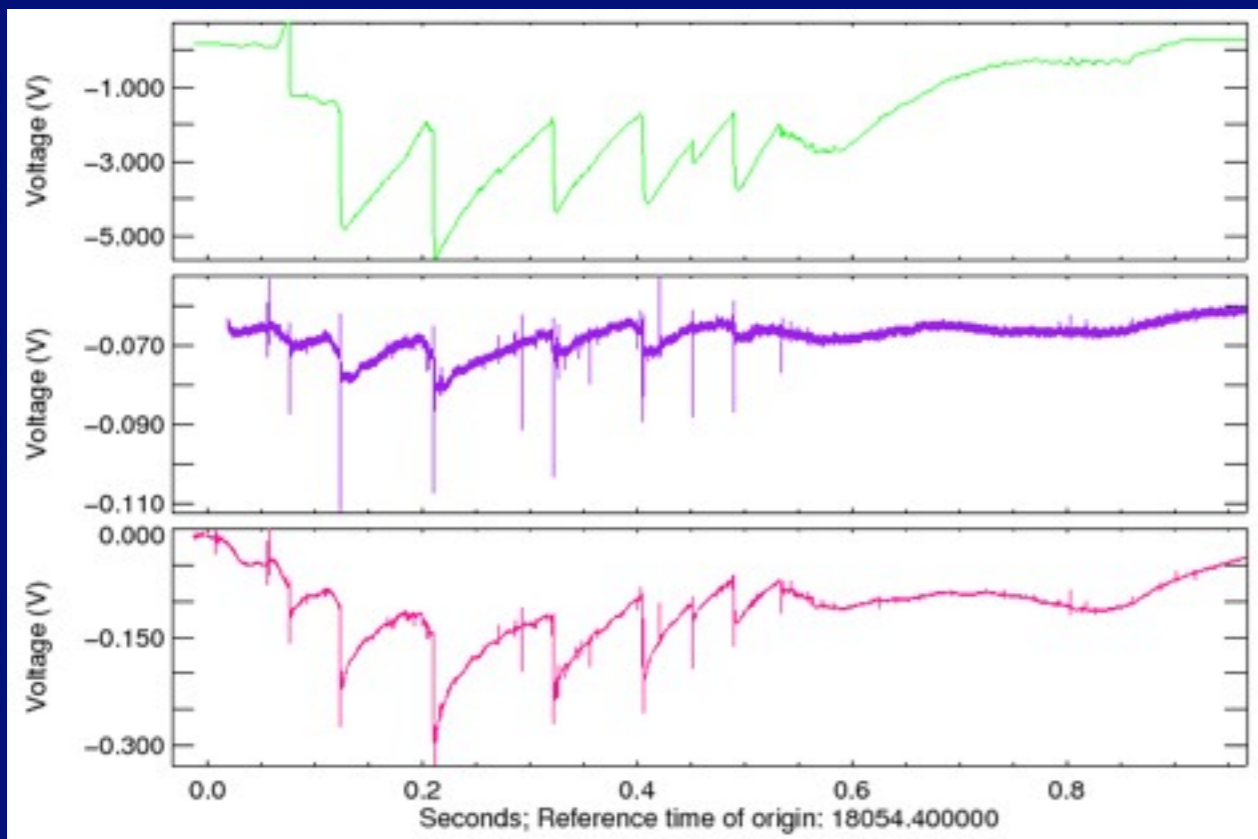
Narrowband VHF ~80 MHz
Sample Rate: 20MHz

HAMMA

Huntsville Alabama Marx Meter Array

Wideband ~1Hz-500kHz
Sample Rate: 1MHz

Can identify polarity/type of discharge

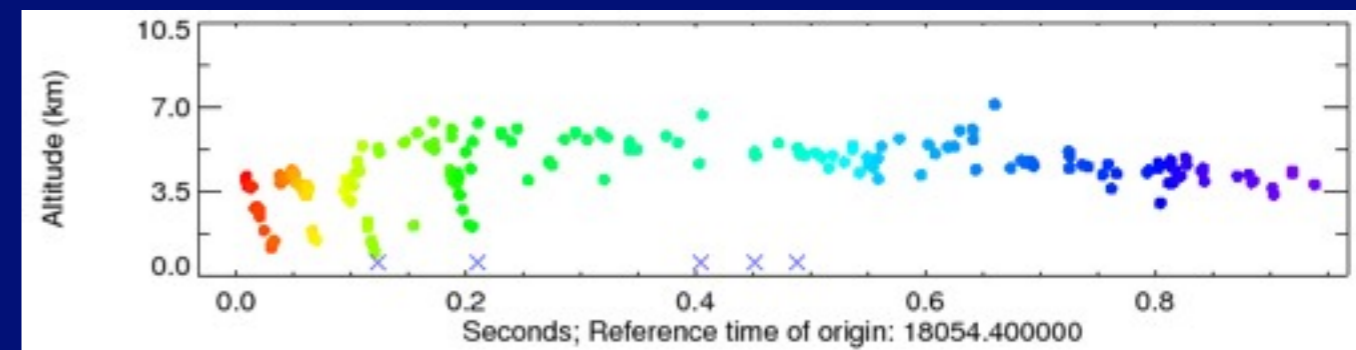
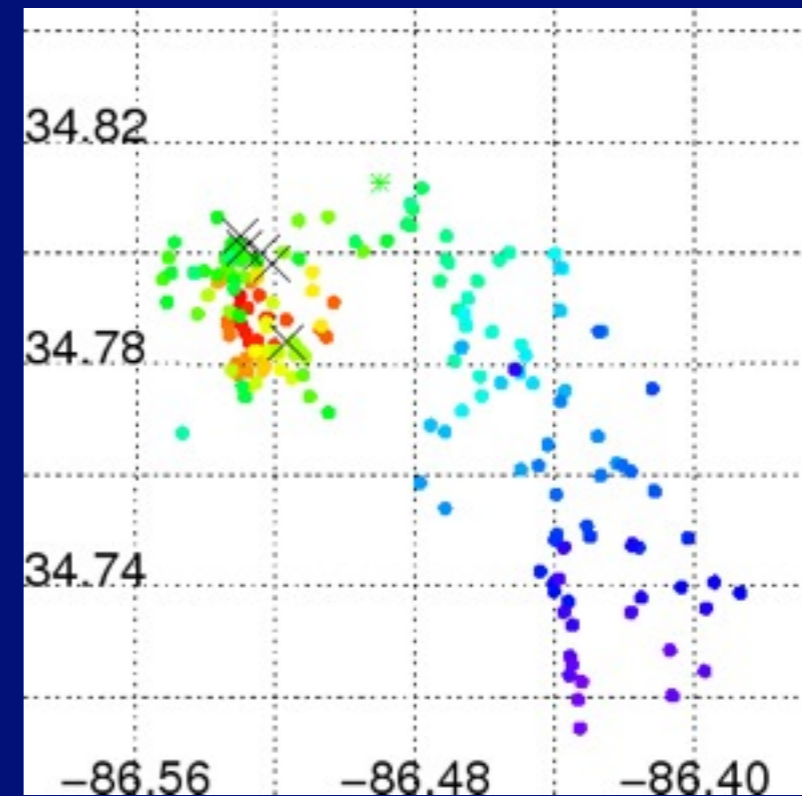


NALMA

North Alabama Lightning Mapping Array

Narrowband VHF ~80 MHz
Sample Rate: 20MHz

Provides accurate "maps" of lightning



HAMMA

Huntsville Alabama Marx Meter Array

Wideband ~1Hz-500kHz
Sample Rate: 1MHz

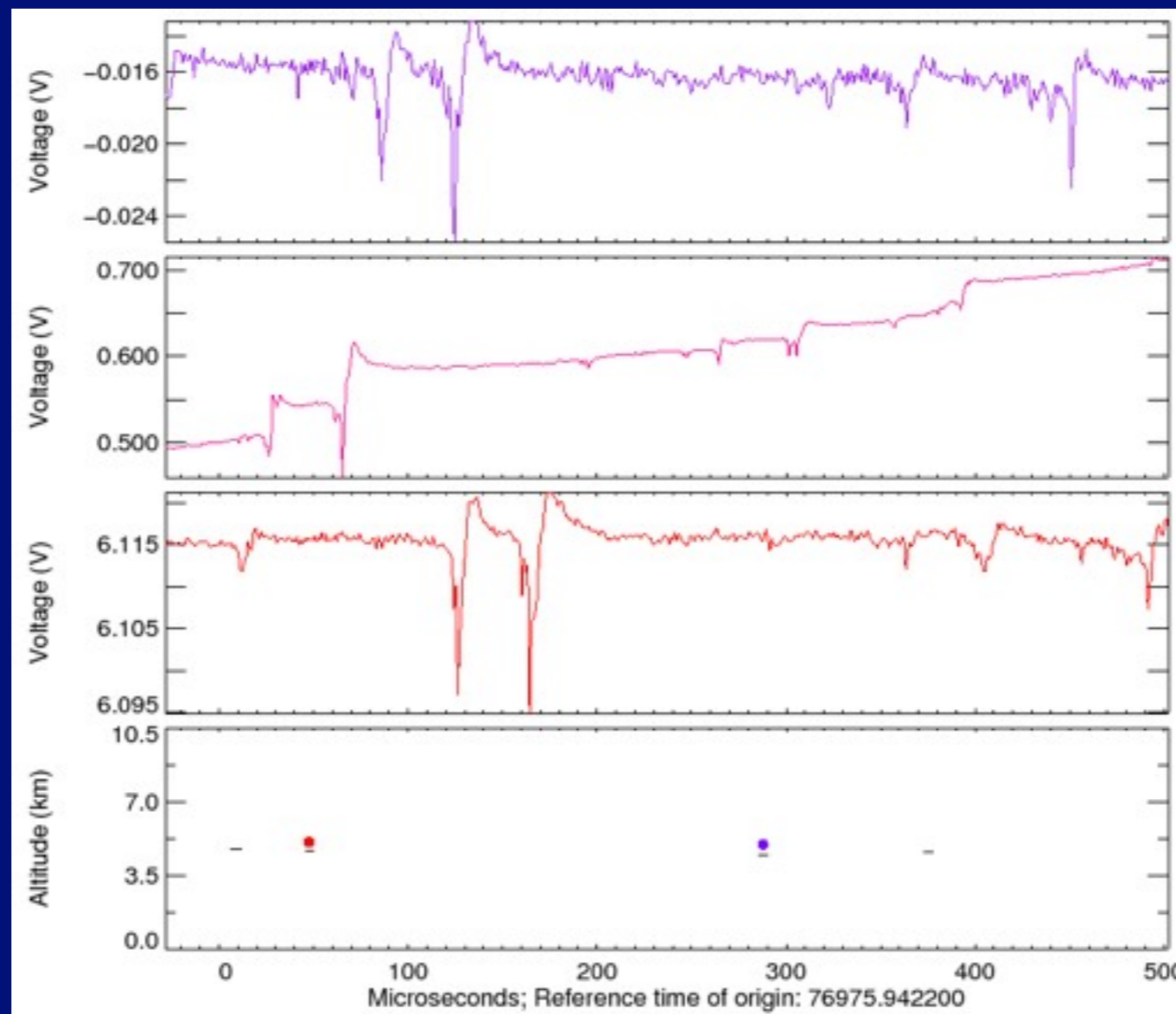
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NALMA

North Alabama Lightning Mapping Array

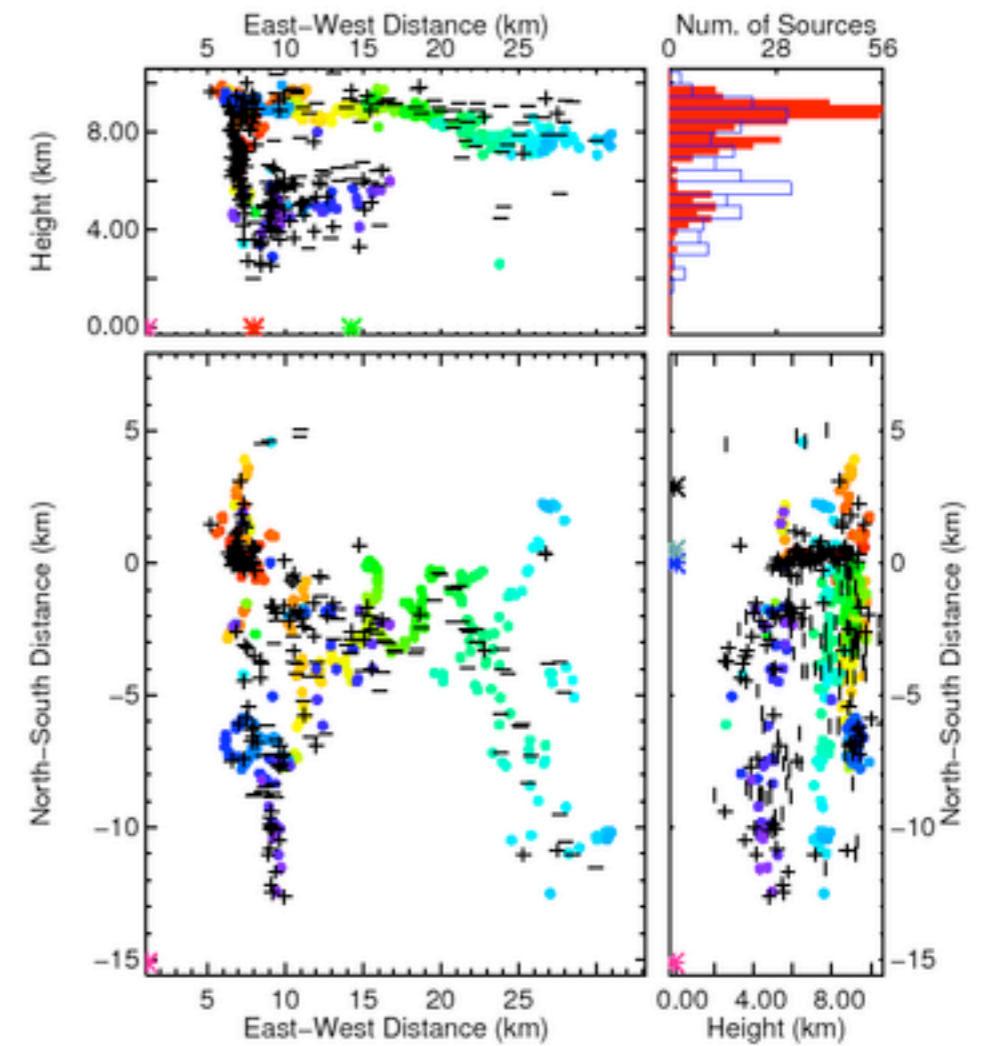
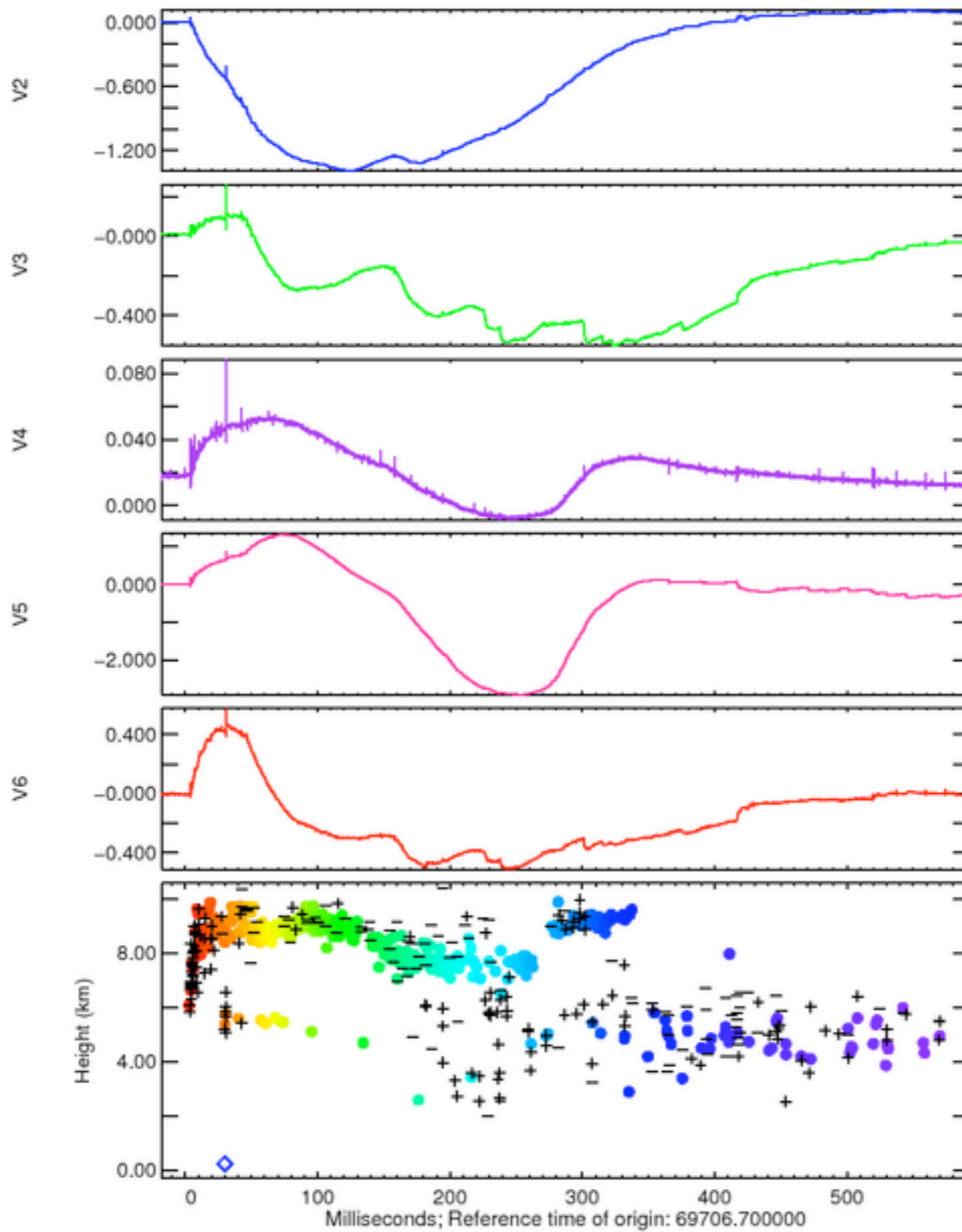
Narrowband VHF ~80 MHz
Sample Rate: 20MHz

Provides accurate “maps” of lightning



Both arrays can locate sources via time-of-arrival; HAMMA can also provide energetic information

2010/06/25 19:21:46

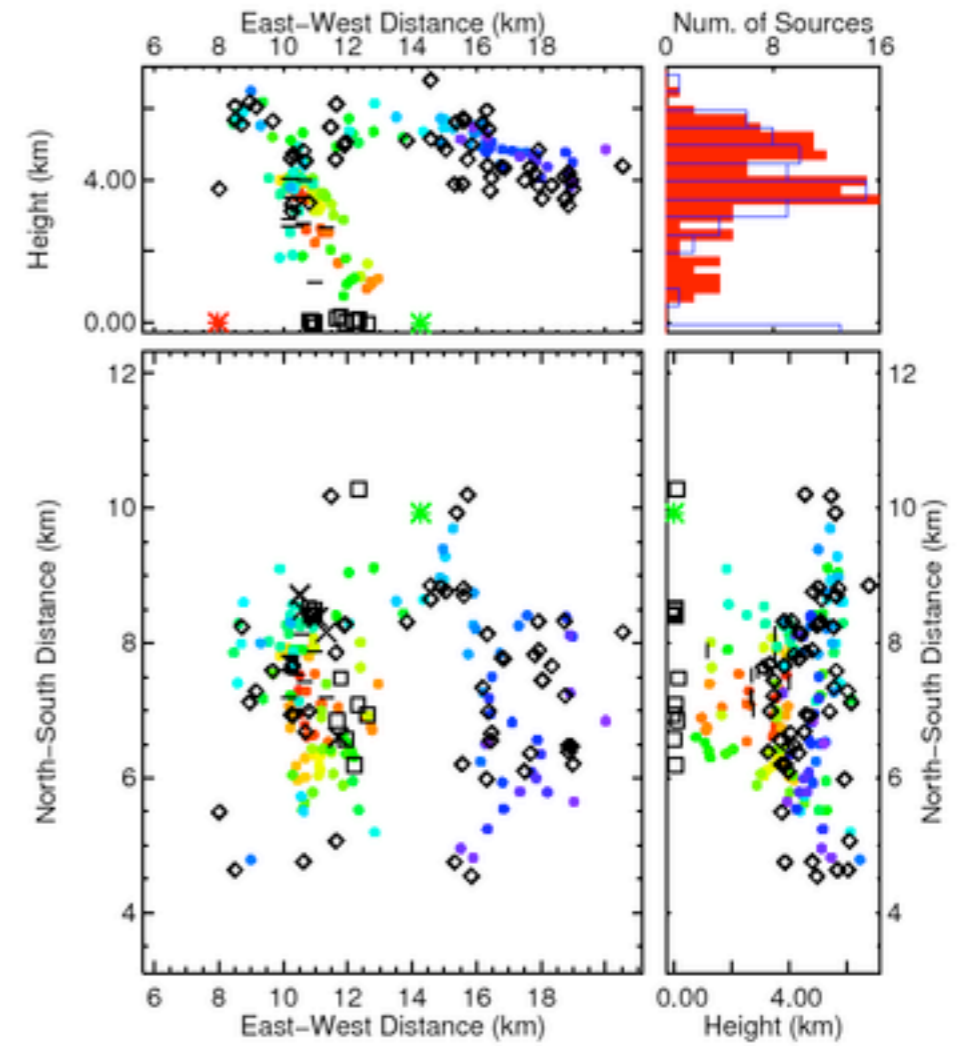
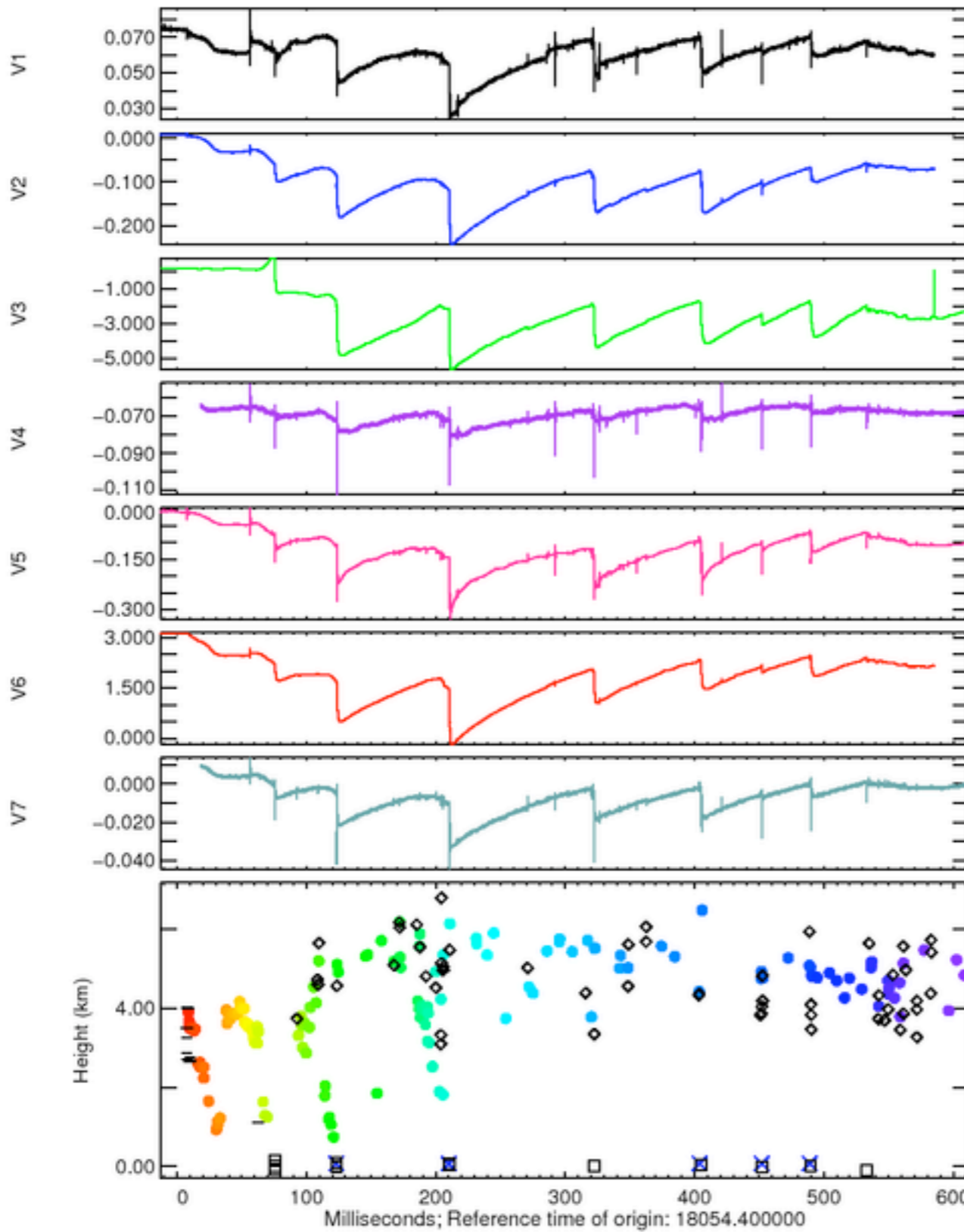


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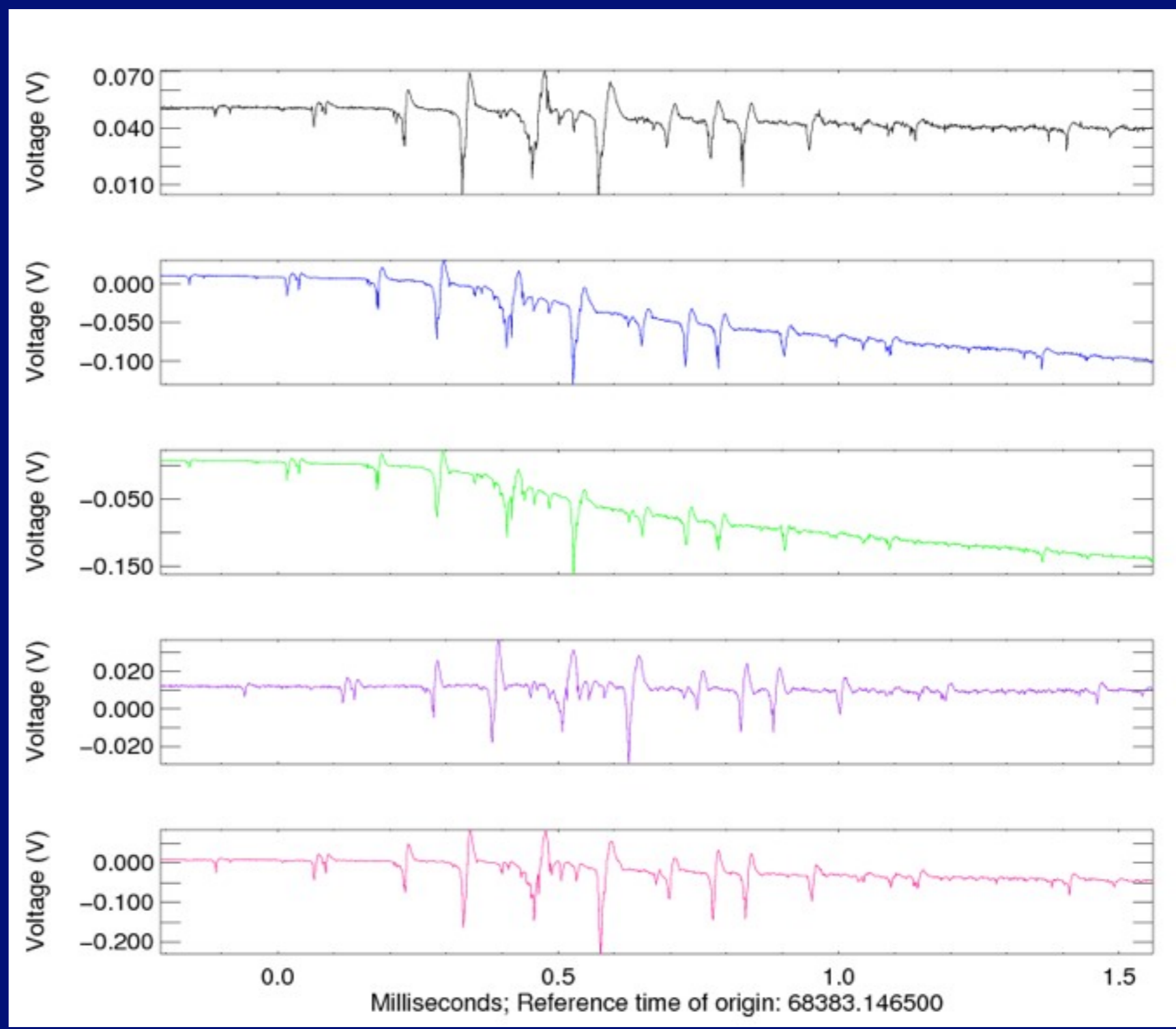
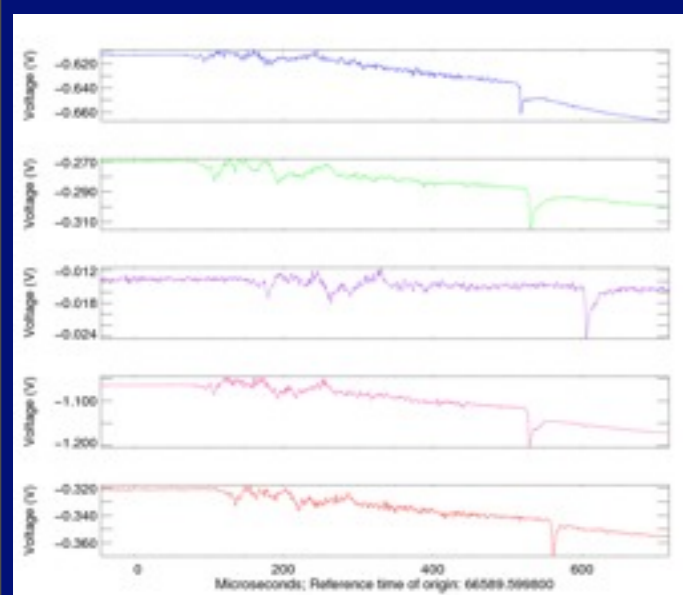
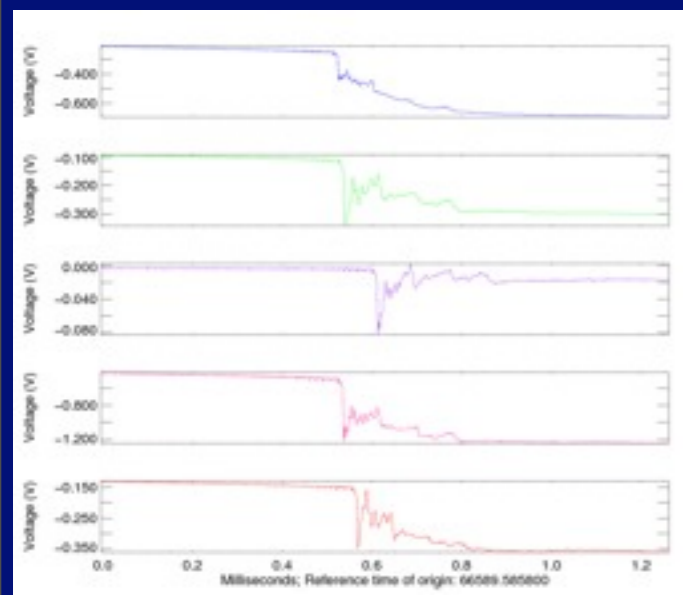
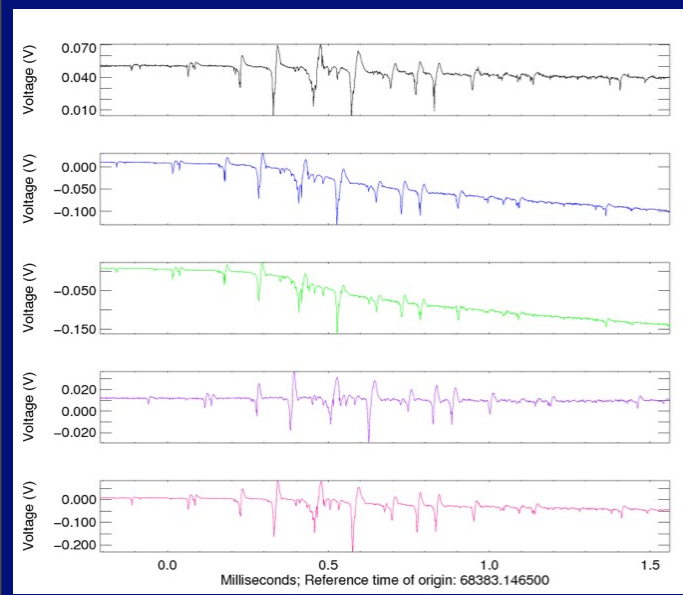
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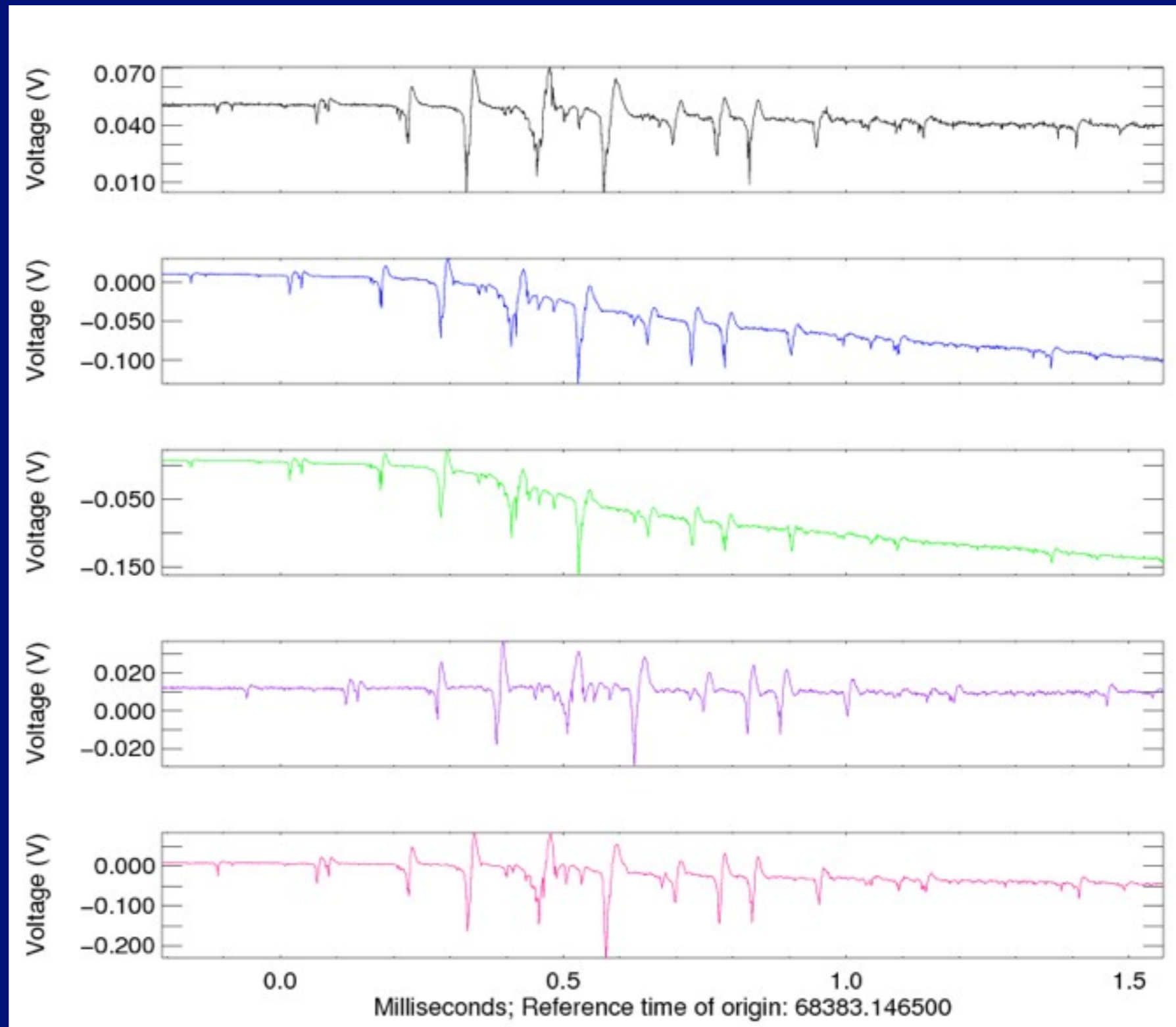
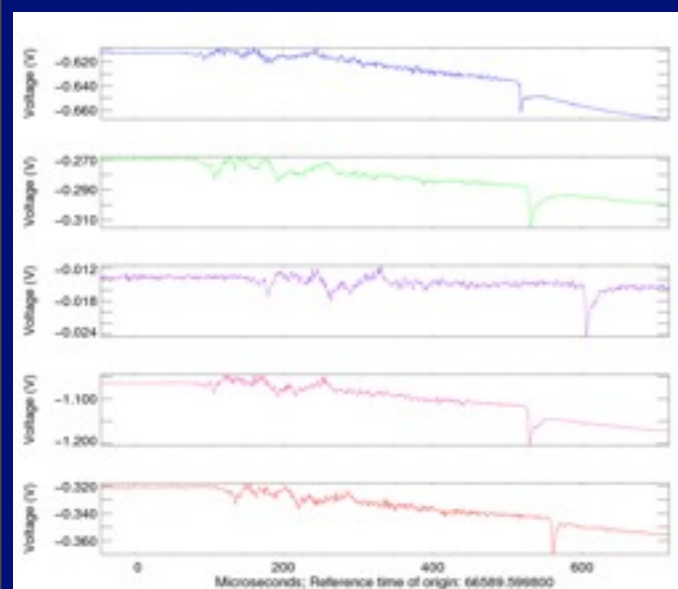
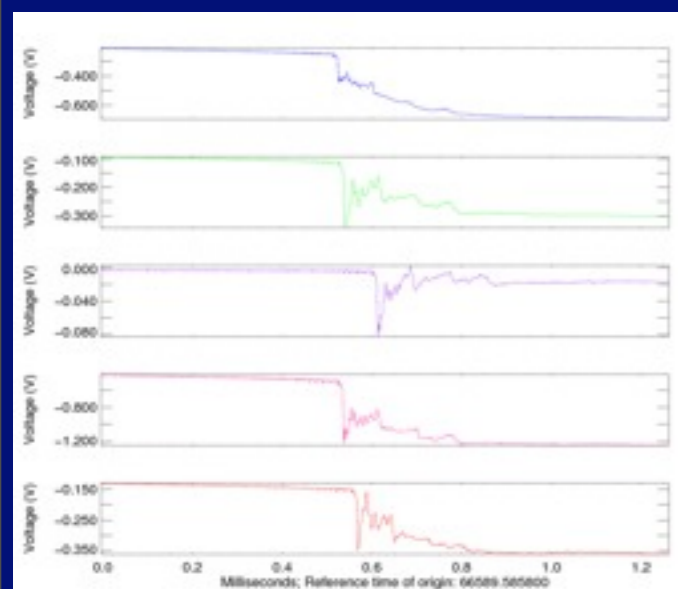
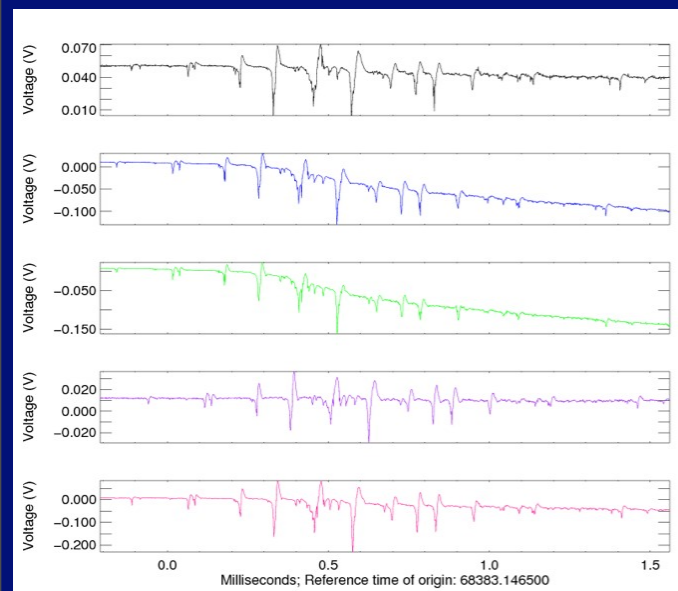


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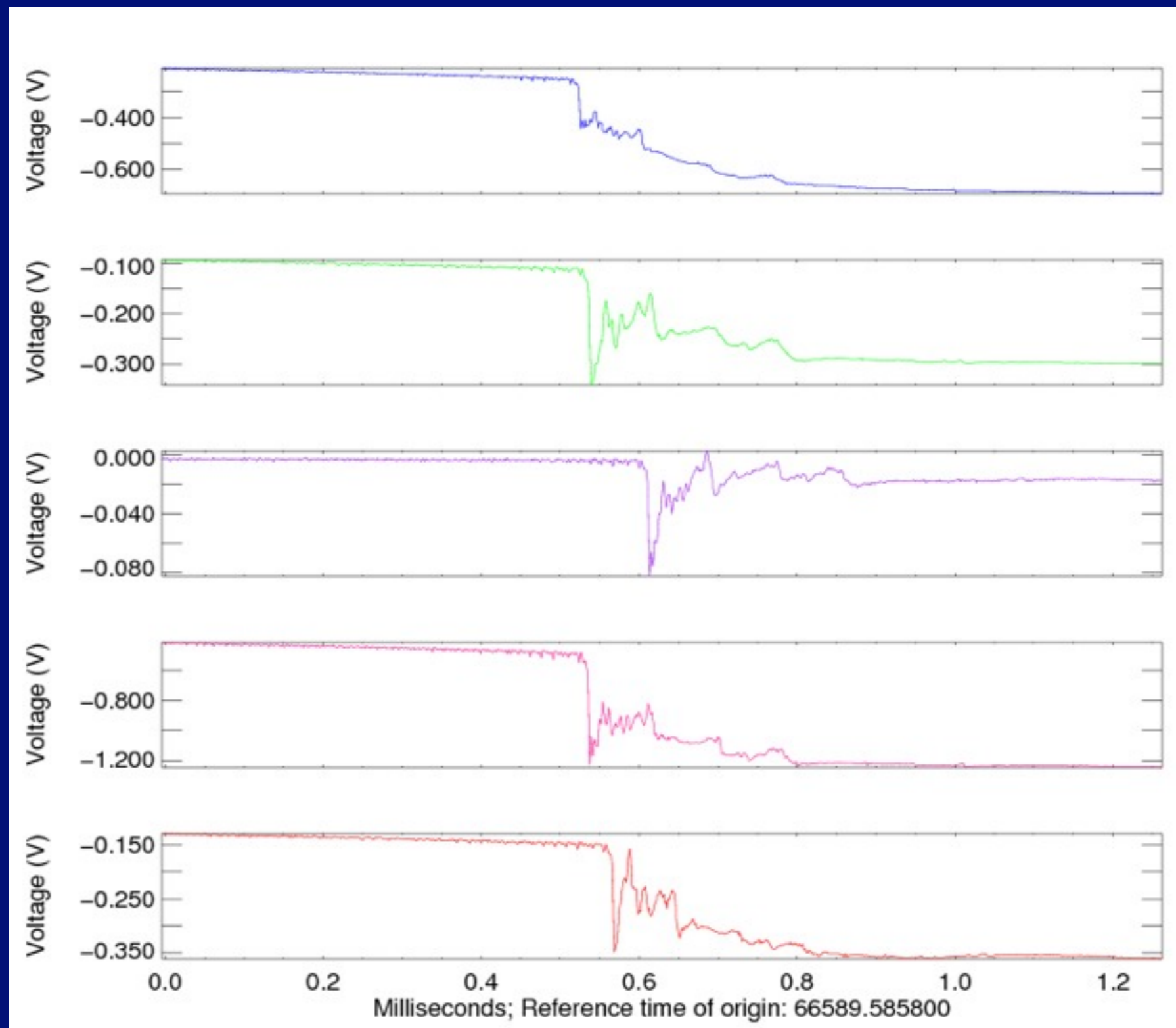
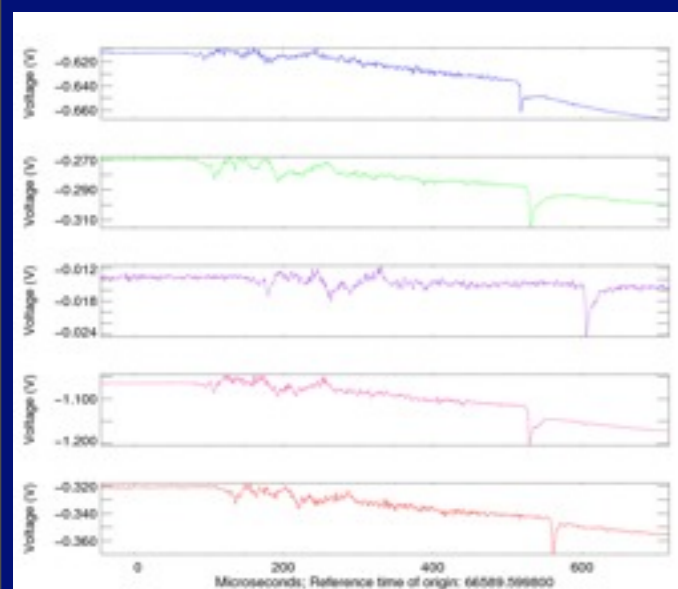
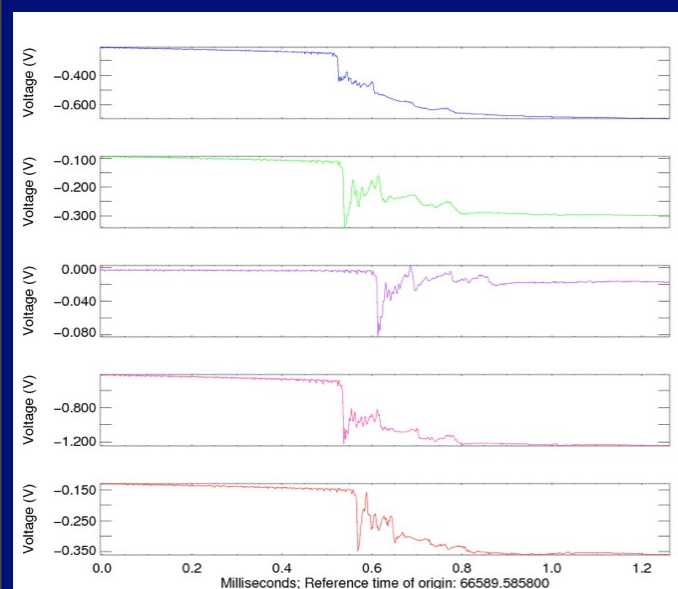
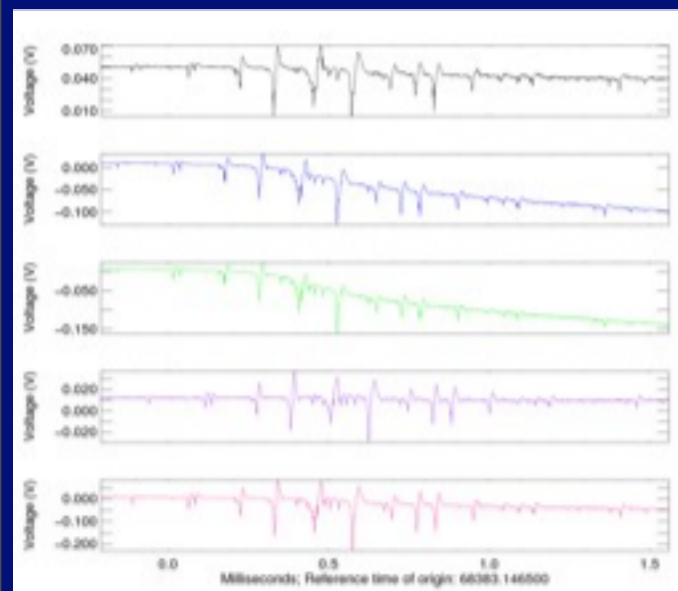
Stop time: 18055.012631

Time Elapsed: 0.625304

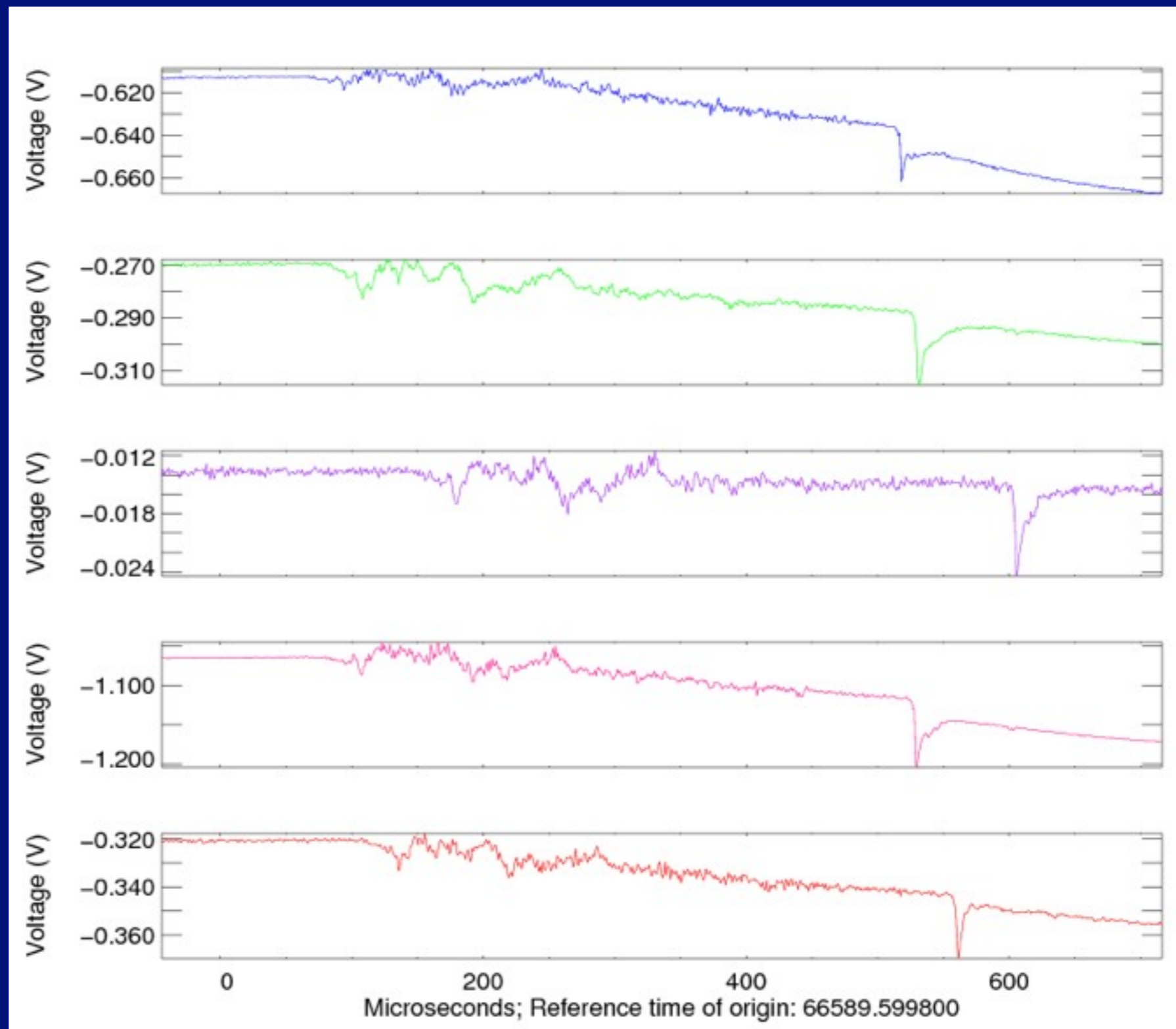
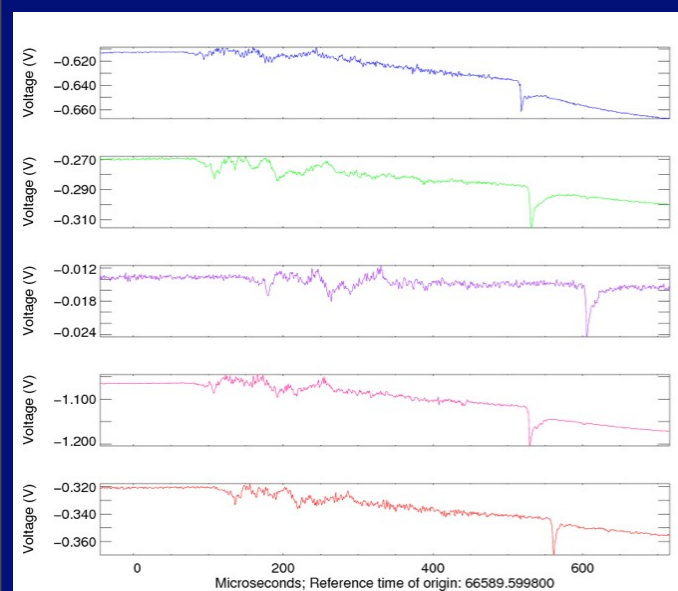
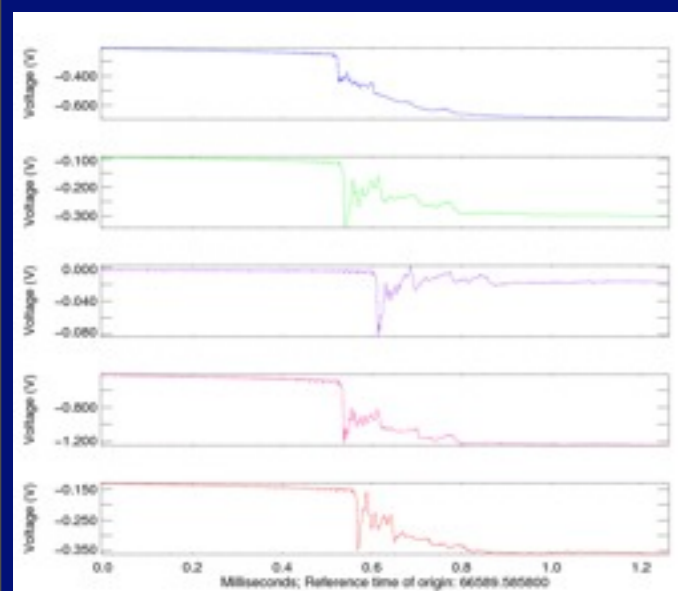
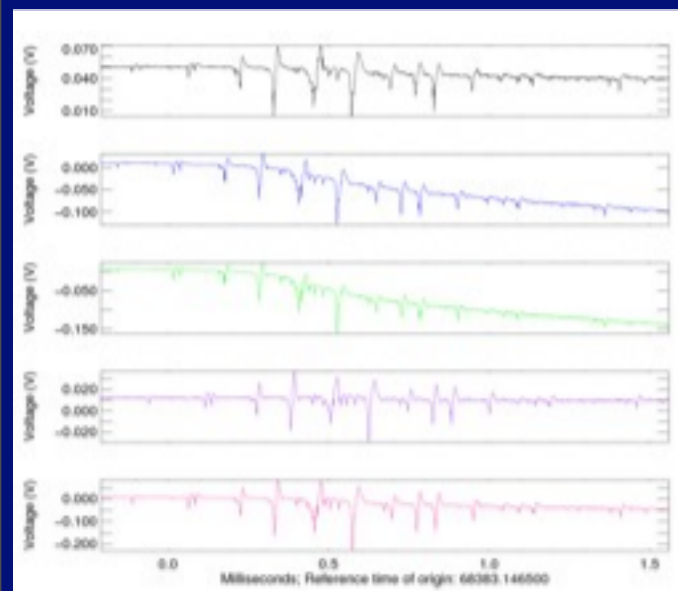




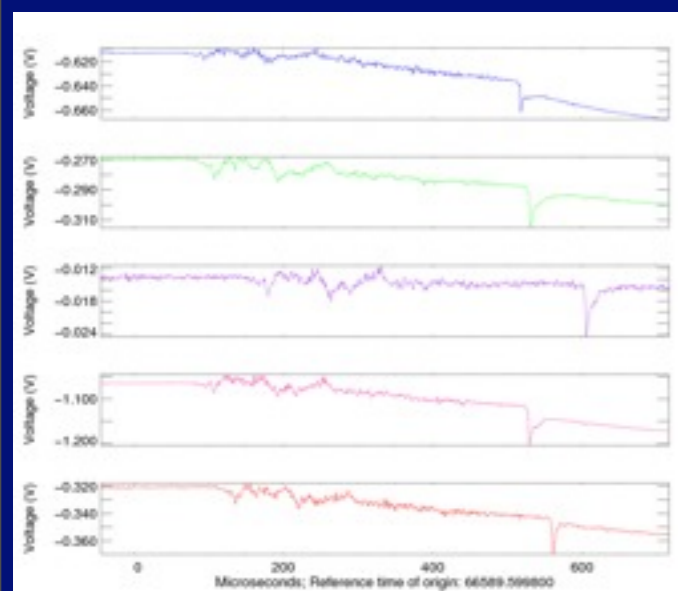
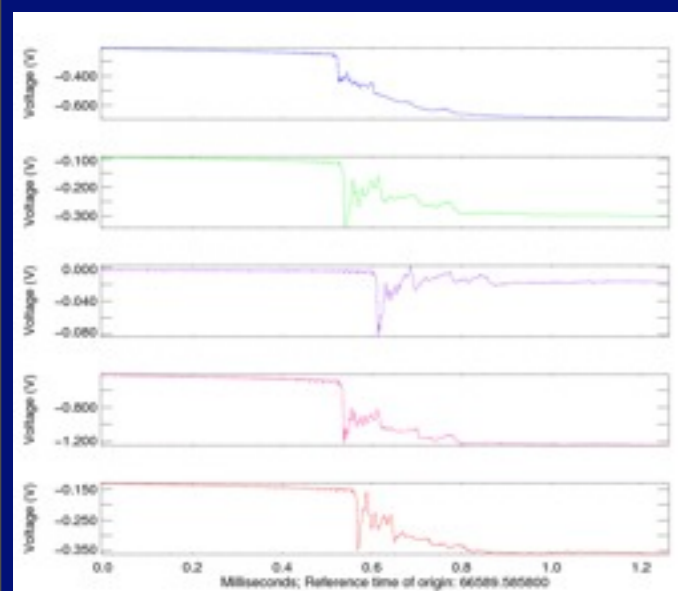
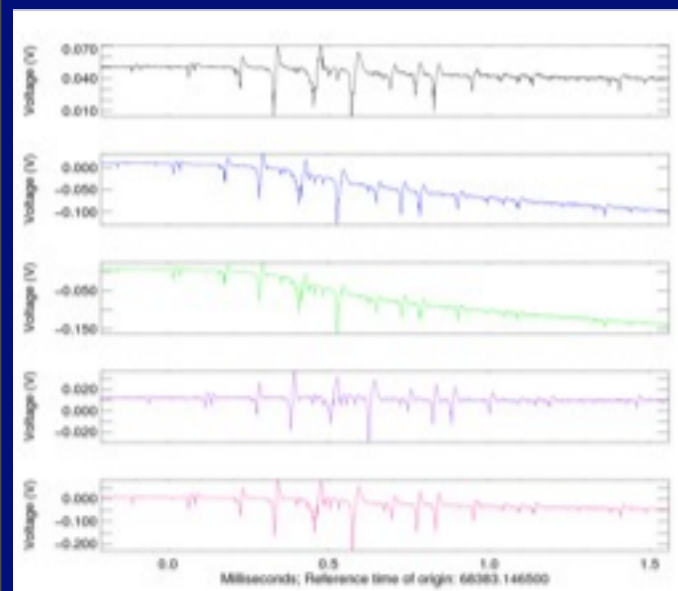
preliminary breakdown pulse train



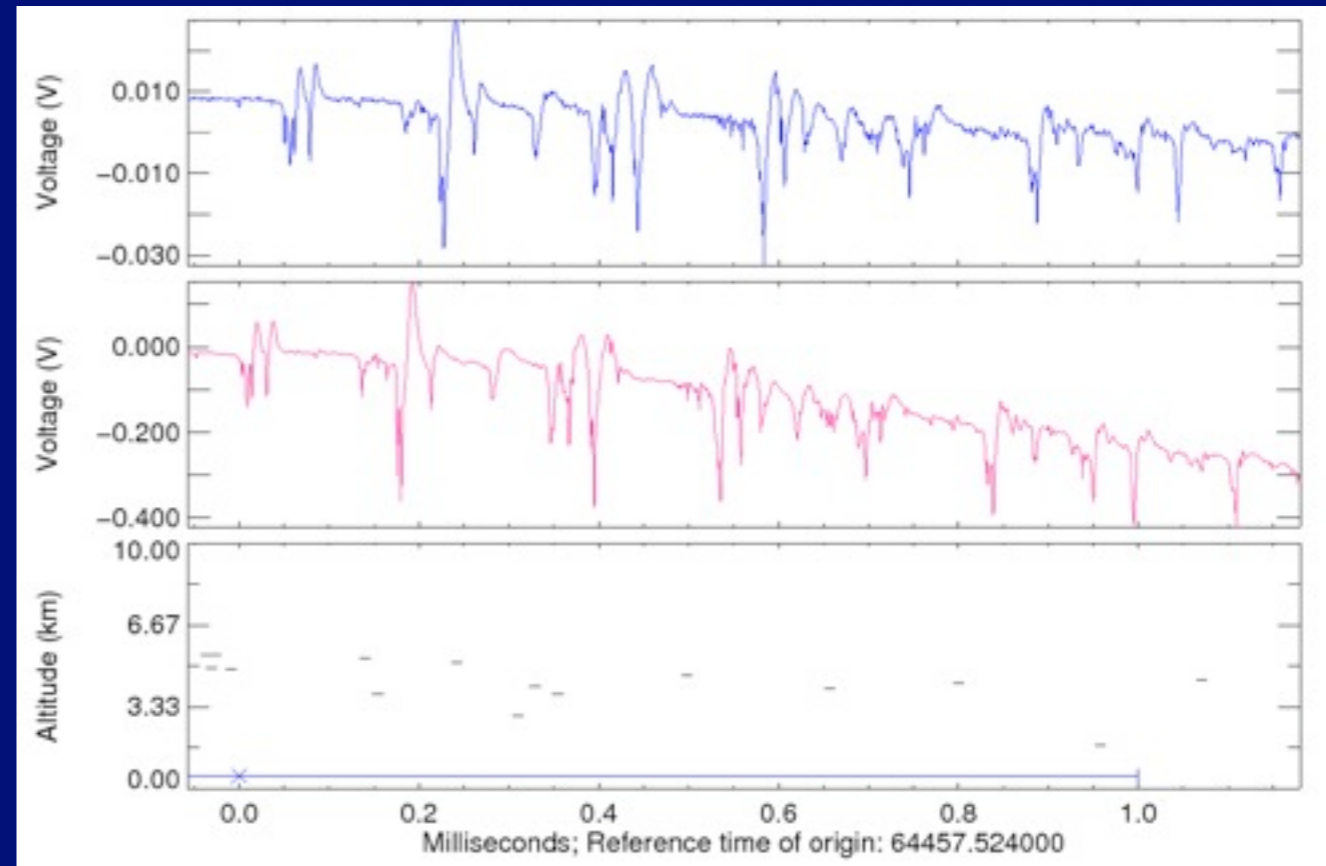
early return stroke



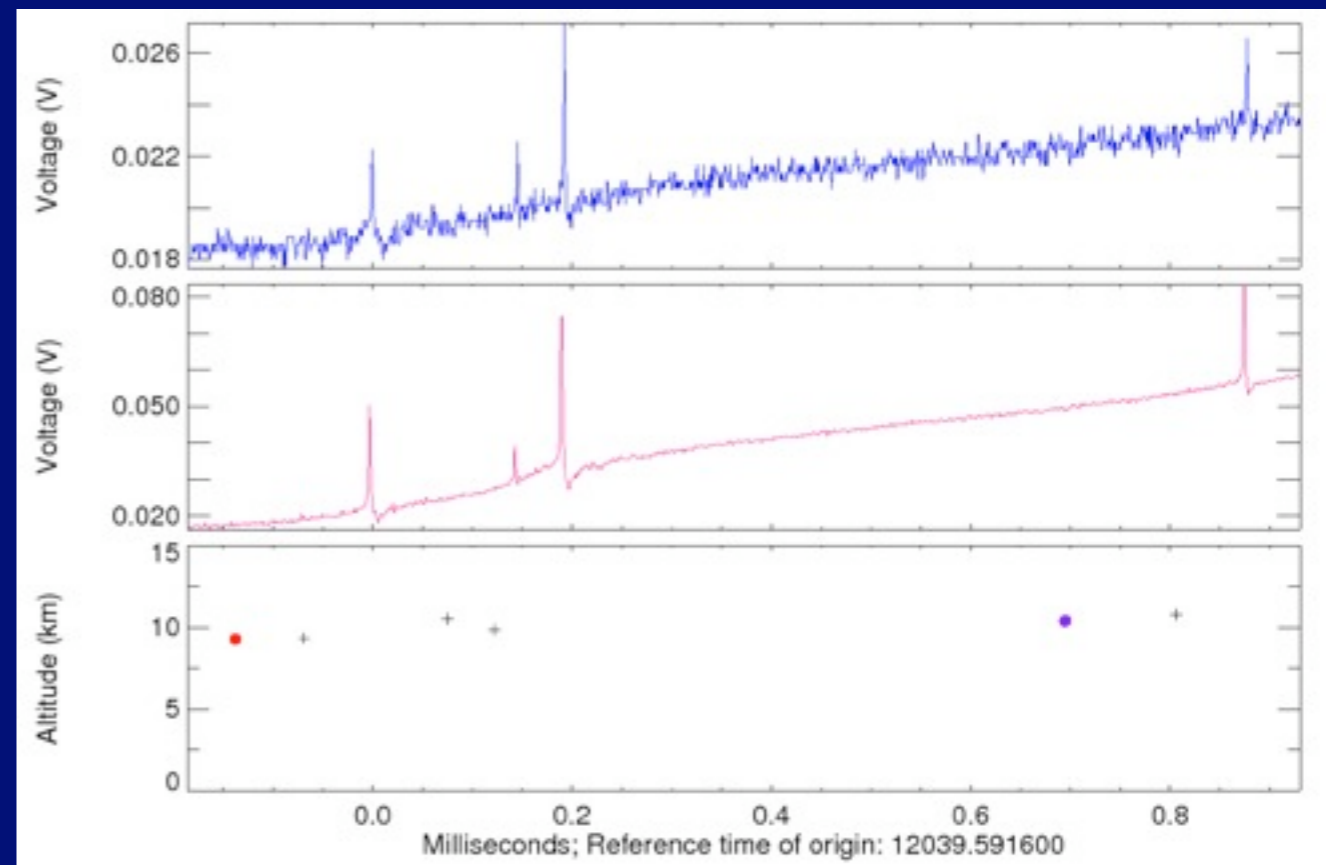
subsequent return stroke



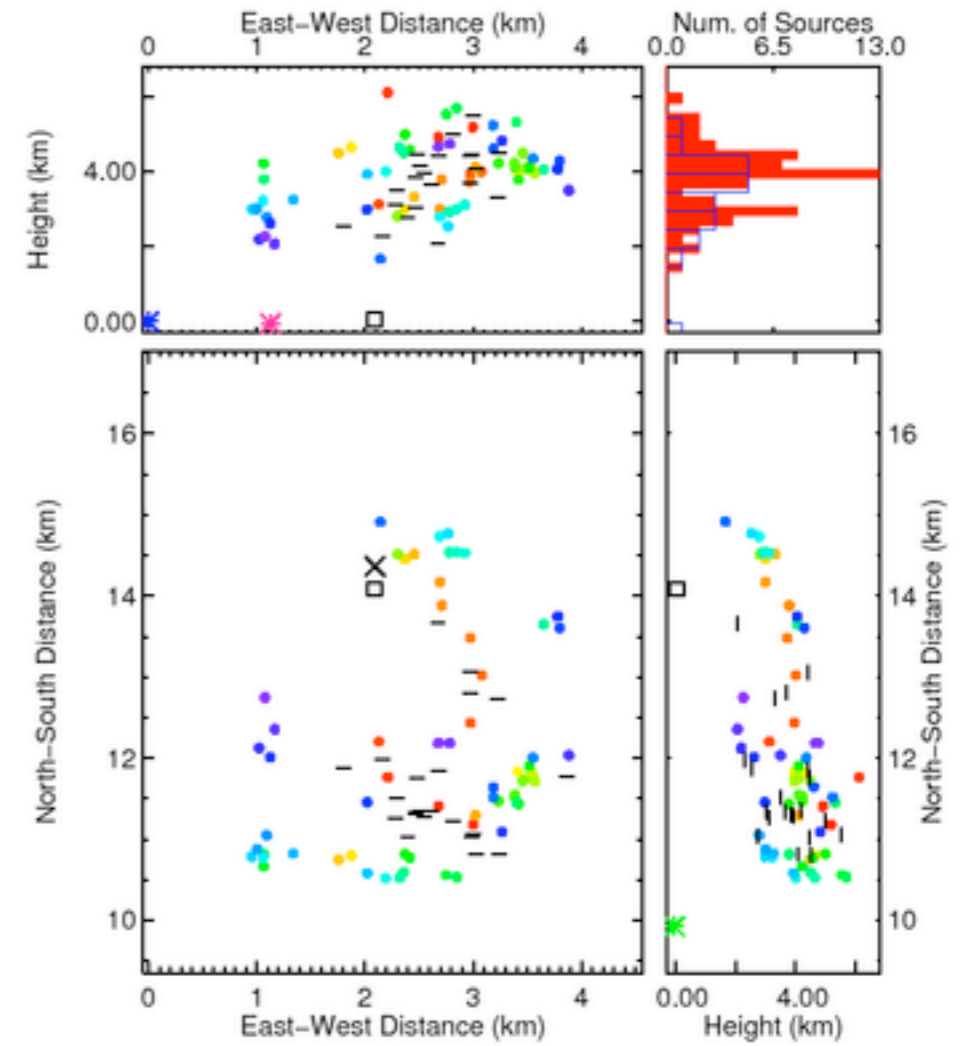
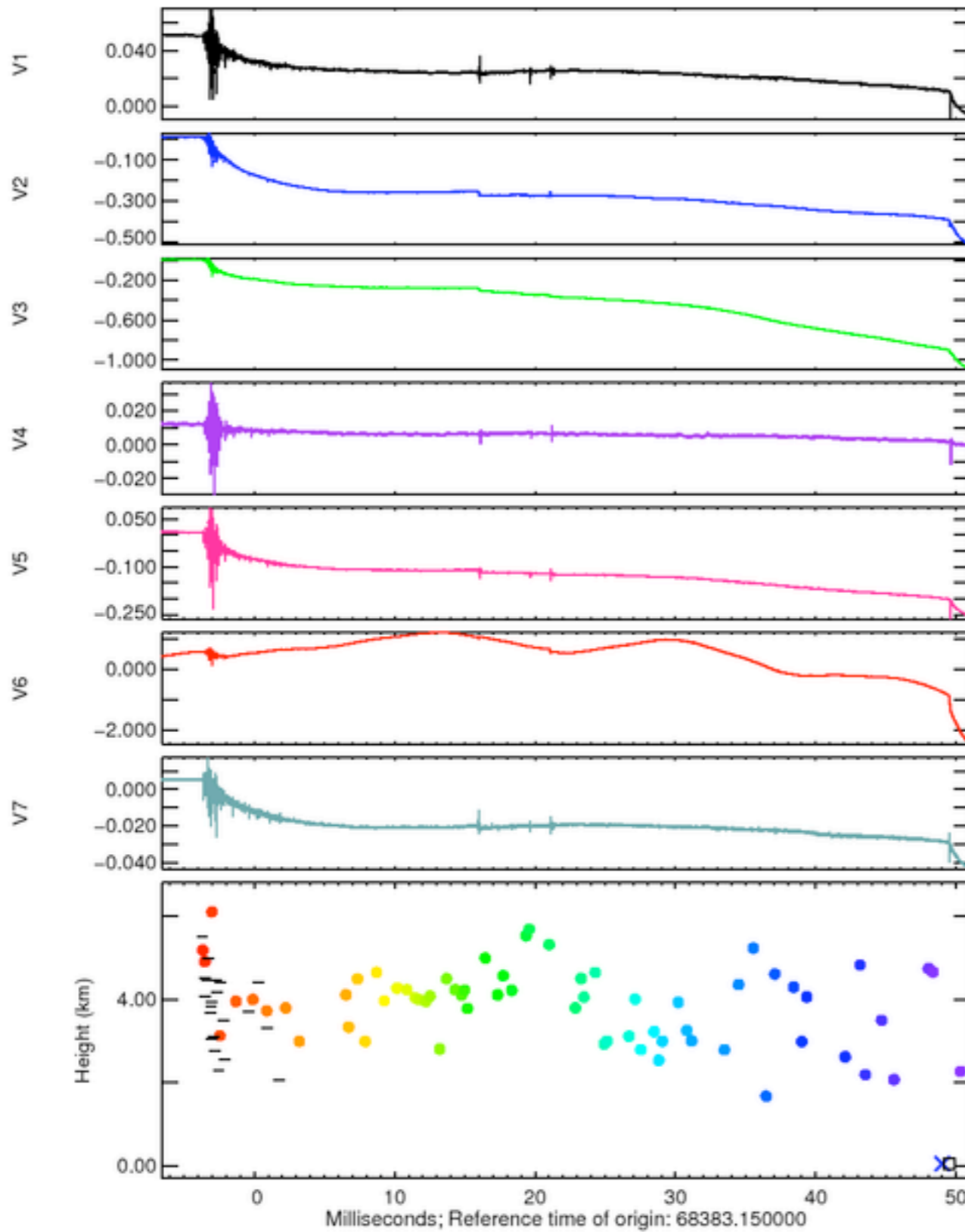
Negative CGs produce a preliminary breakdown pulse train with negative polarity pulses



Flashes with an initial intracloud component begin with positive polarity pulses



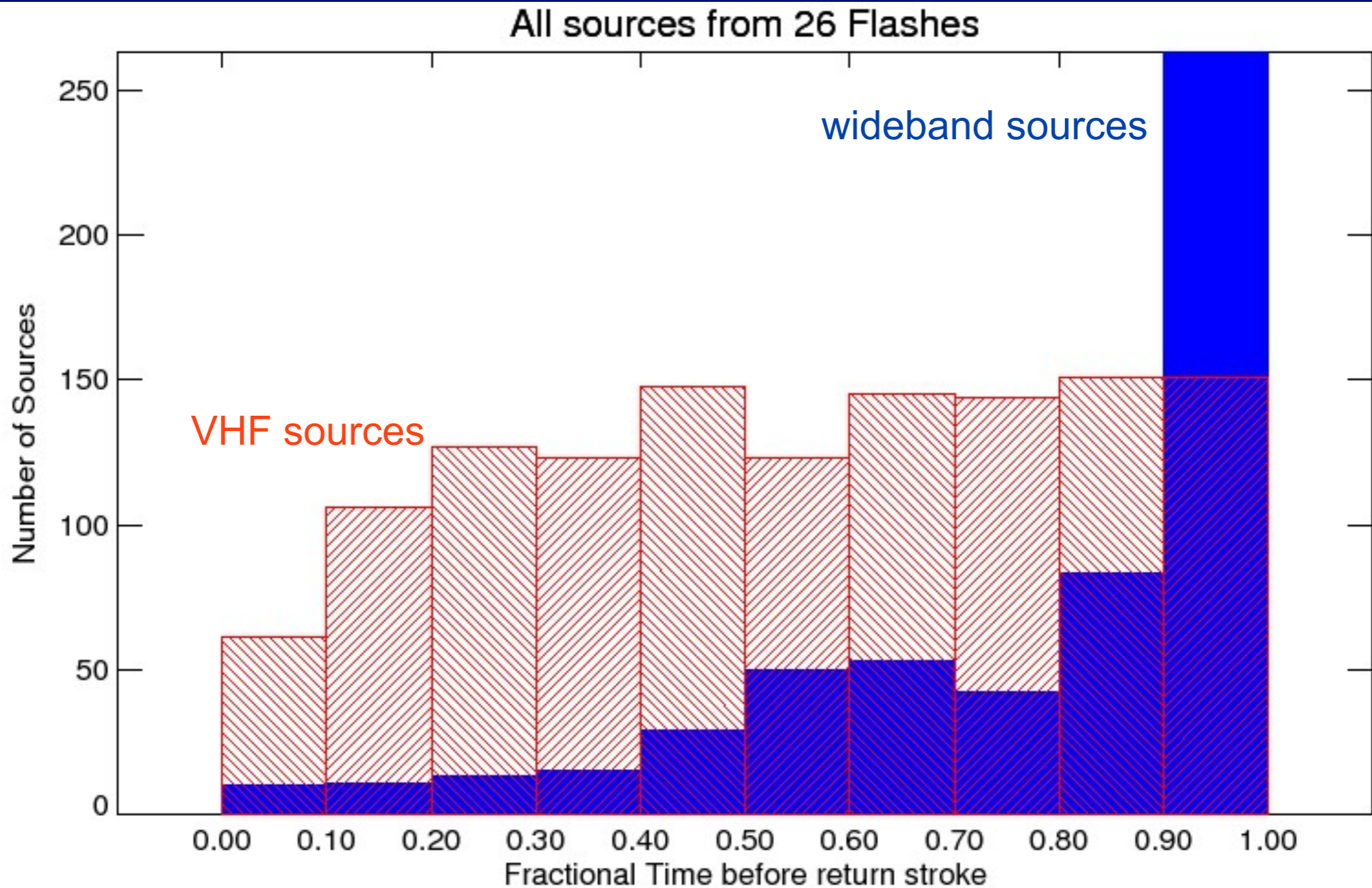
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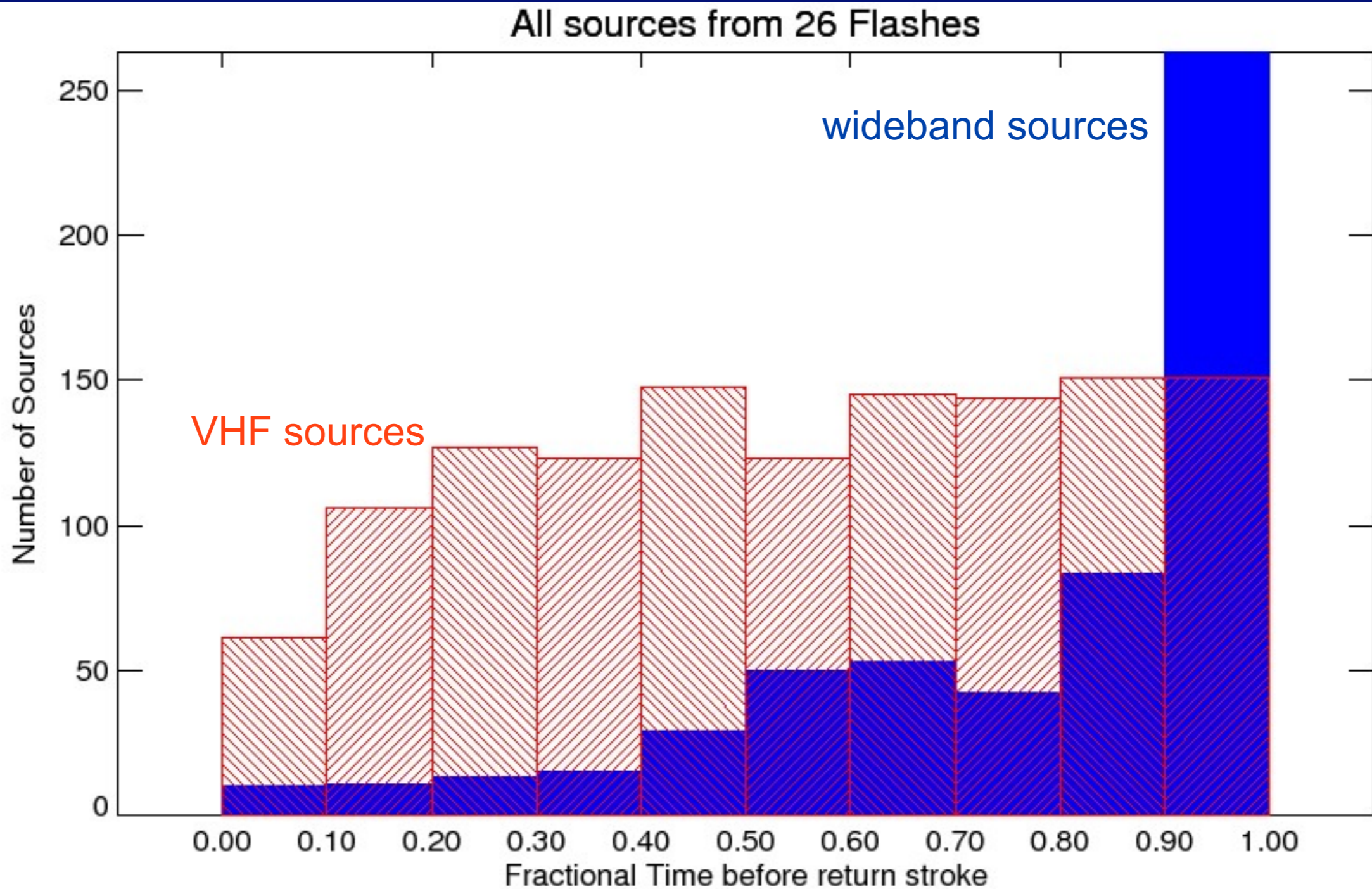
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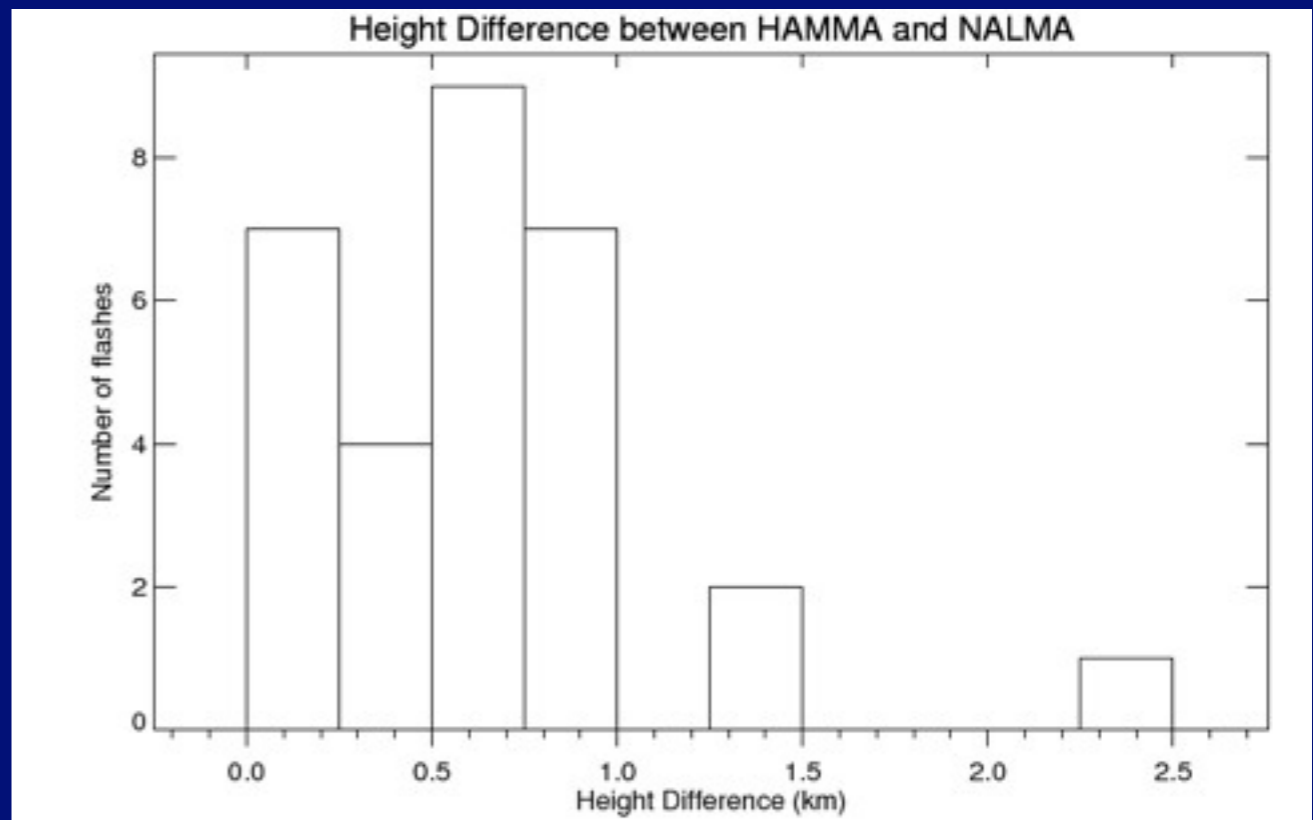
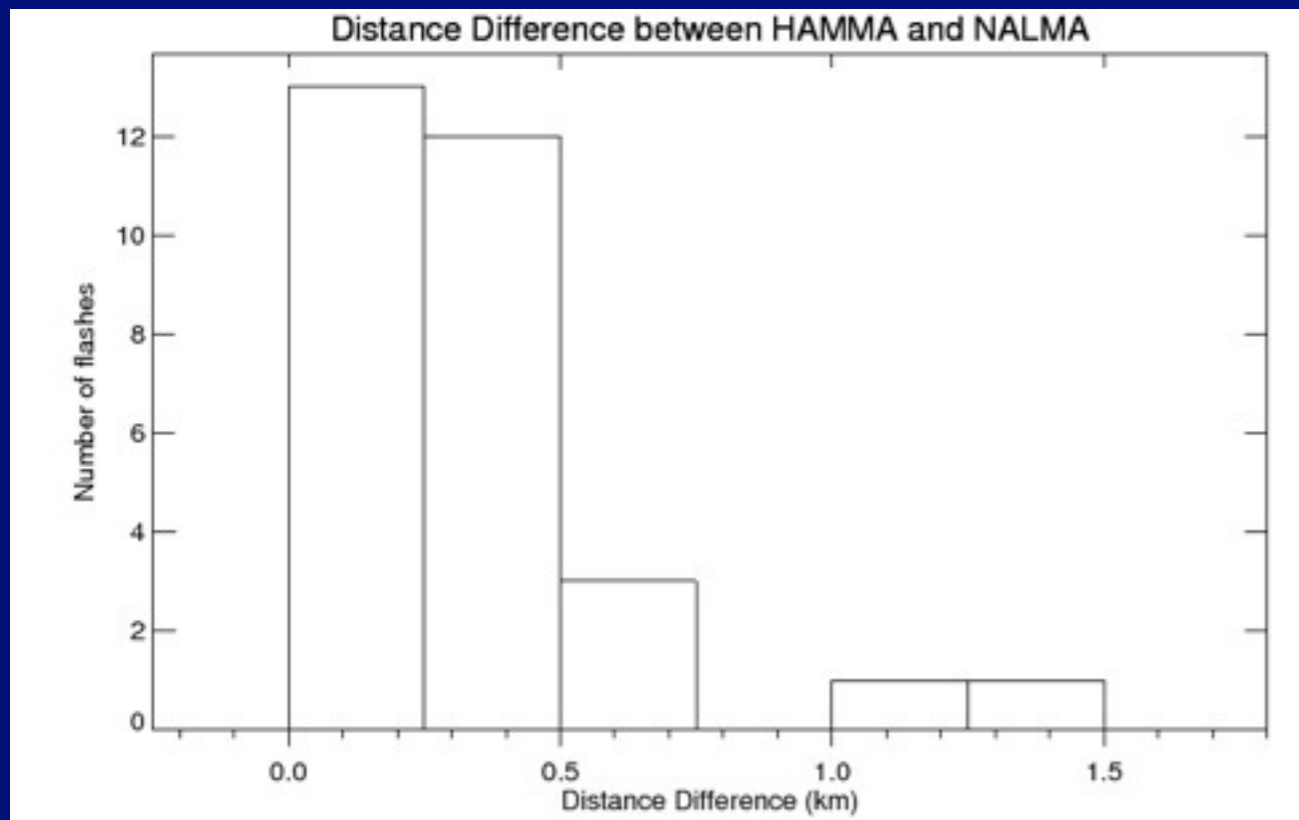
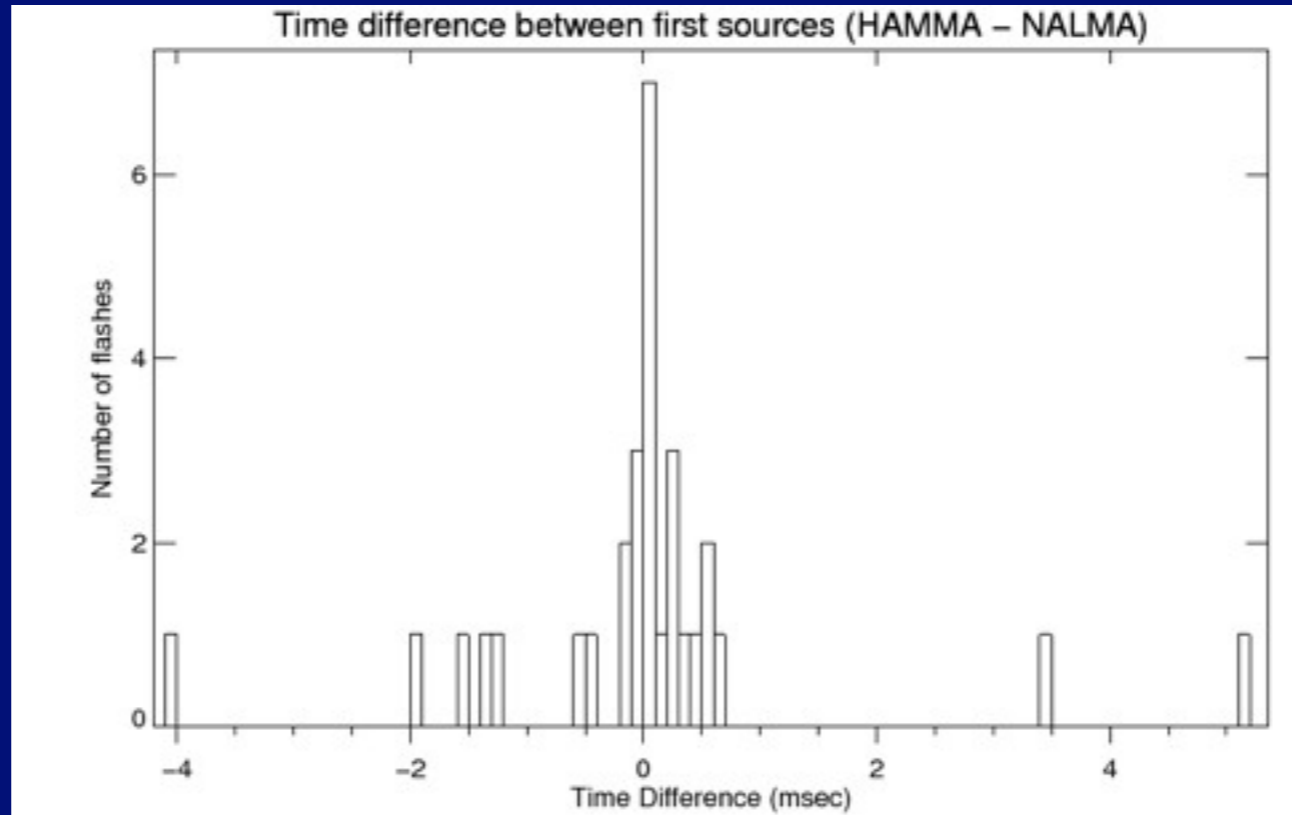


HAMMA is “very” active in the beginning of a flash

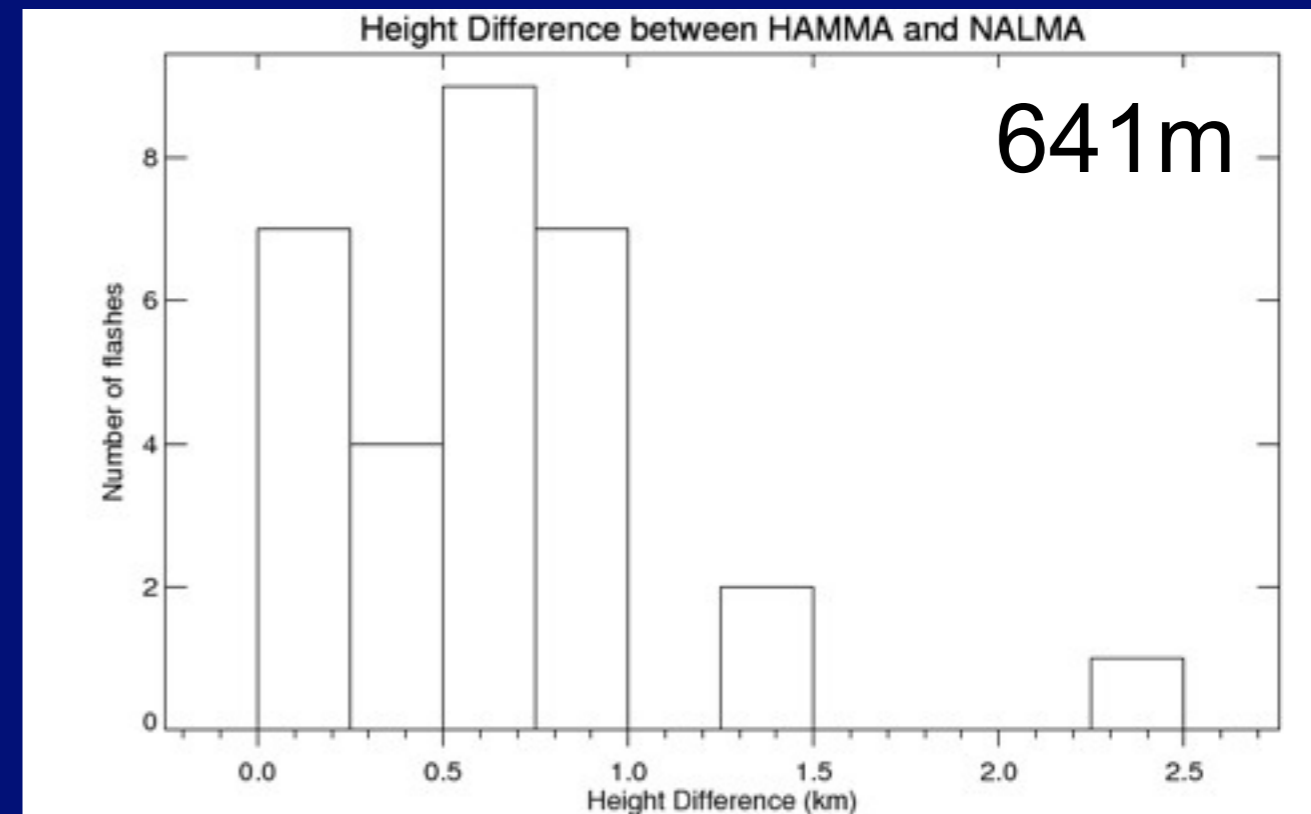
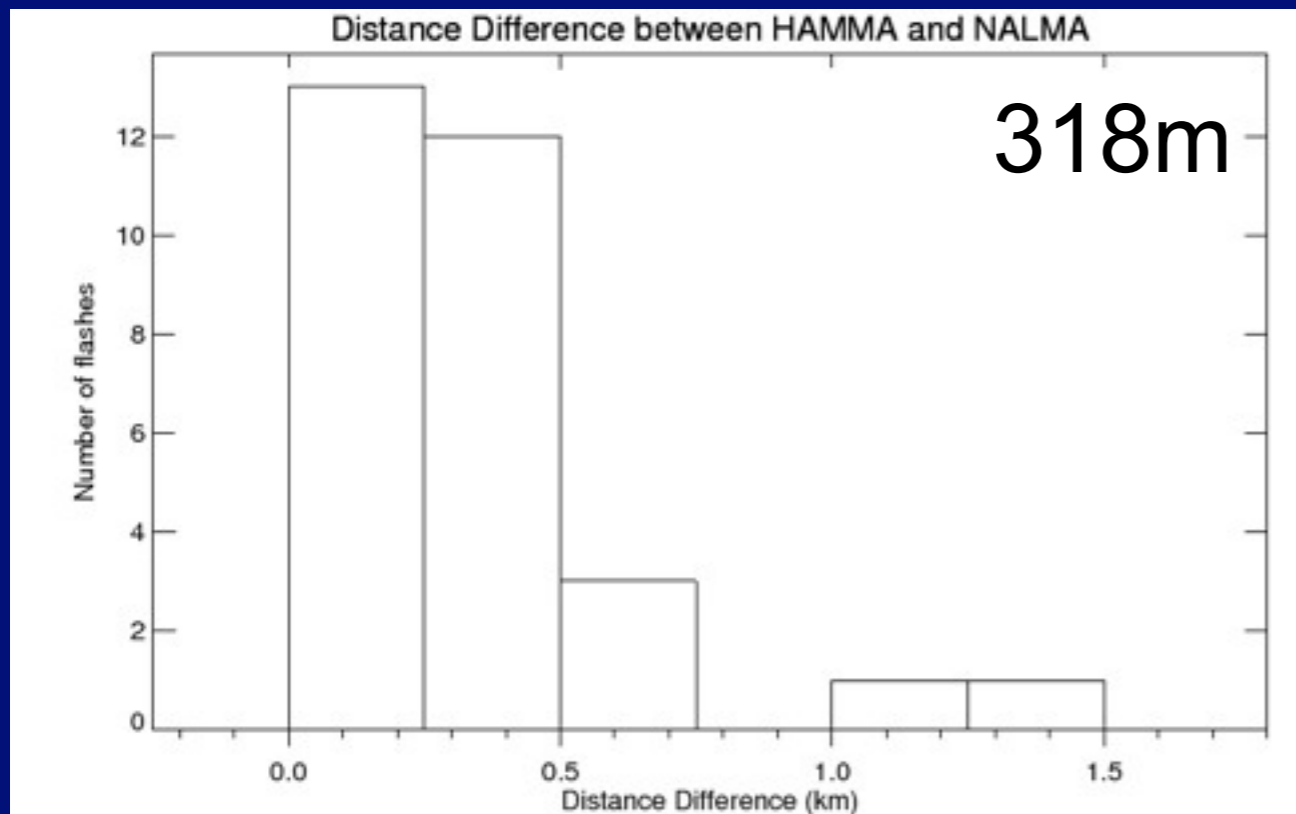
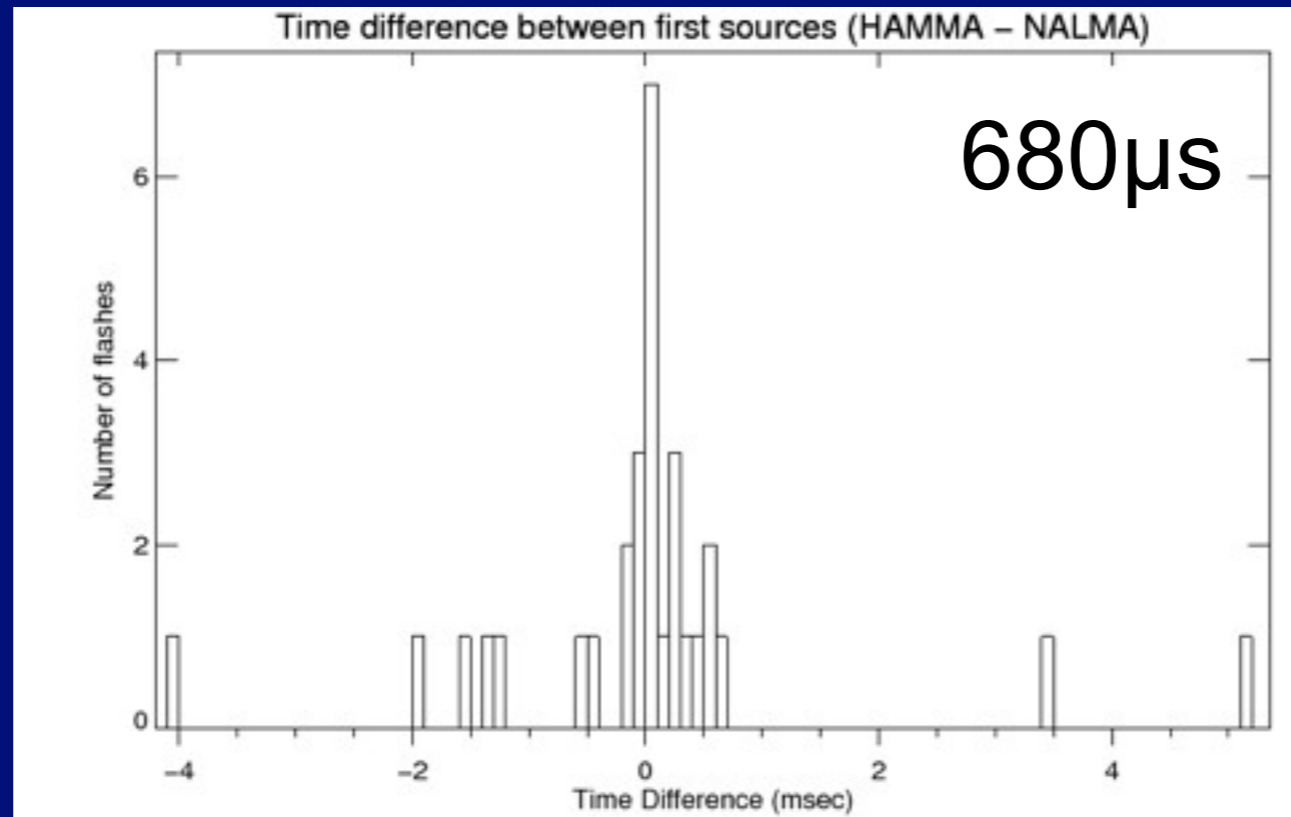


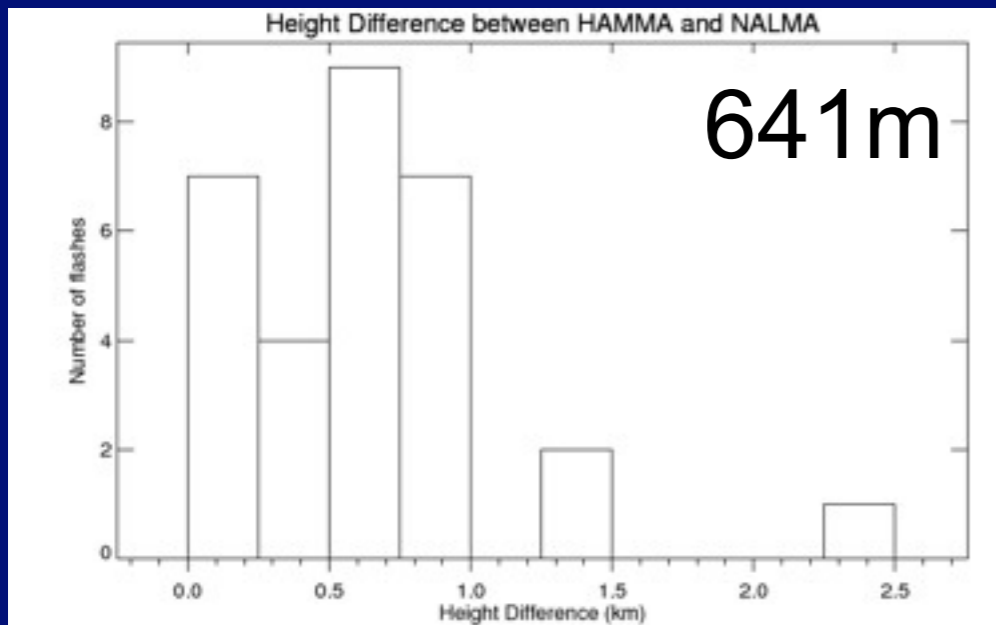
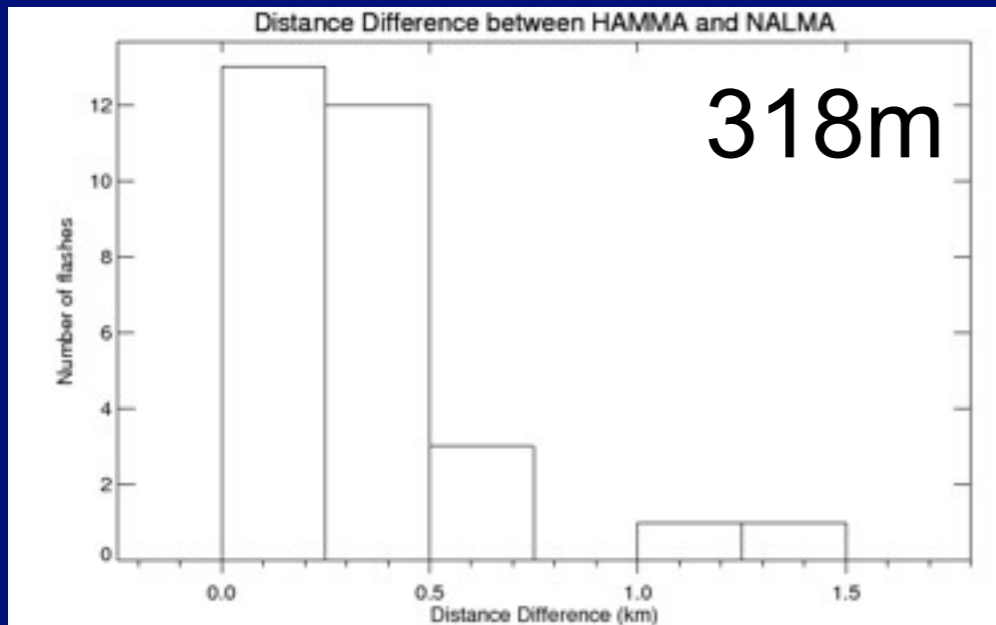
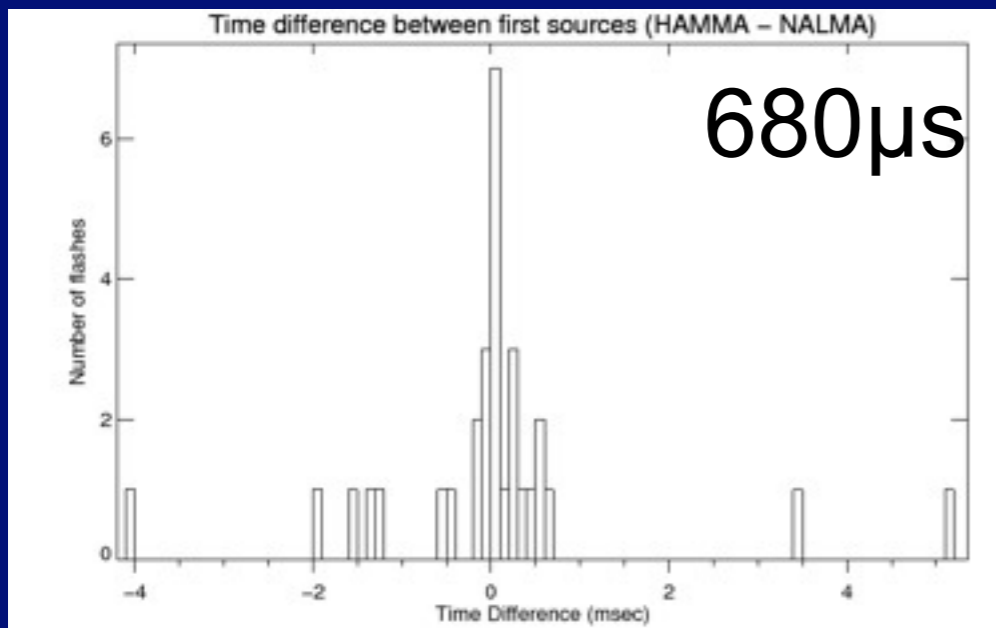
HAMMA is “very” active in the beginning of a flash
=> *wideband sources are well tuned to identify initiation*

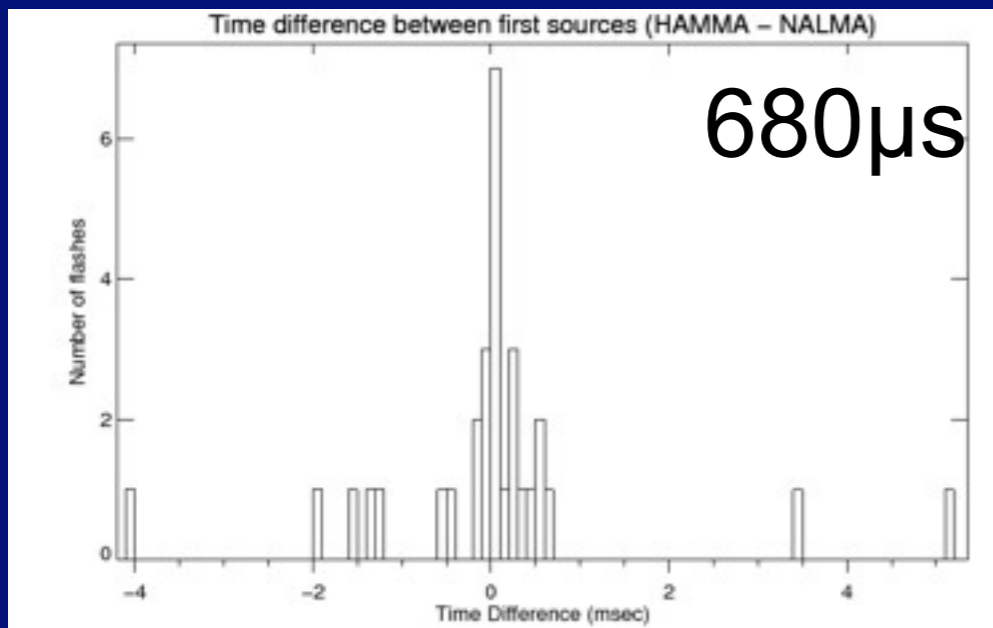
What's the difference in wideband and VHF?



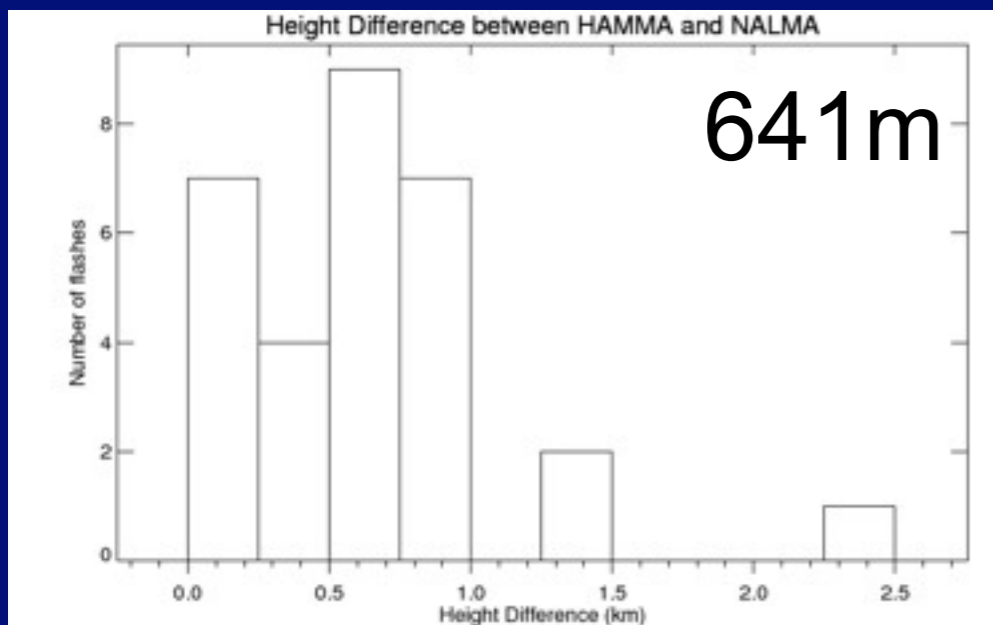
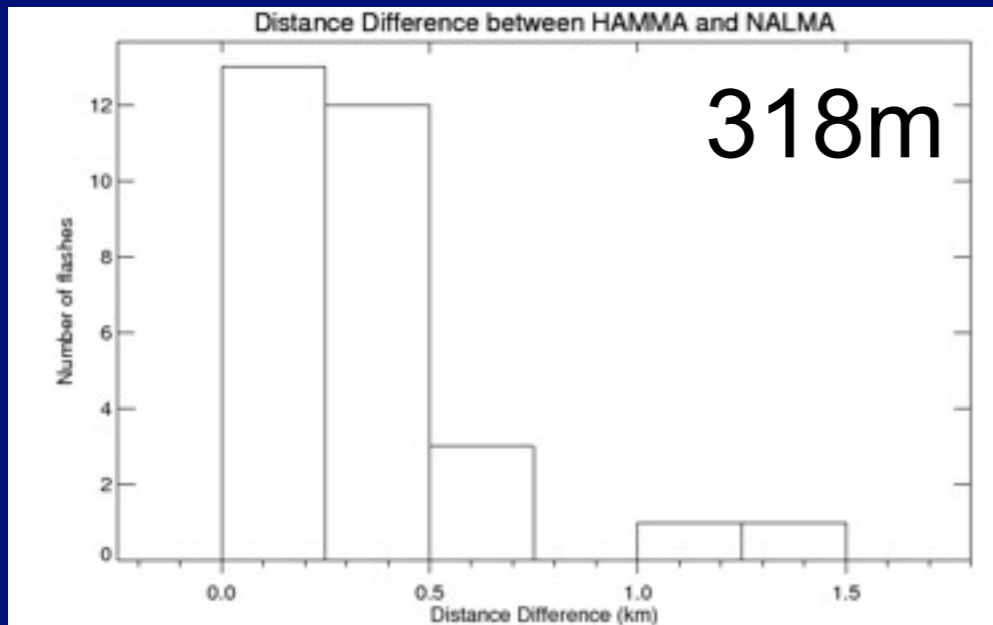
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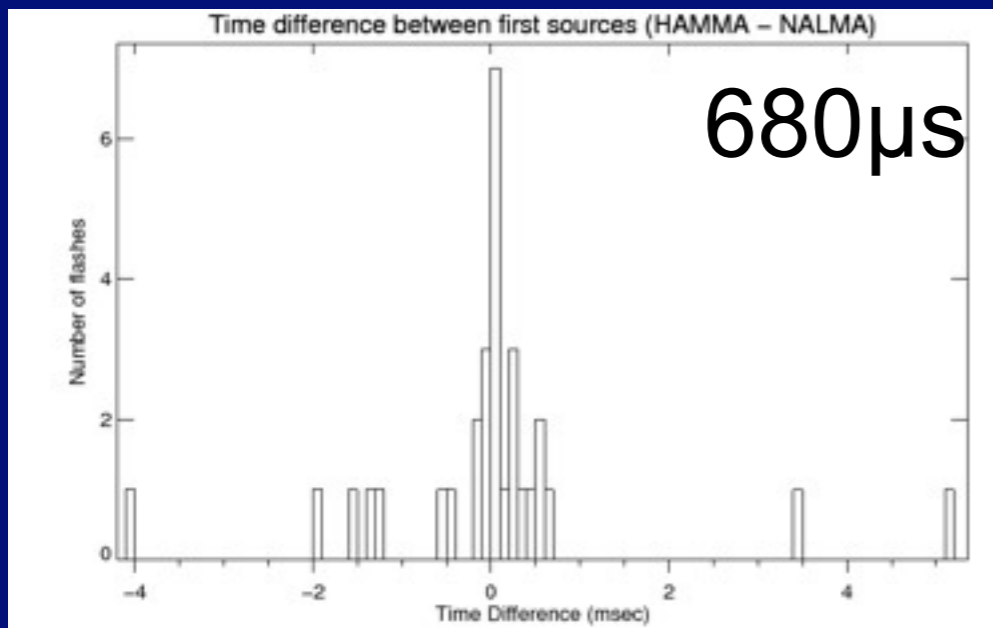




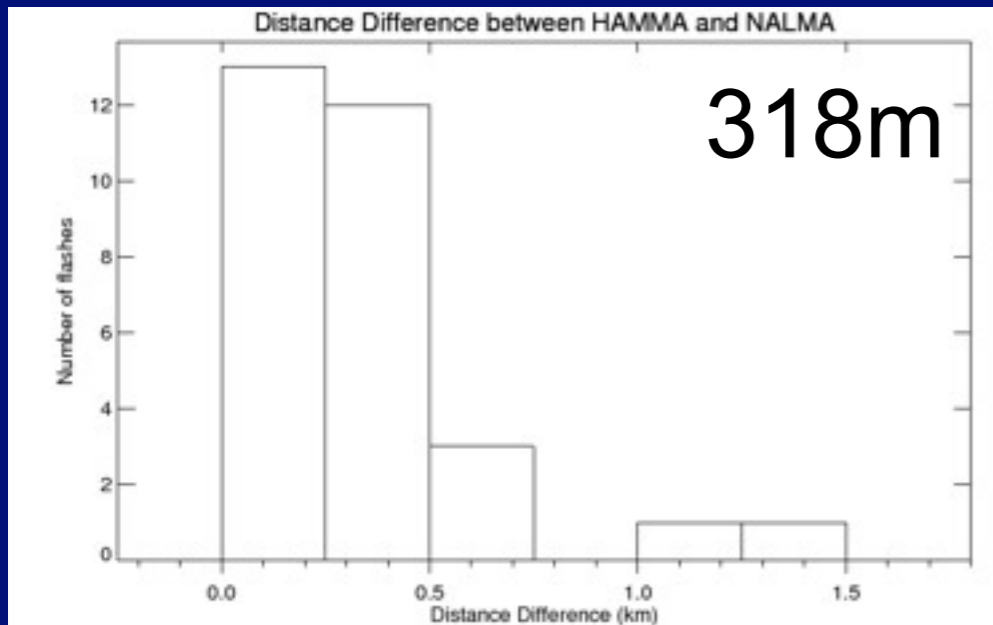


On average, initiation in the wideband and VHF are spatially and temporally similar...

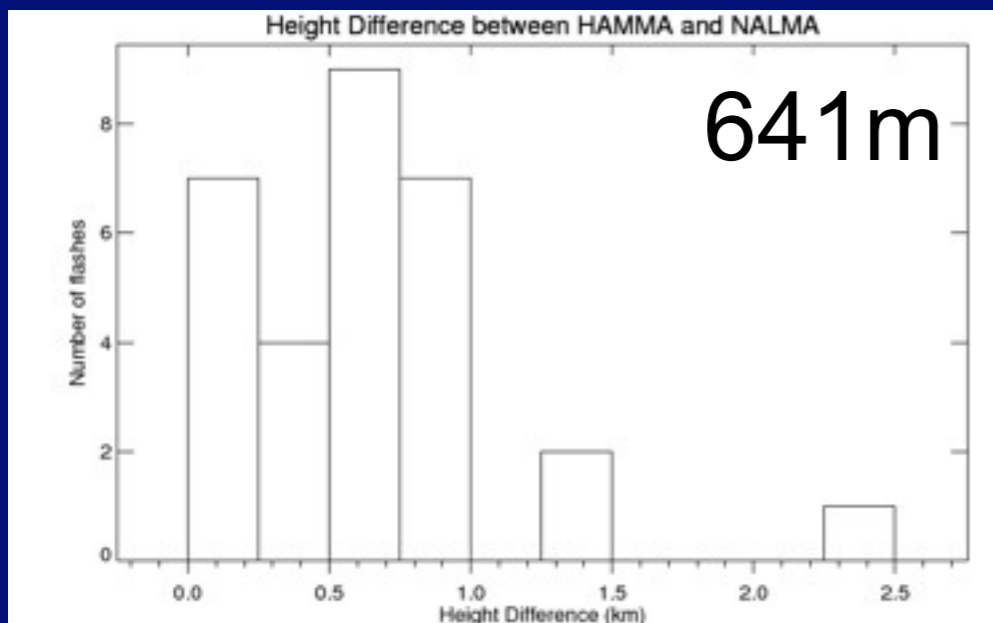




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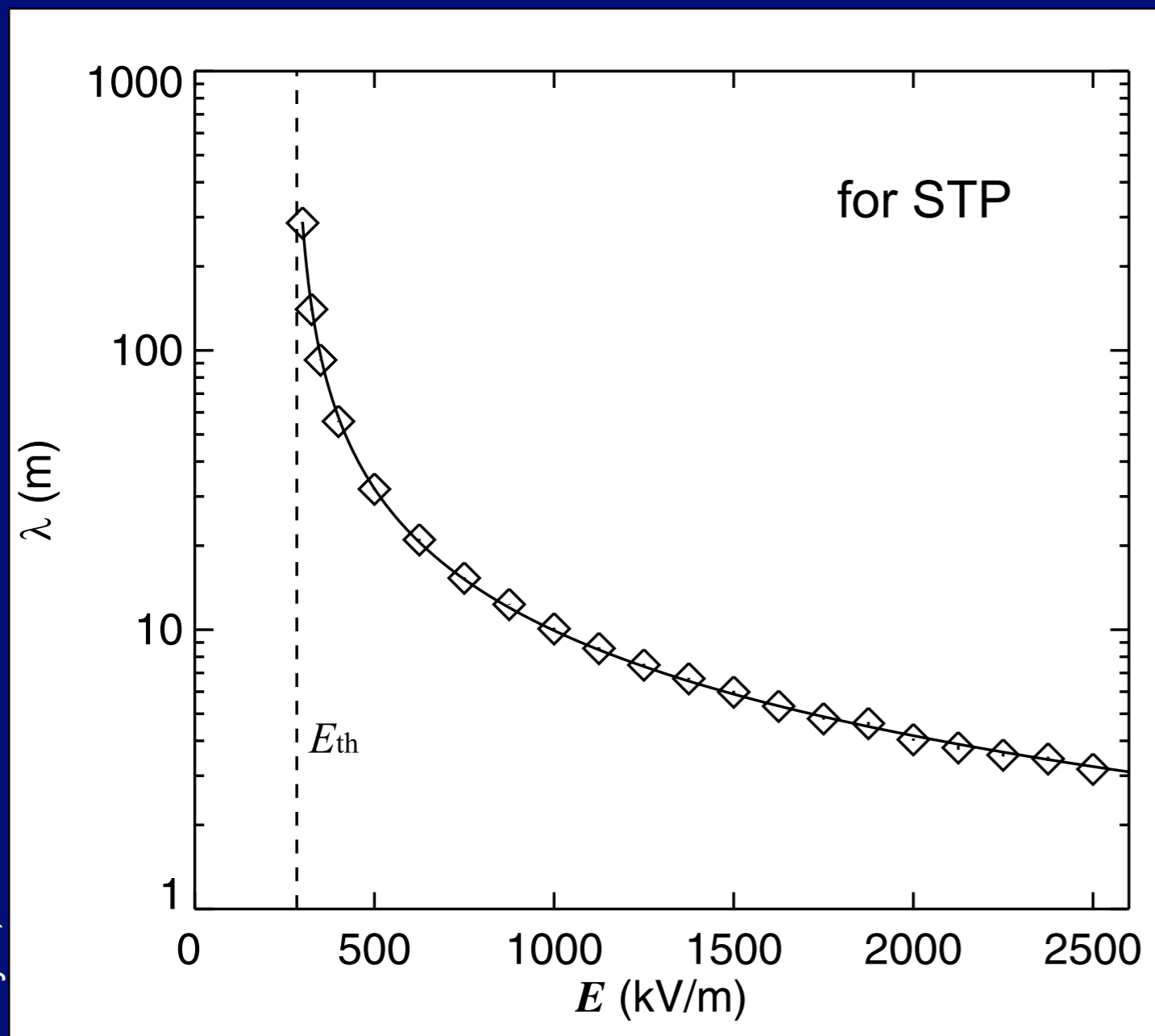


...but what does this mean for the initiation mechanism?

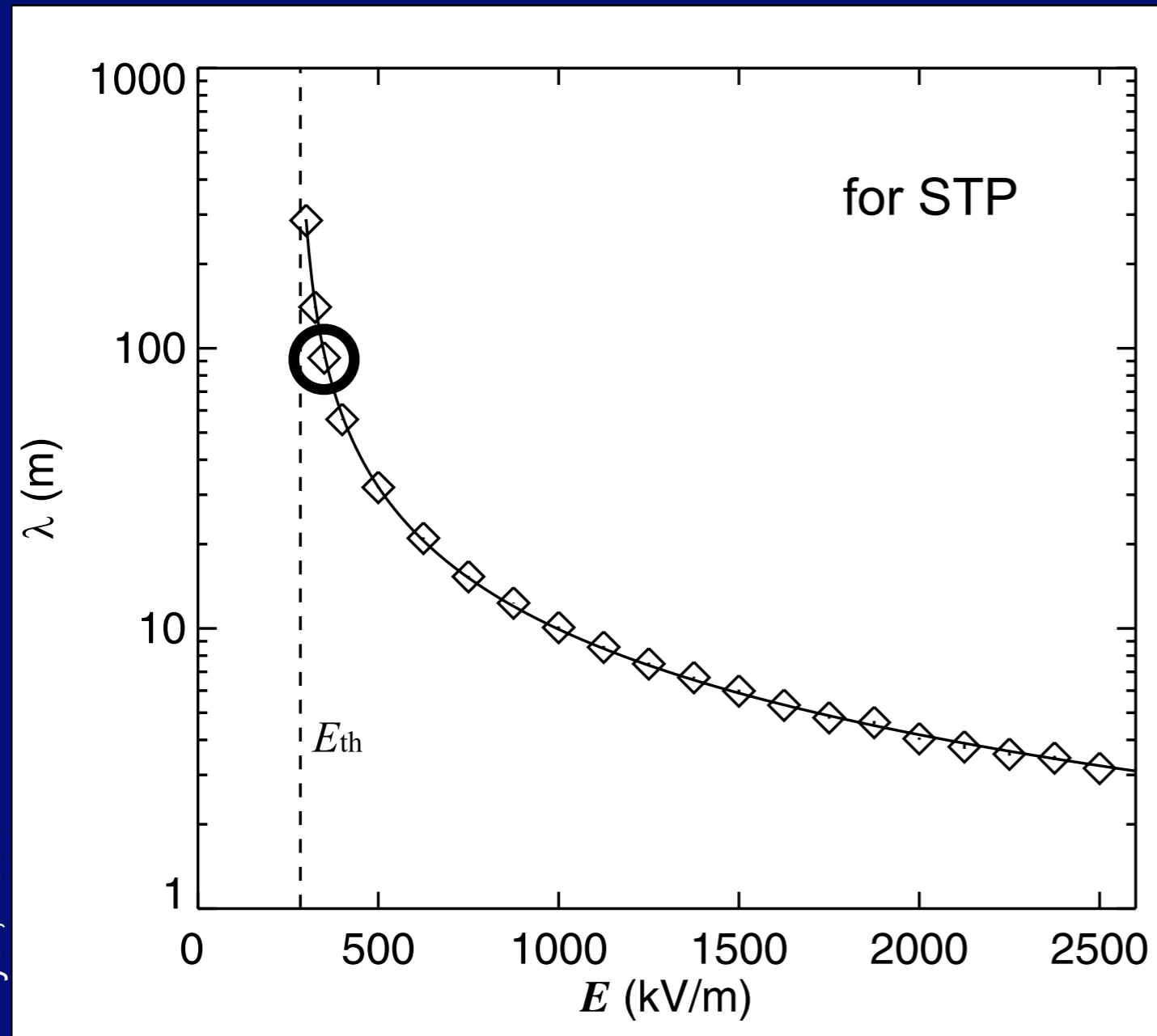


Gurevich (1999) predicts that initiation via relativistic runaway breakdown would produce wideband radiation prior to VHF radiation...

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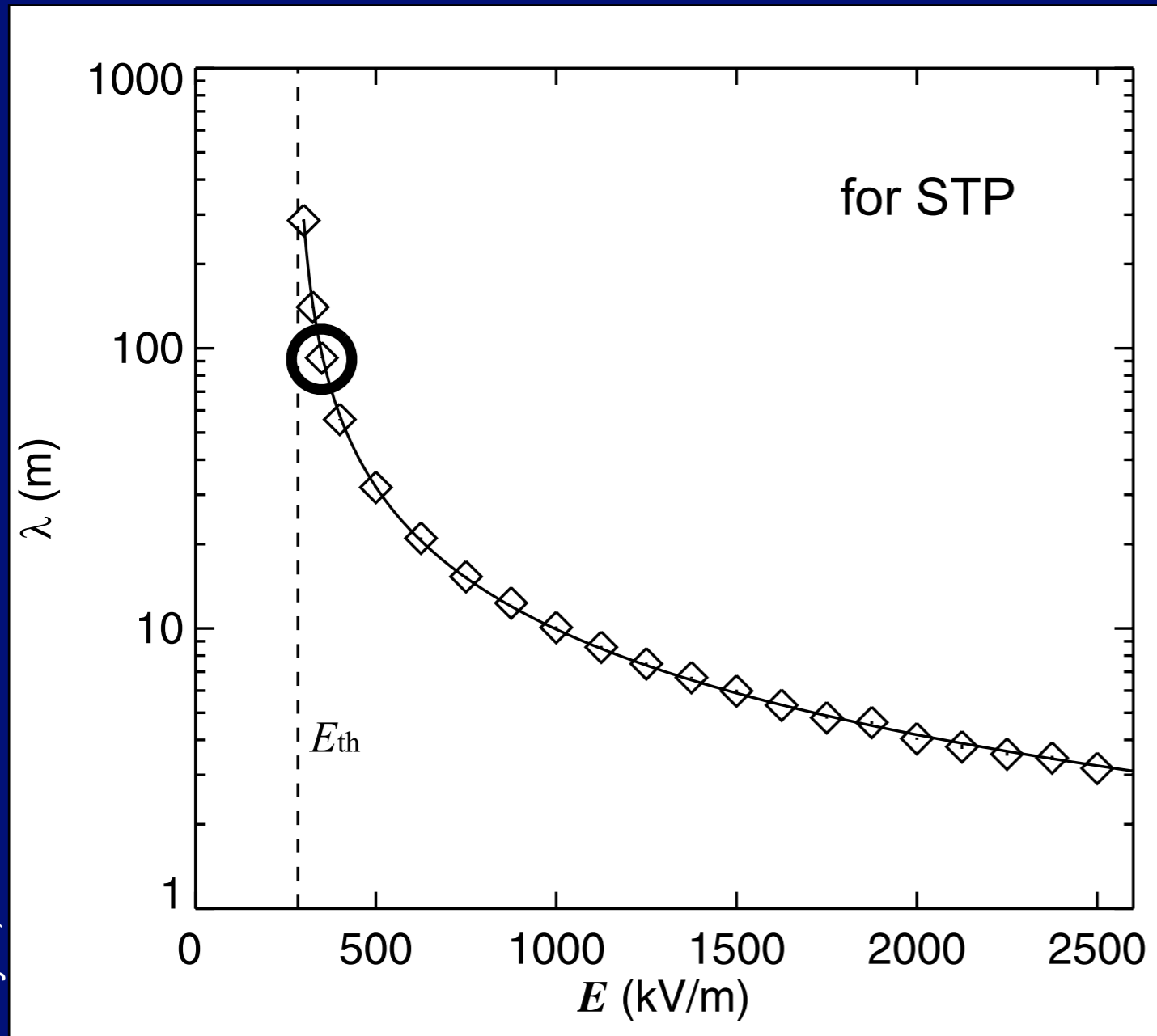


Assume:

initiation height of 6km

electric field of 175 kV/m

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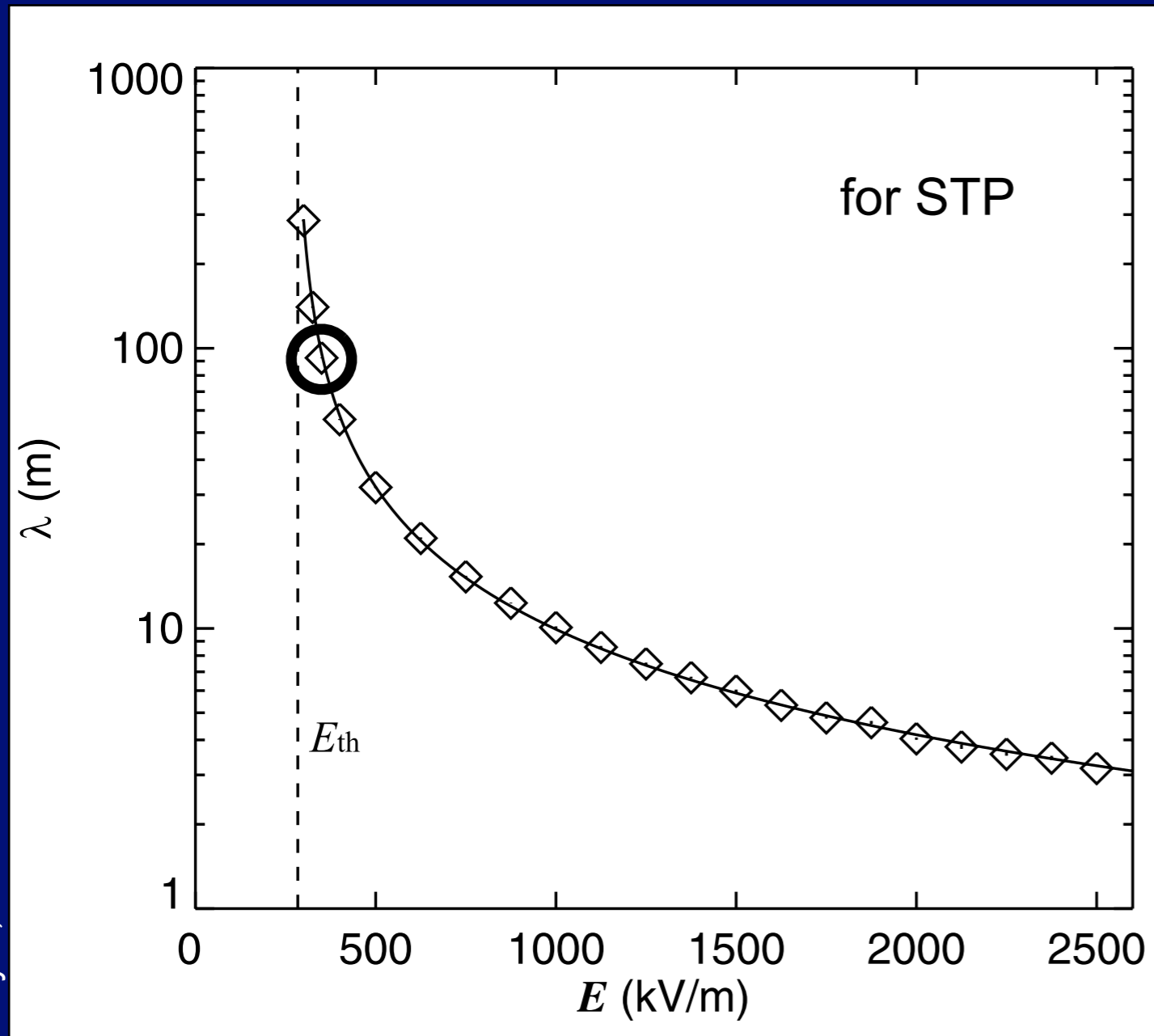
electric field of 175 kV/m

Implies:

avalanche length is ~200m

total length of the high field region is ~2km

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Assume:

initiation height of 6km

electric field of 175 kV/m

Implies:

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total length of the high field region is ~2km

Further, Coleman and Dwyer (2006) show that the avalanche progresses with a speed of $0.89c$

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Measurement	$685\mu\text{s}$	Bitzer, 2011

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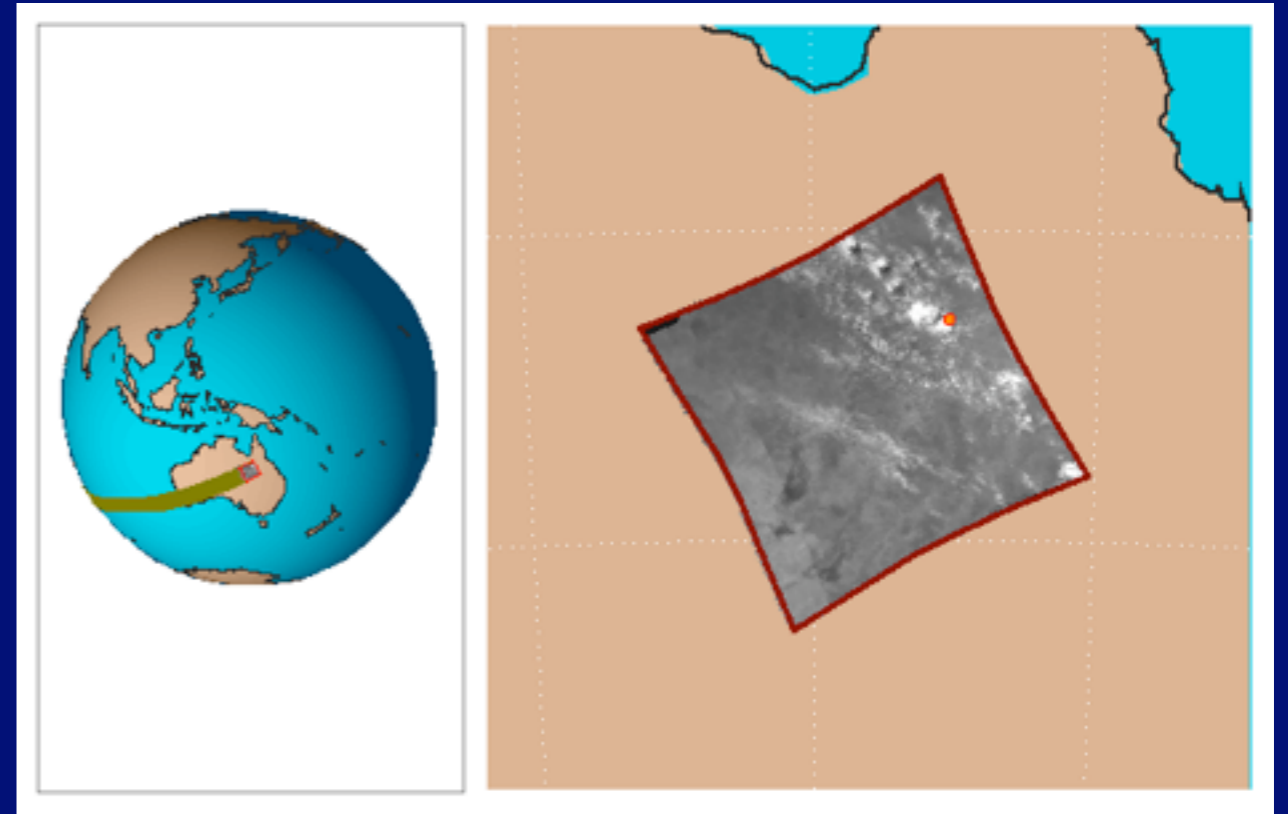
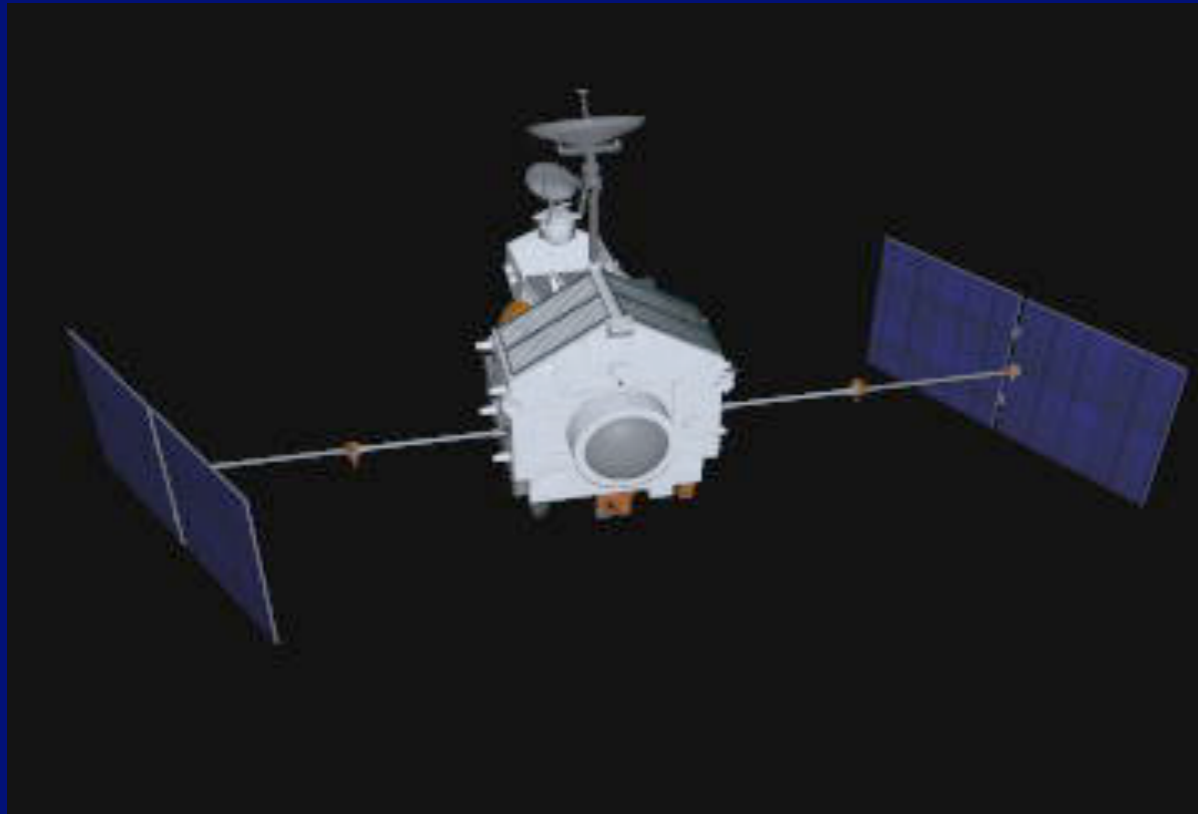
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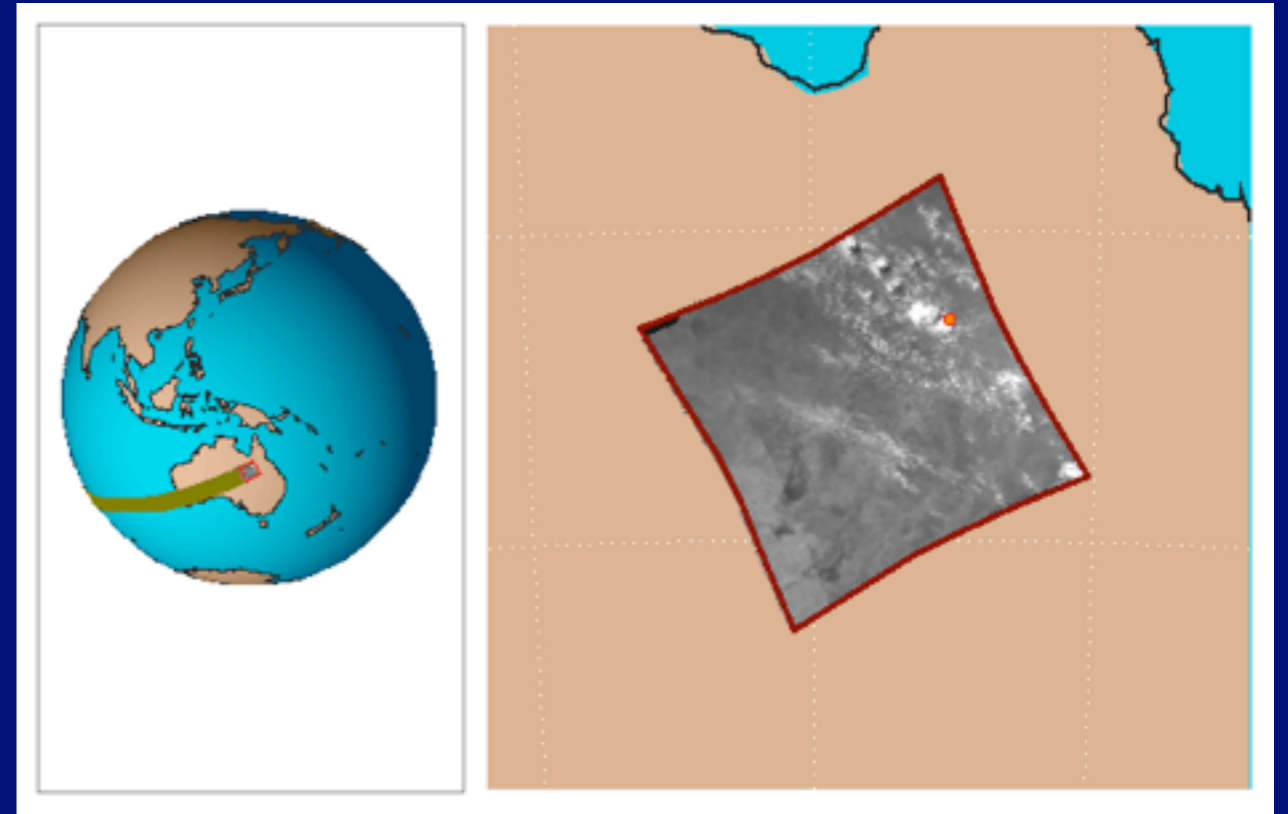
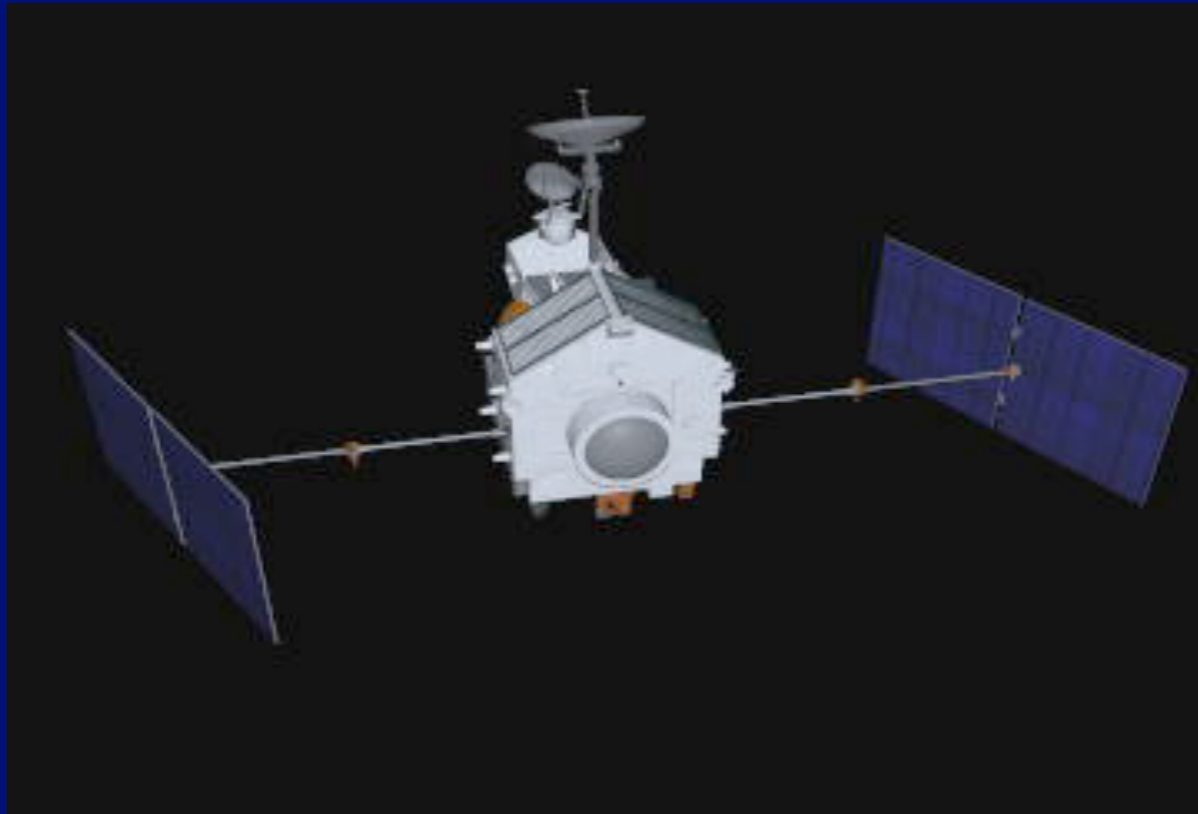
Further....

simulations (Dwyer, 2010) show the region of slow electrons is far too diffuse to yield the required conductivity.

LIS (and by extension, GLM)
does not
detect *flashes...*

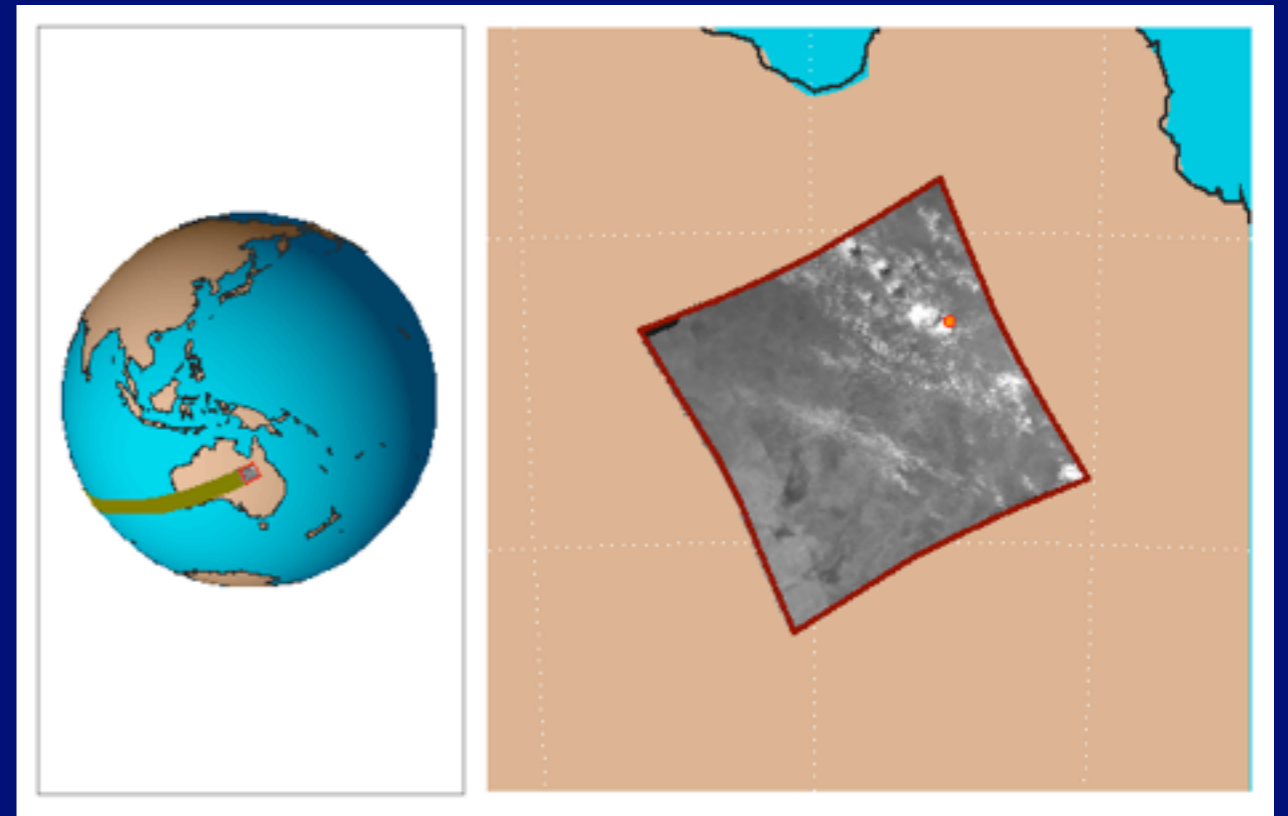


LIS is an **optical event detector**



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*These events can be classified into **strokes** groups and flashes.*

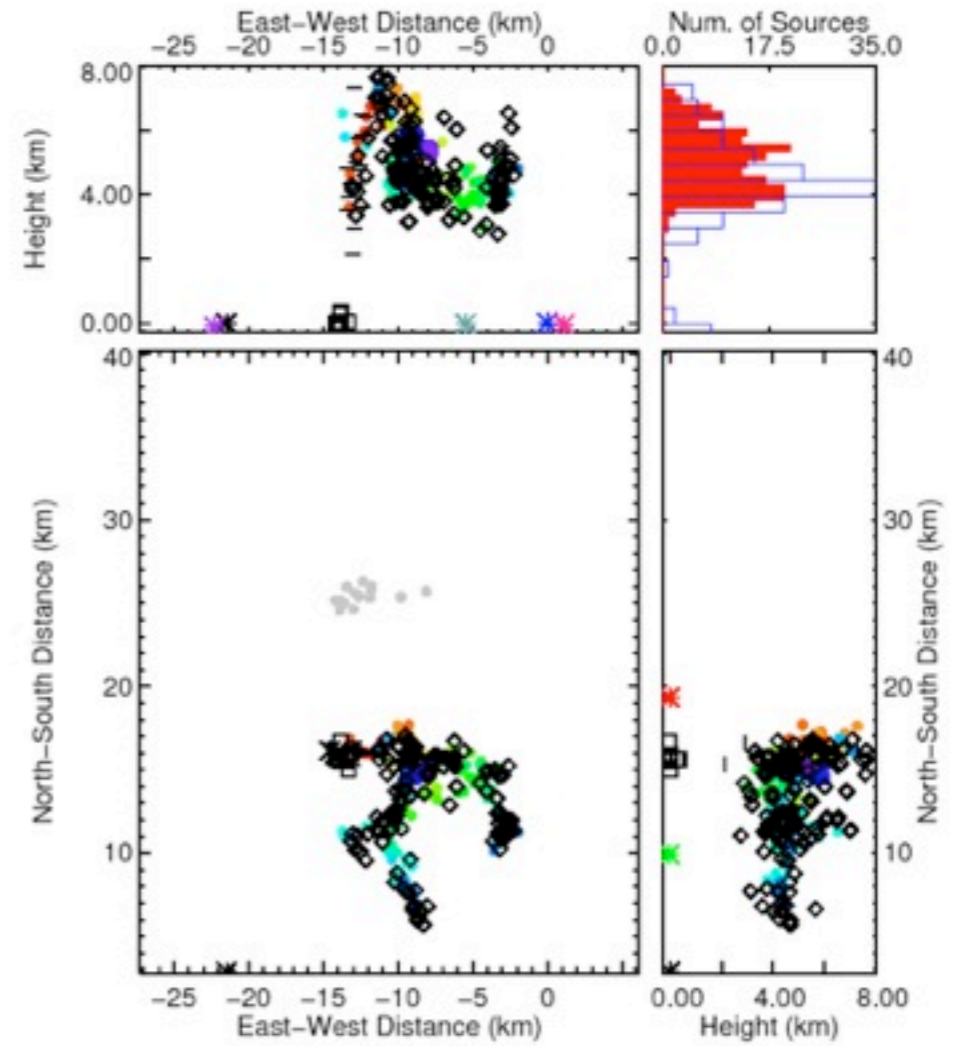
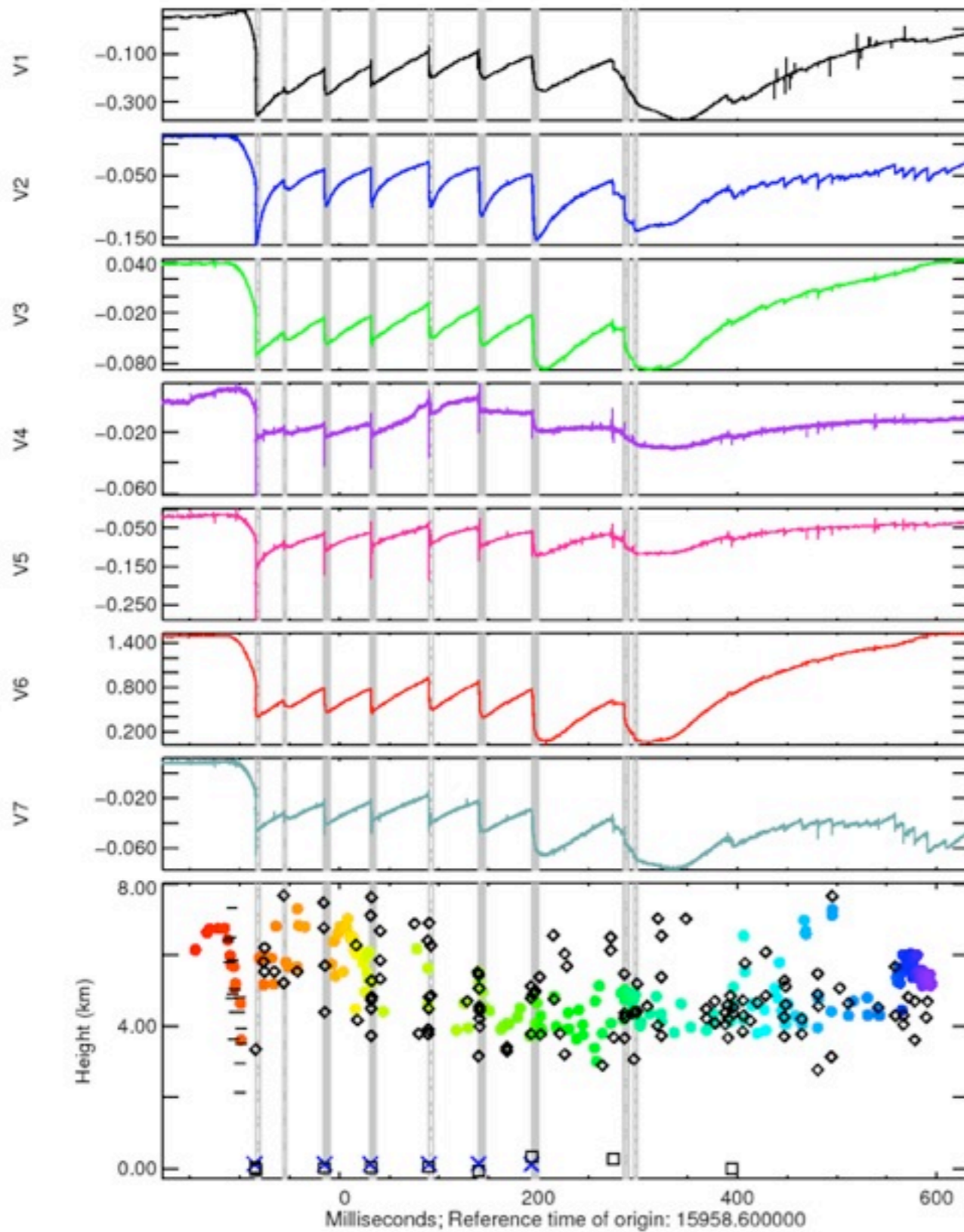


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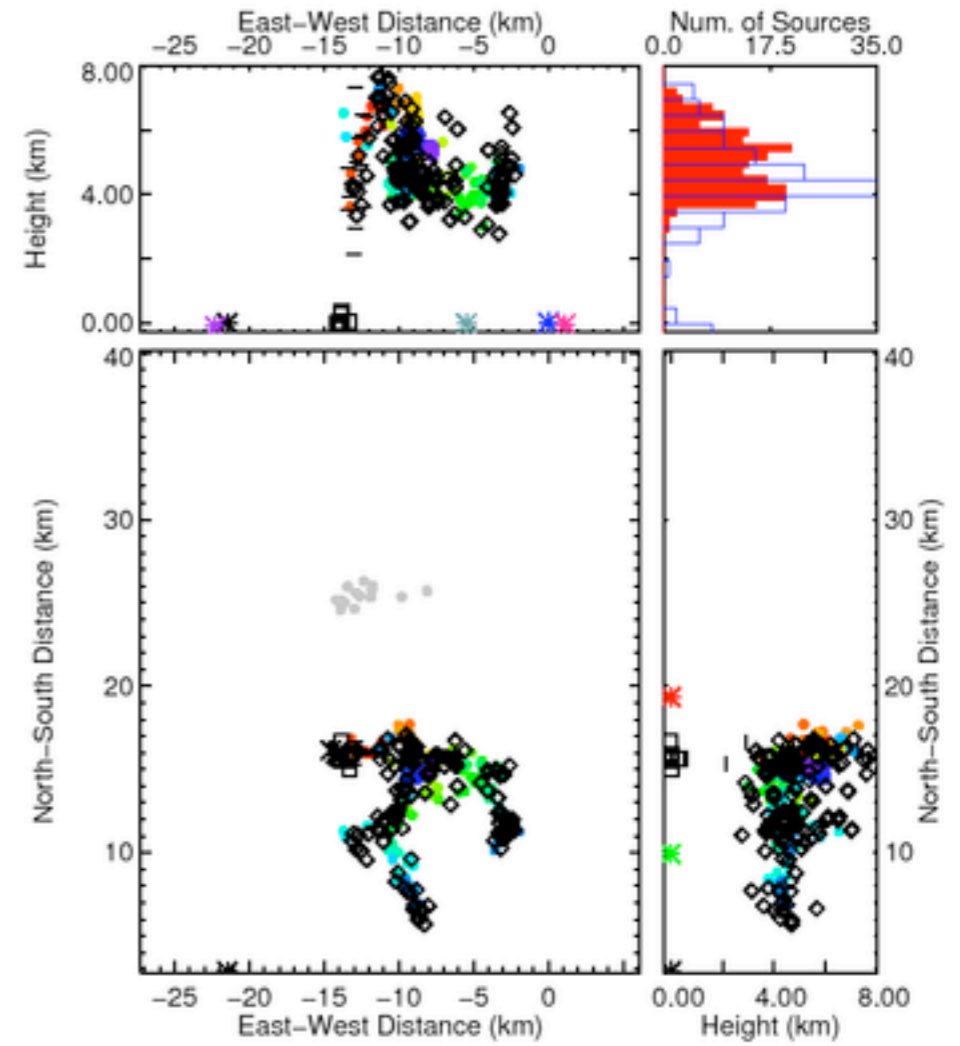
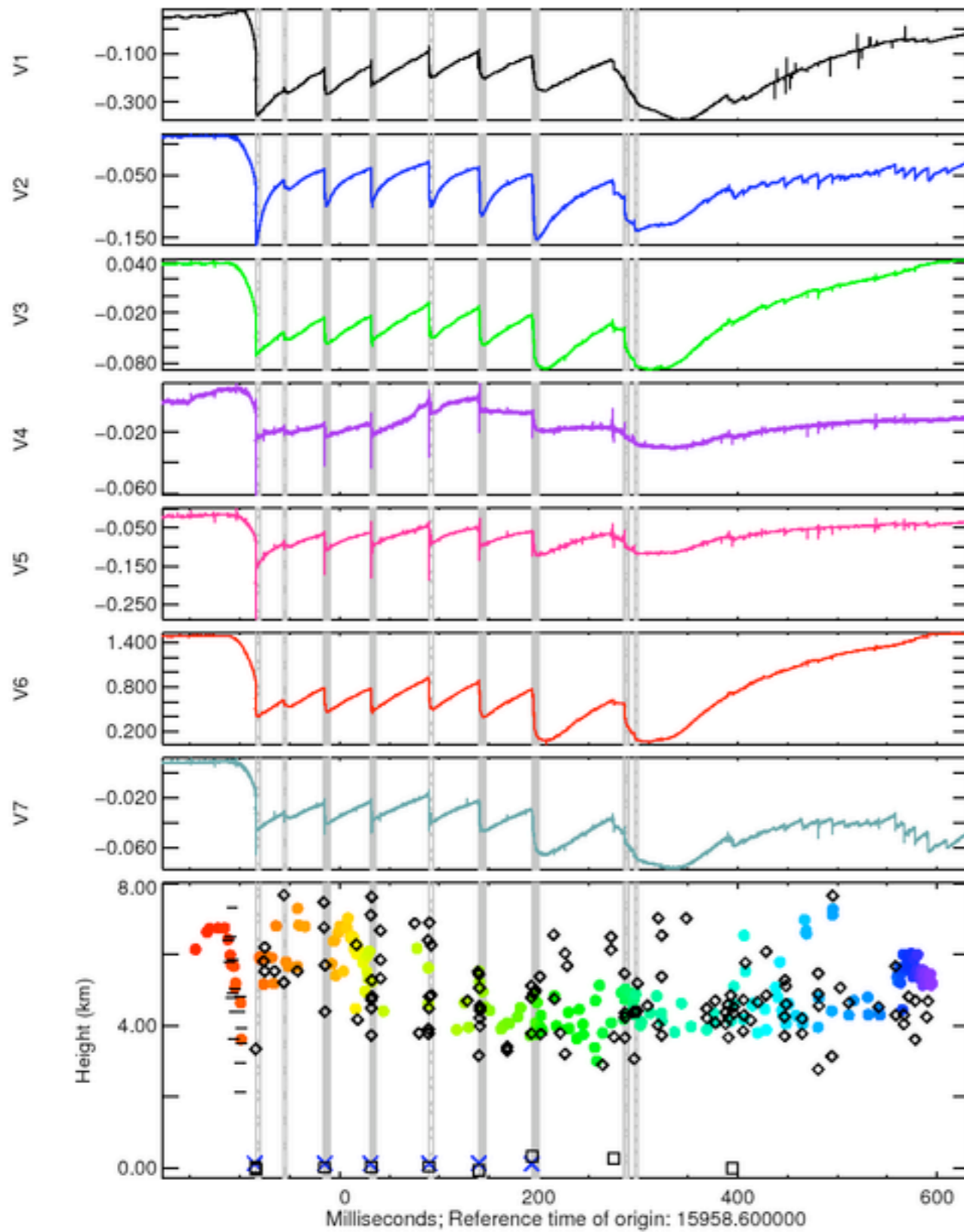
Validation of LIS/GLM measurements should use instruments sensitive to the part of the lightning discharge which produces optical emission

2010/10/25 04:25:58



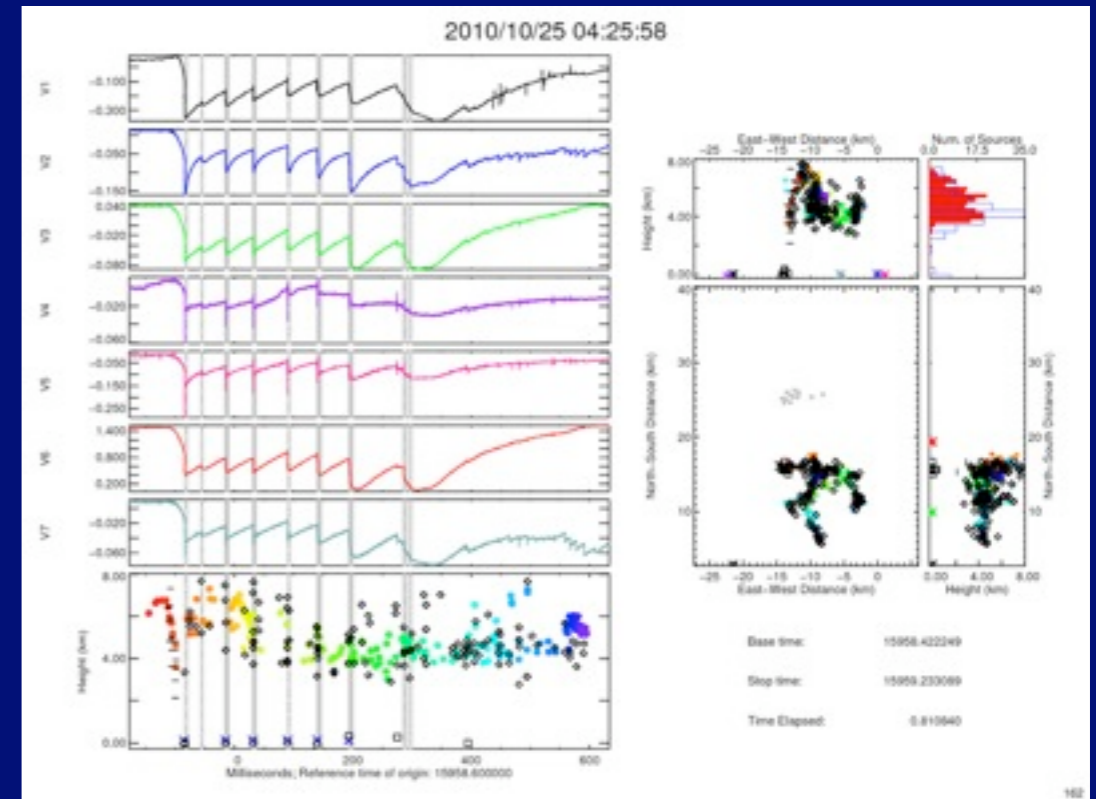
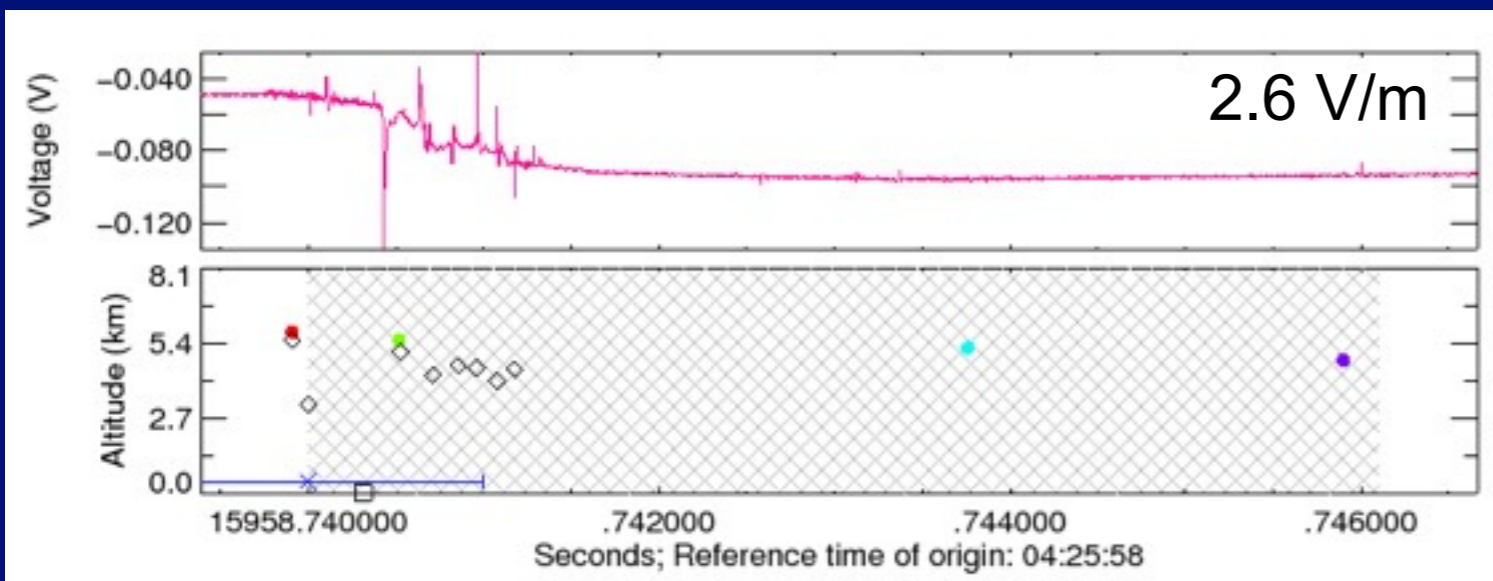
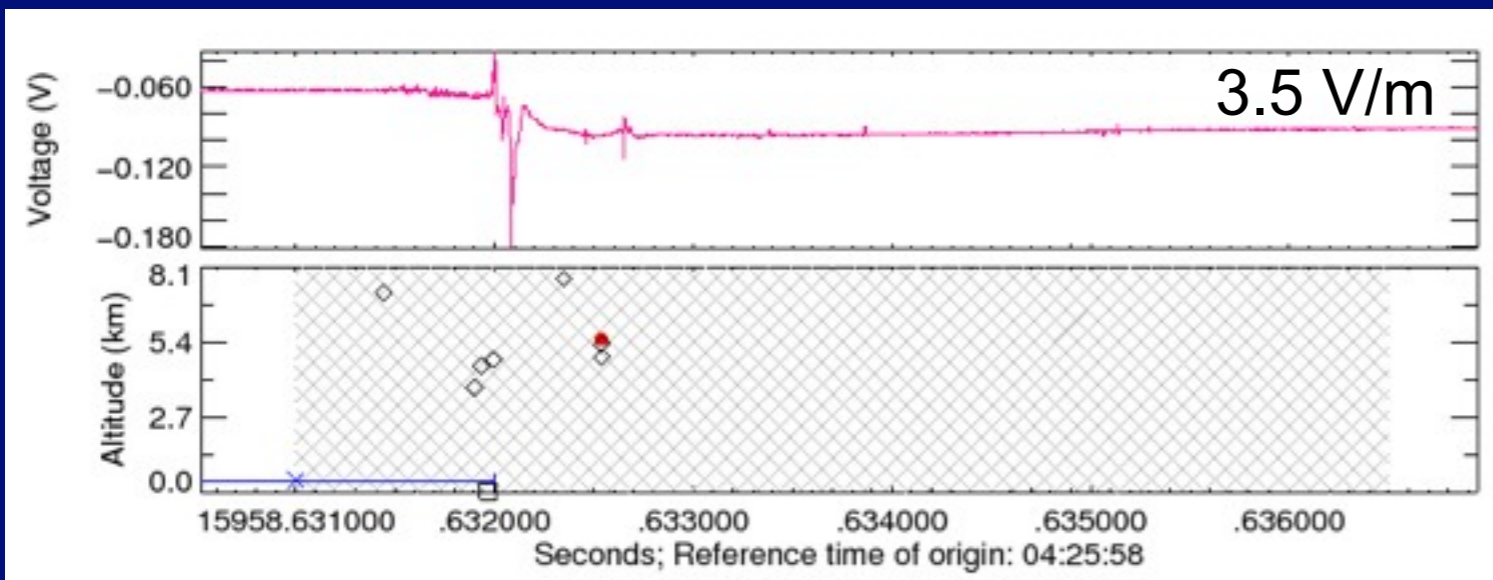
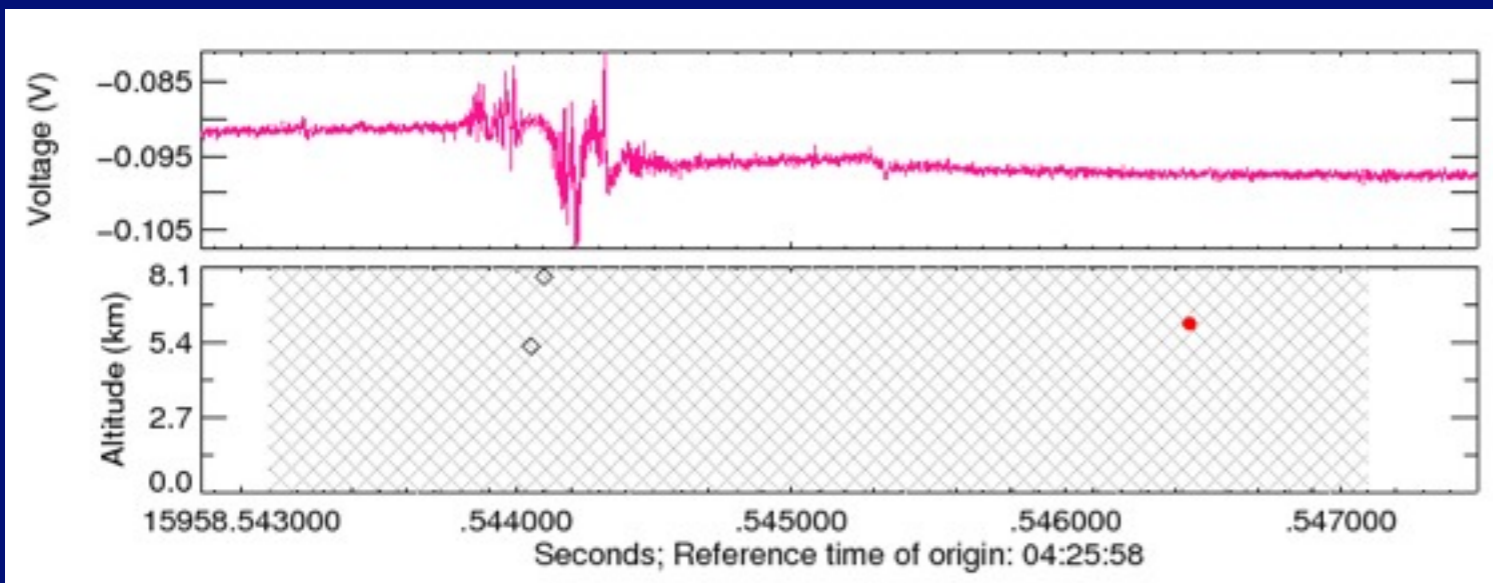
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Time Elapsed: 0.810840

2010/10/25 04:25:58

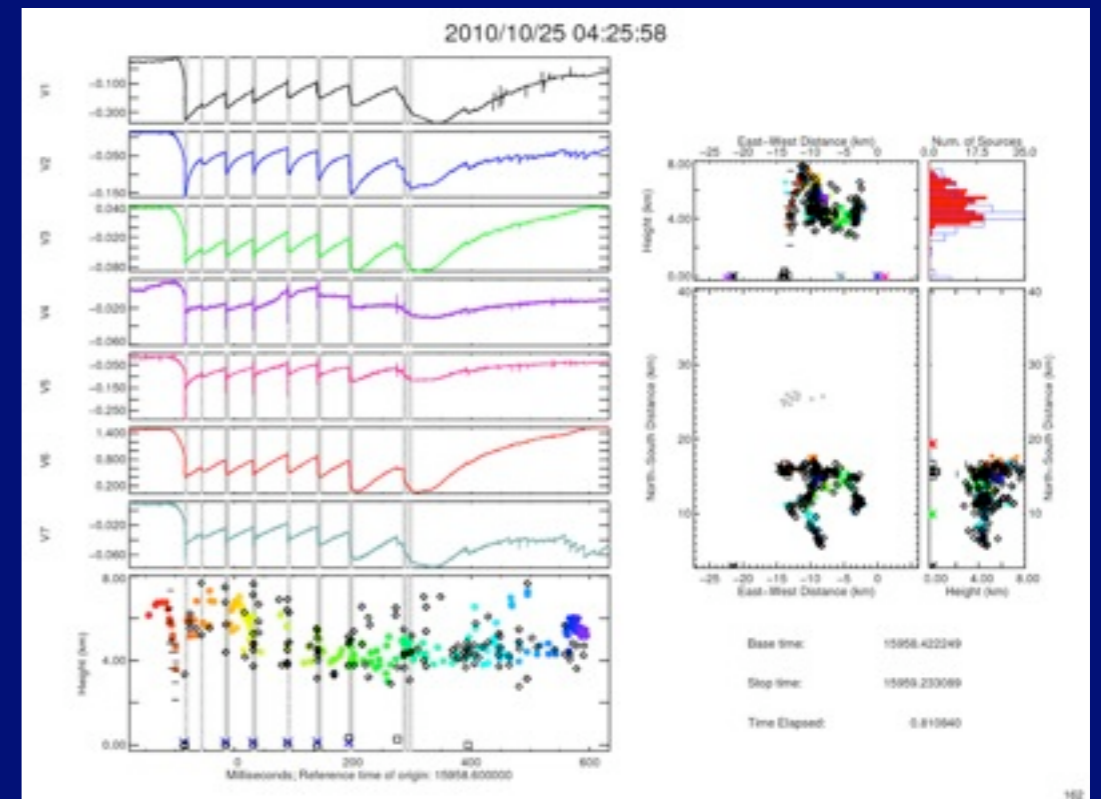
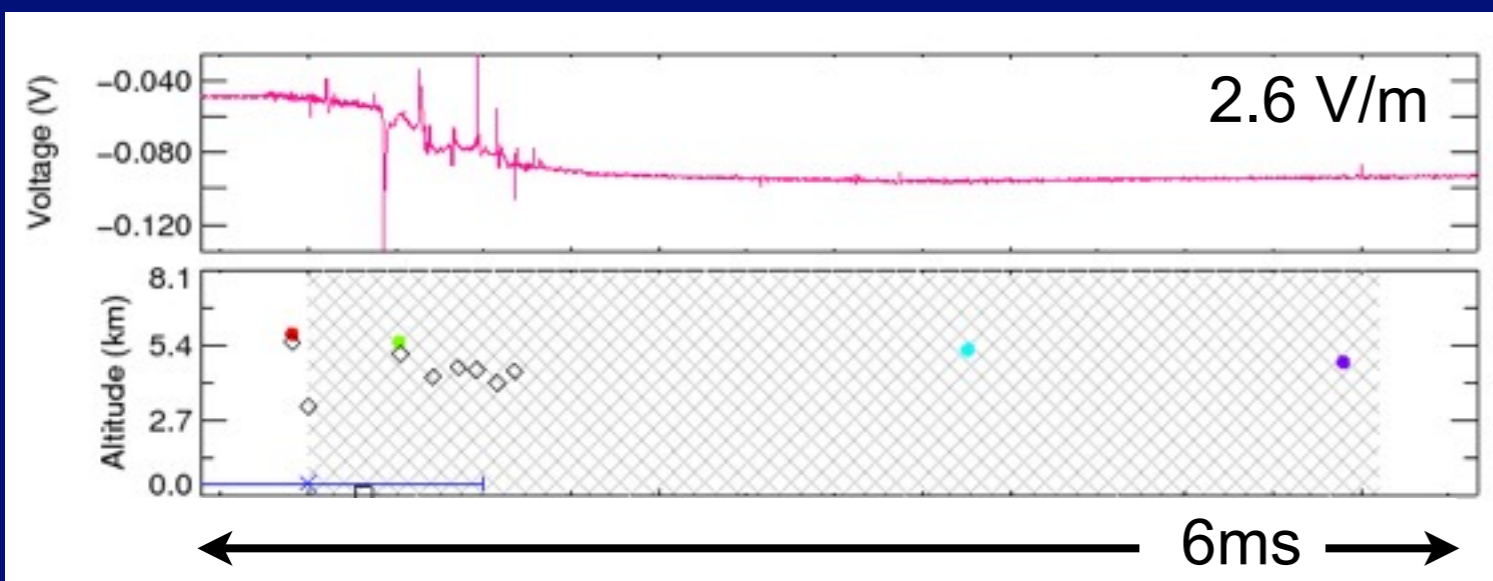
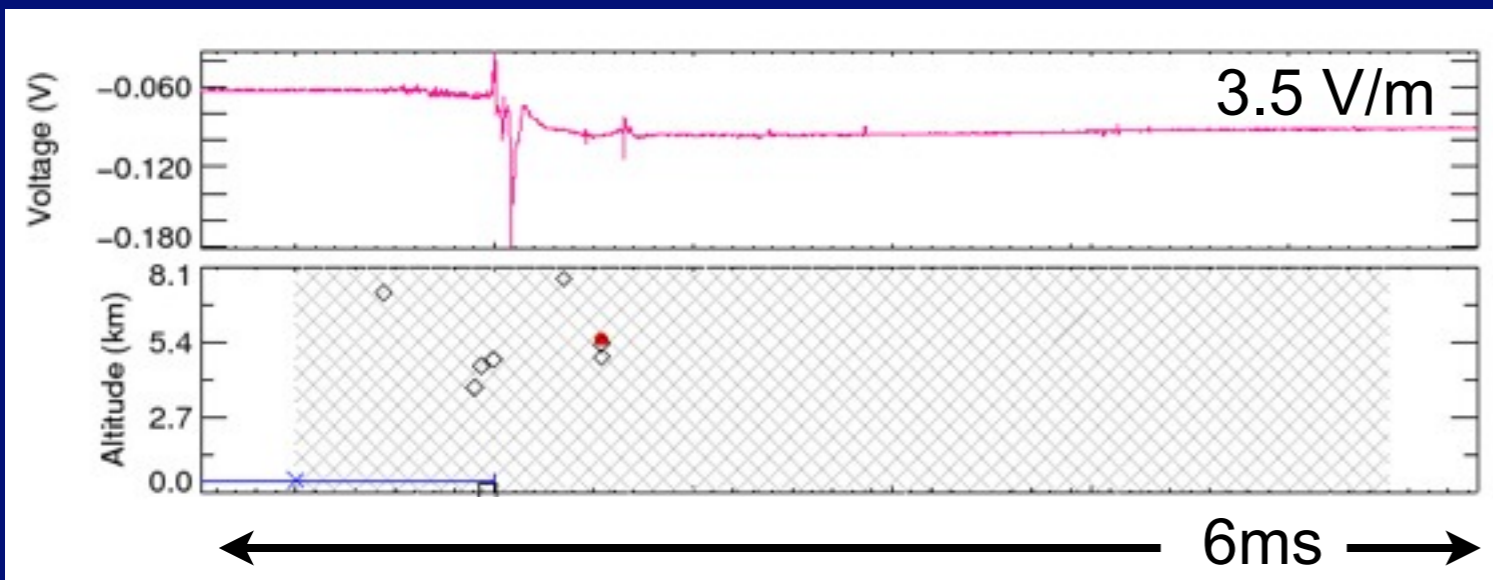
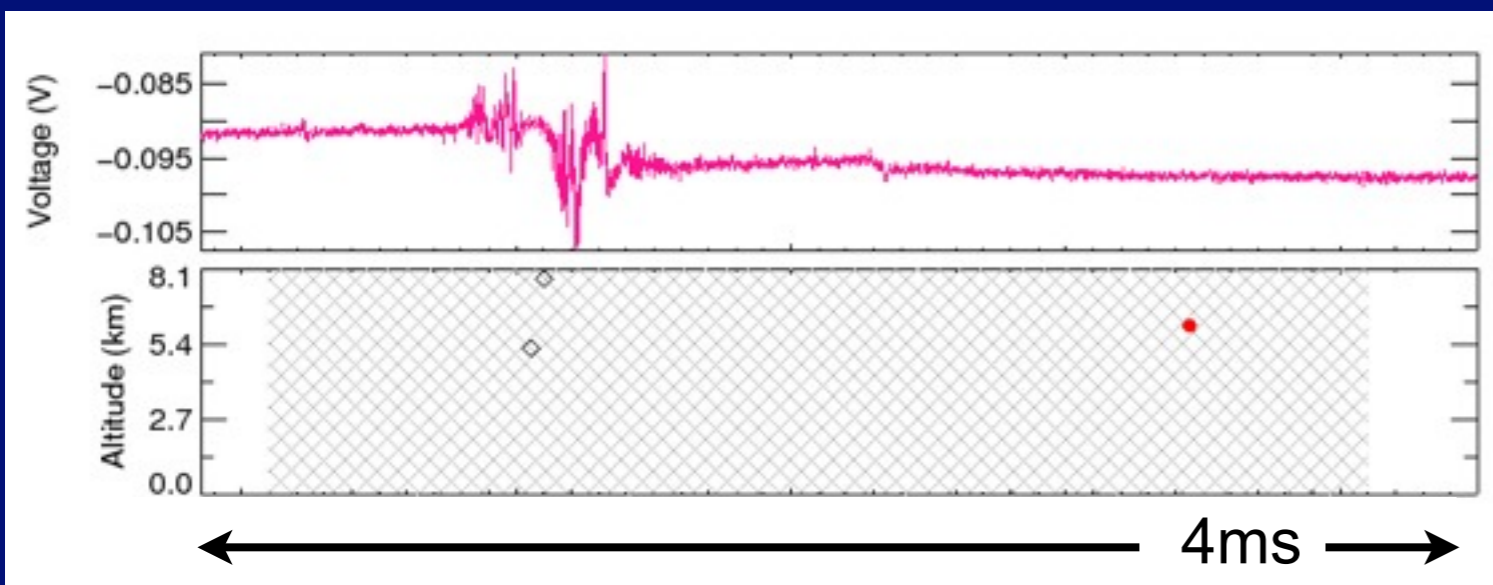


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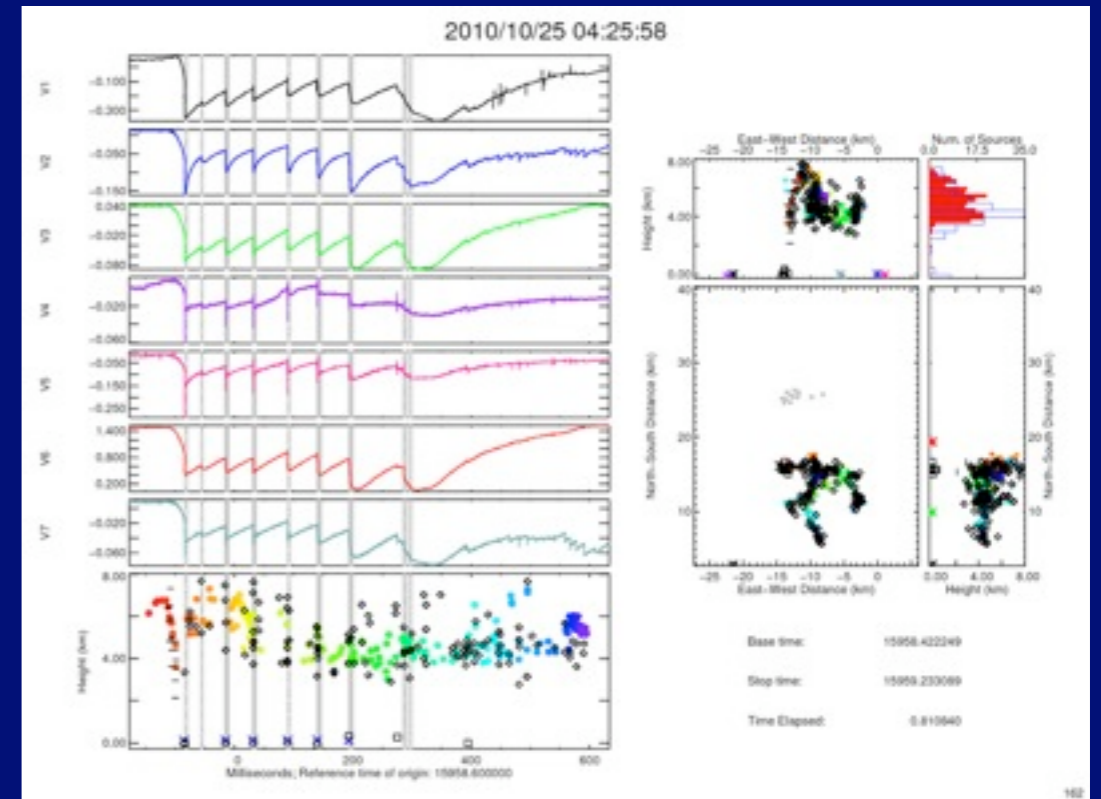
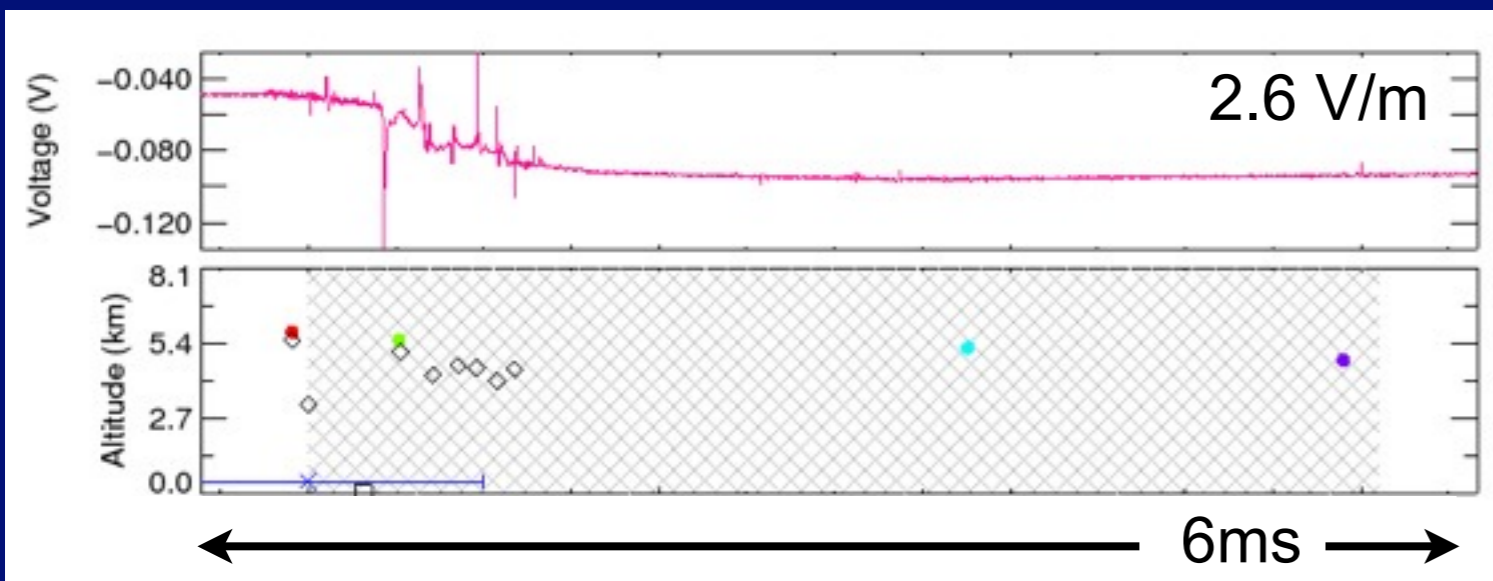
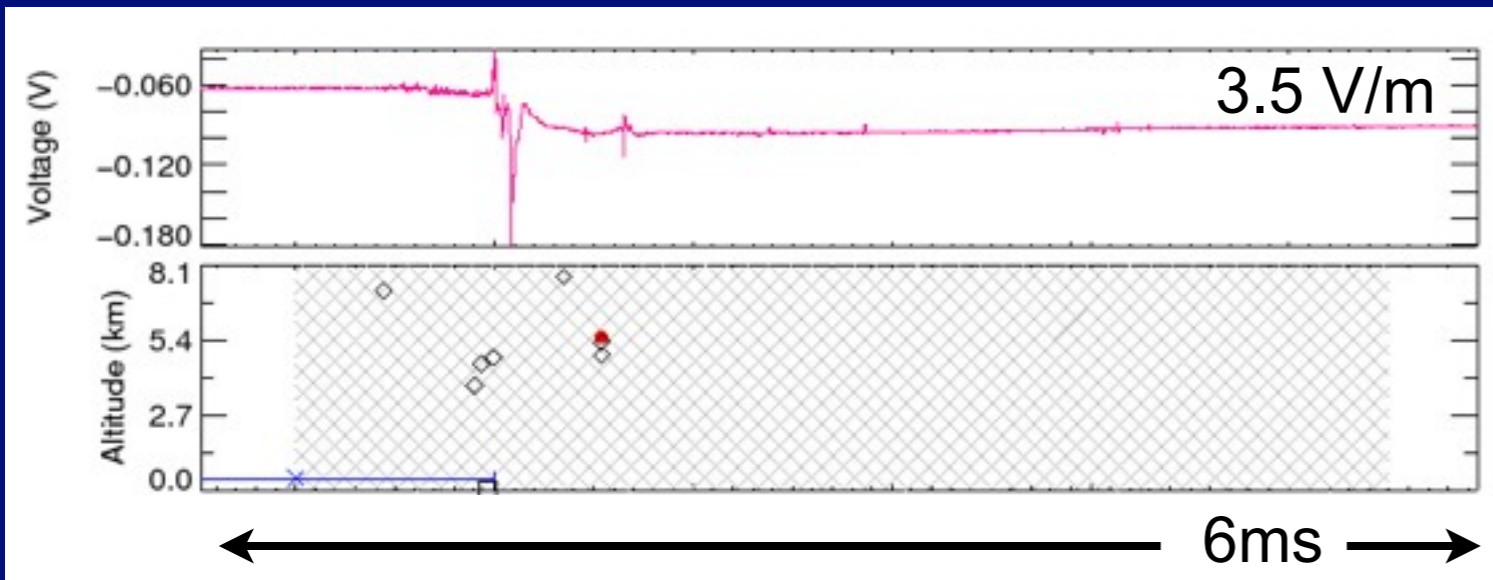
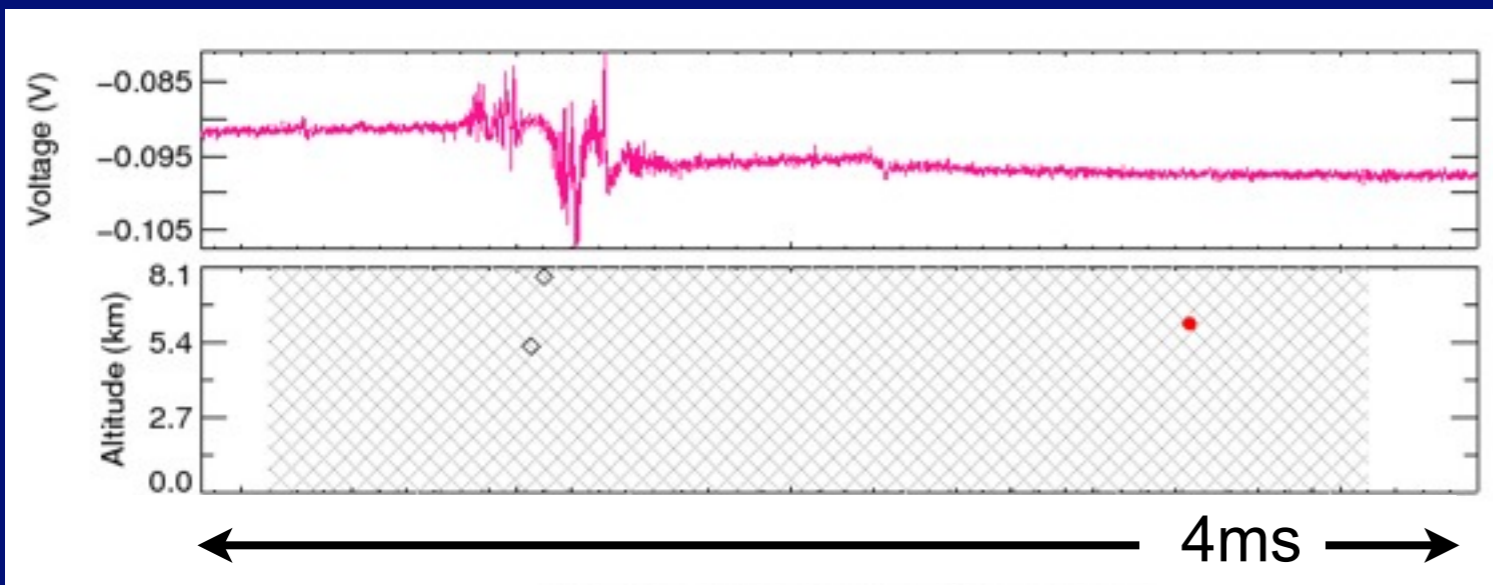
What is a hit?



What is a hit?

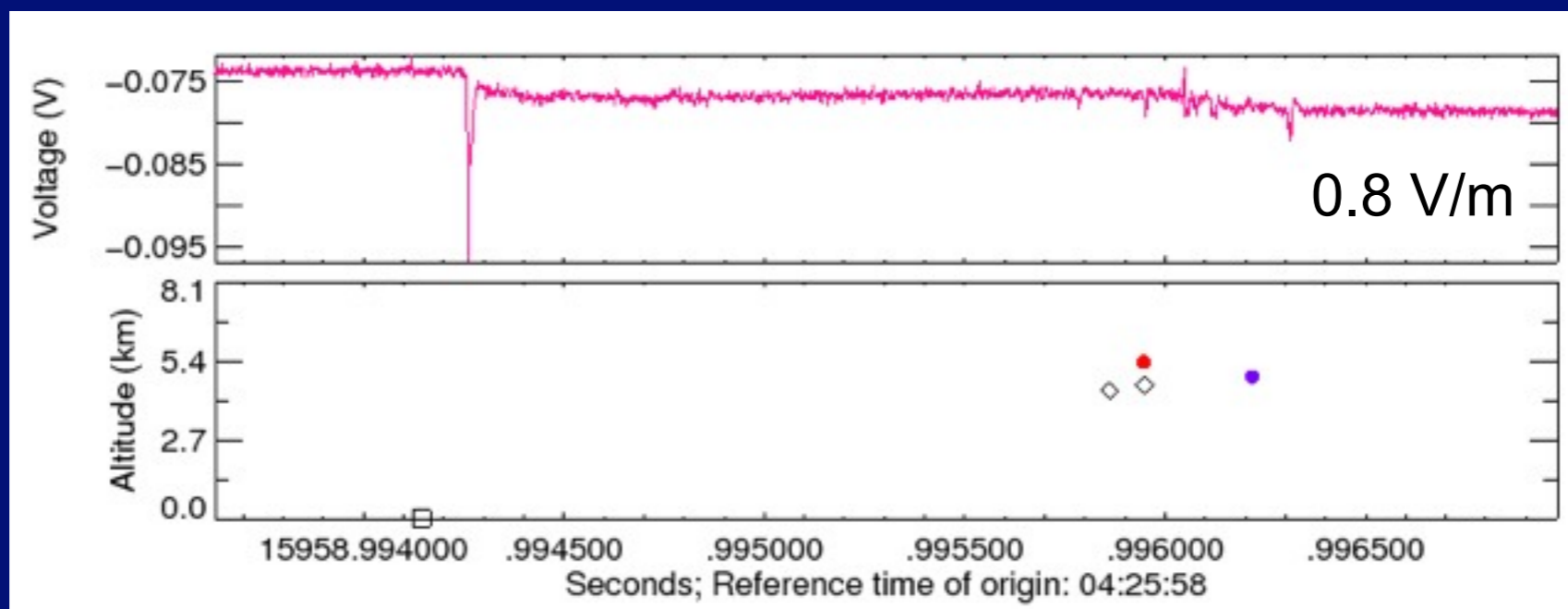
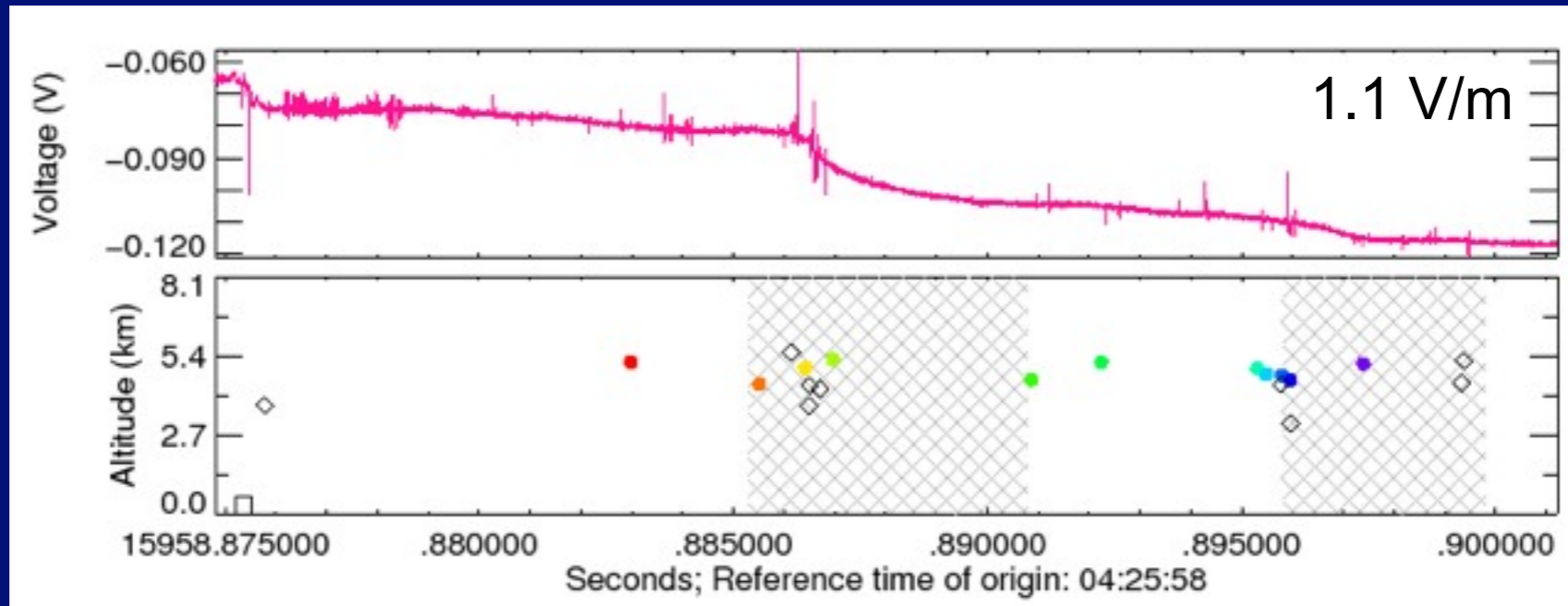


What is a hit?

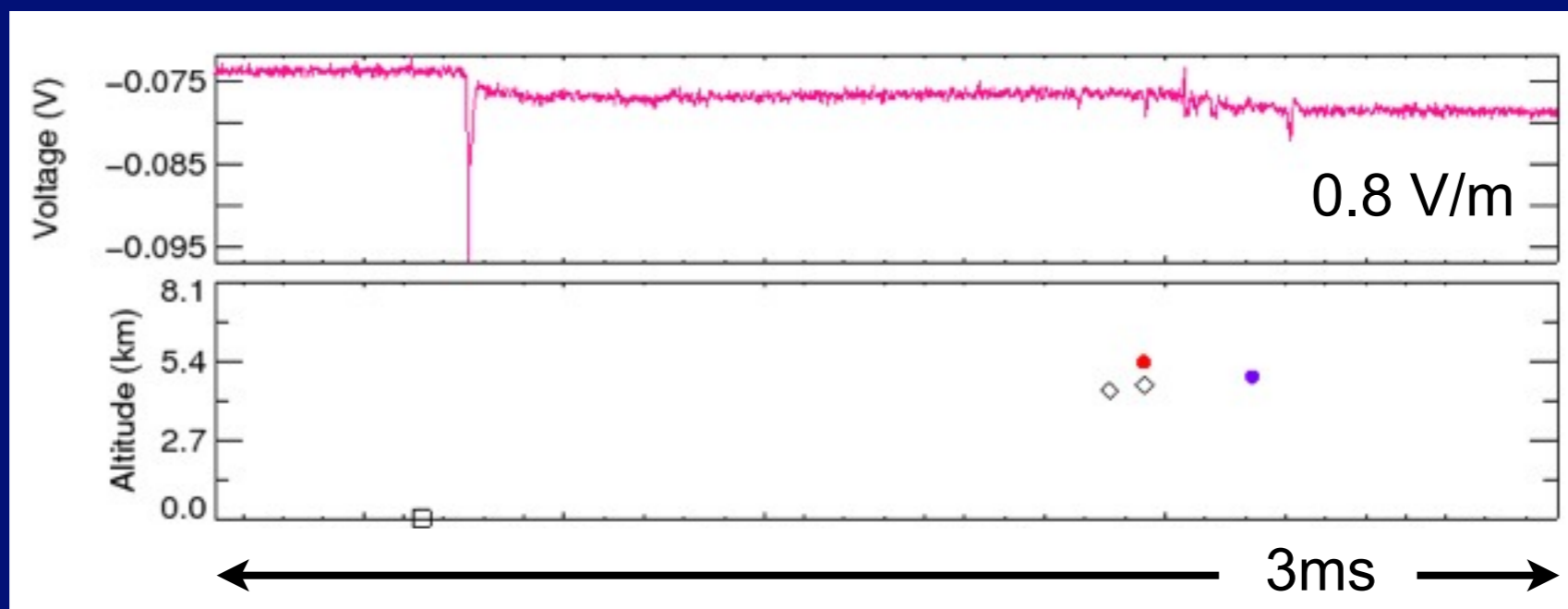
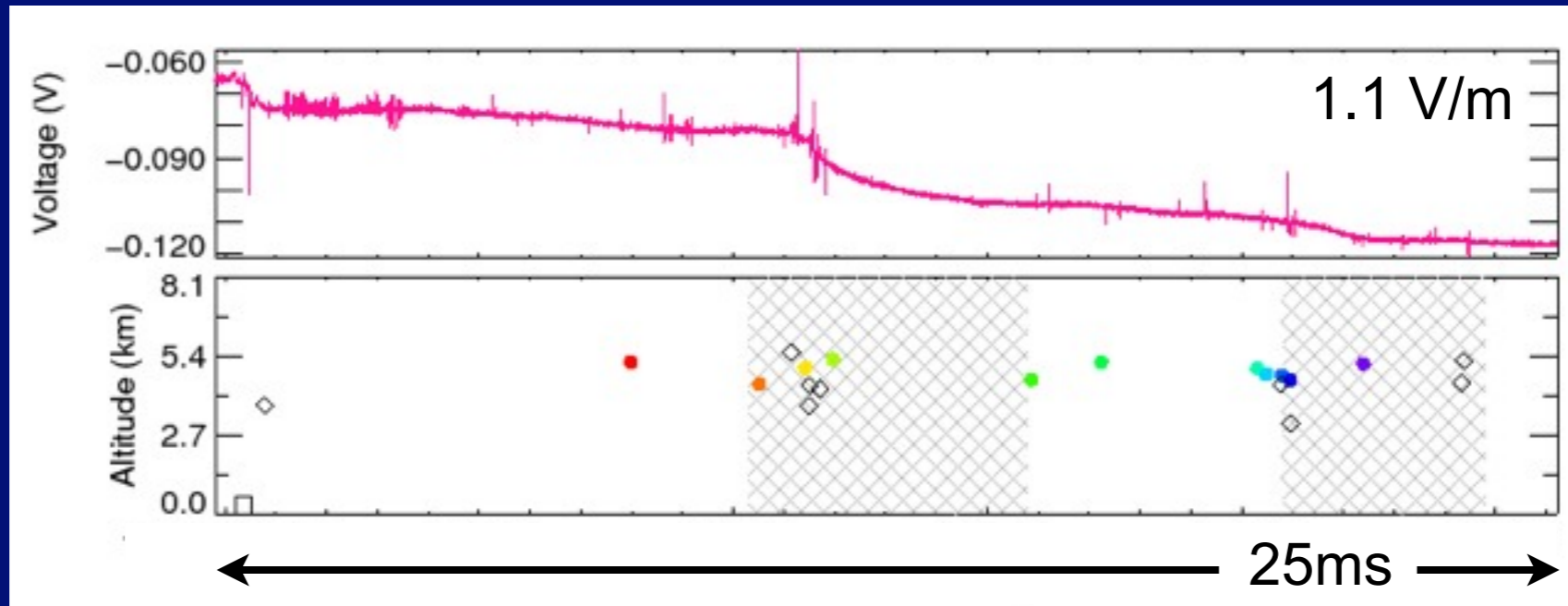


Wideband record indicates activity during LIS groups

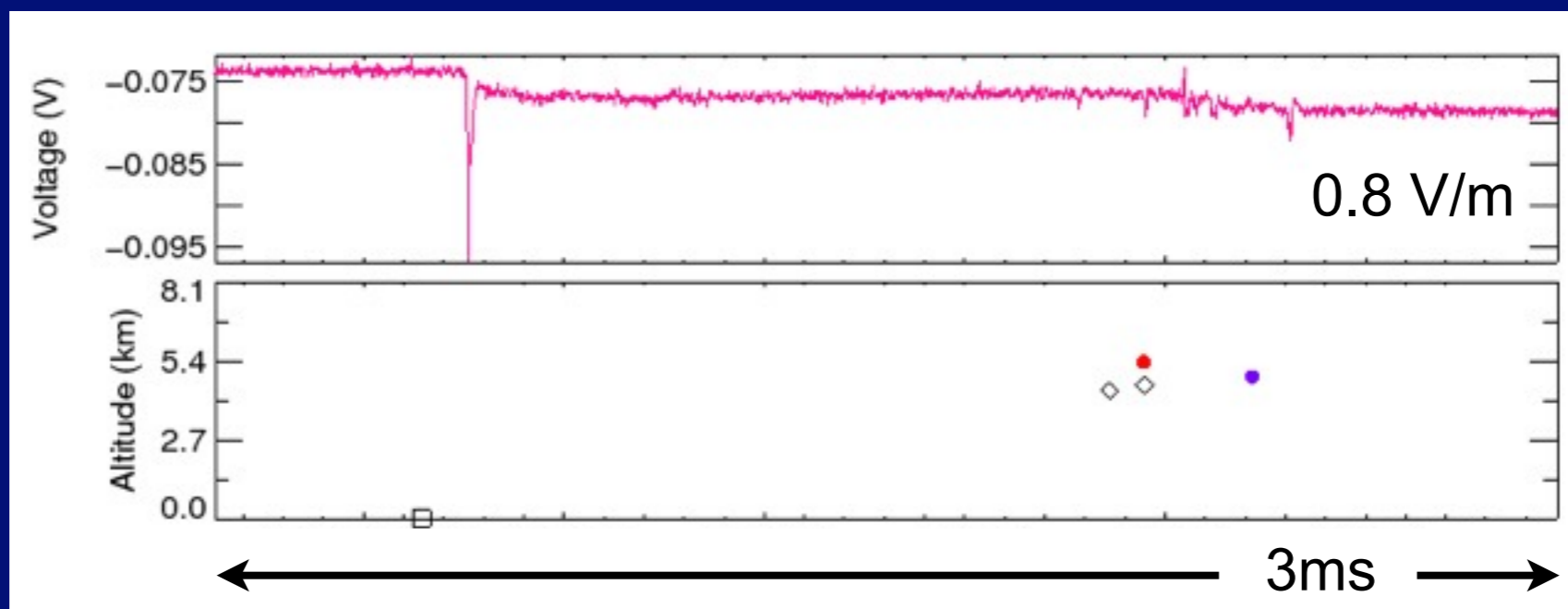
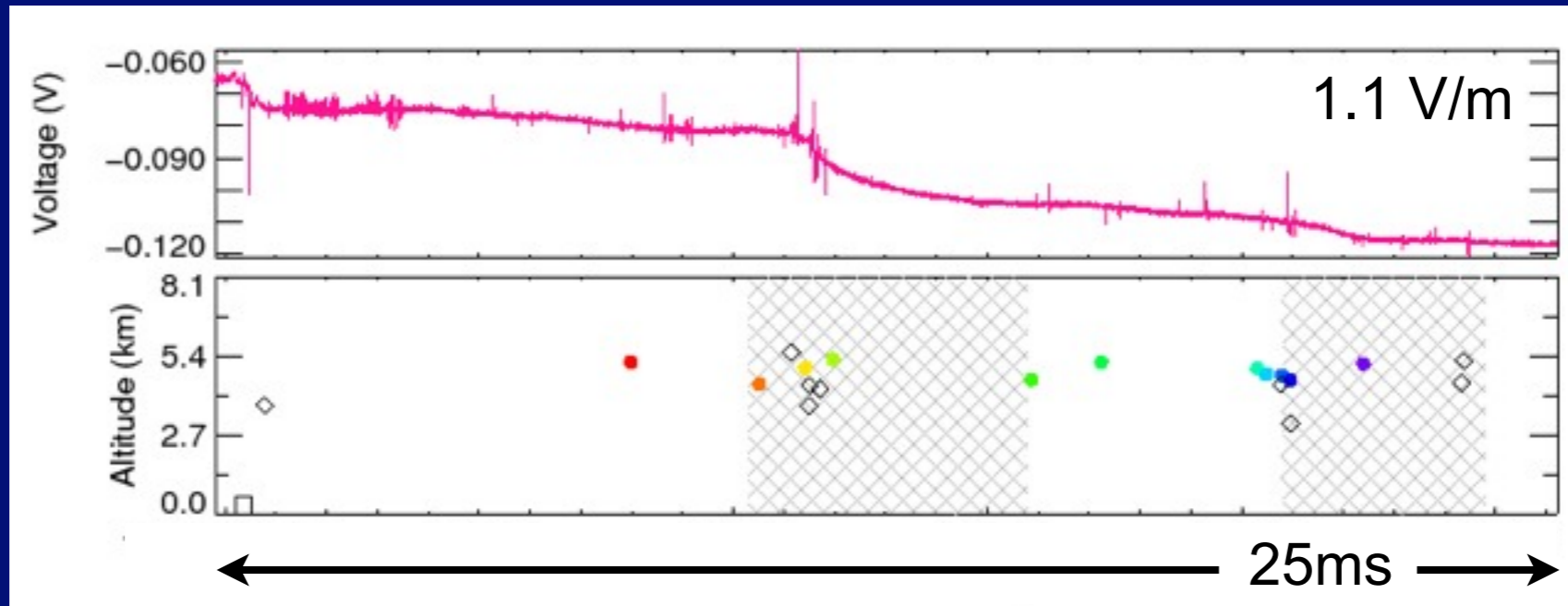
What is a miss?



What is a miss?

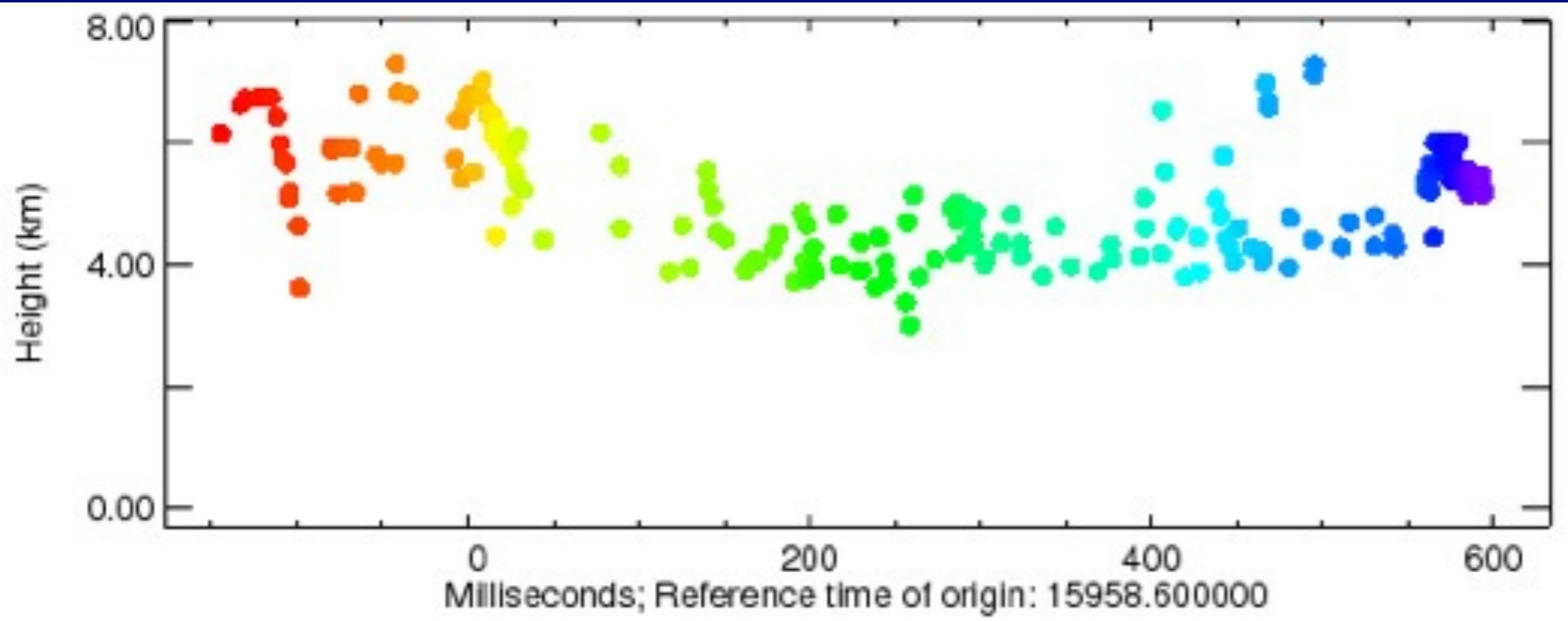


What is a miss?

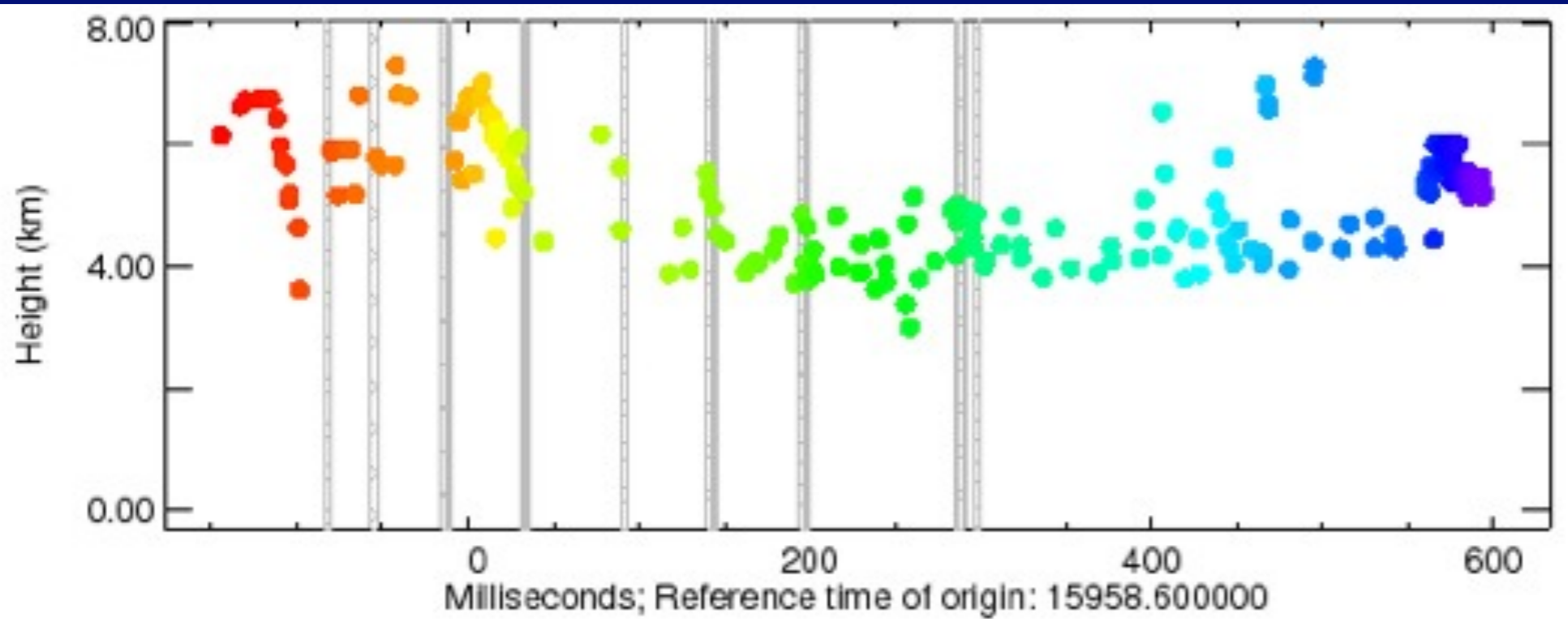


Electric field measurements indicate that LIS missed these strokes because they are not very energetic

What can VHF tell us?



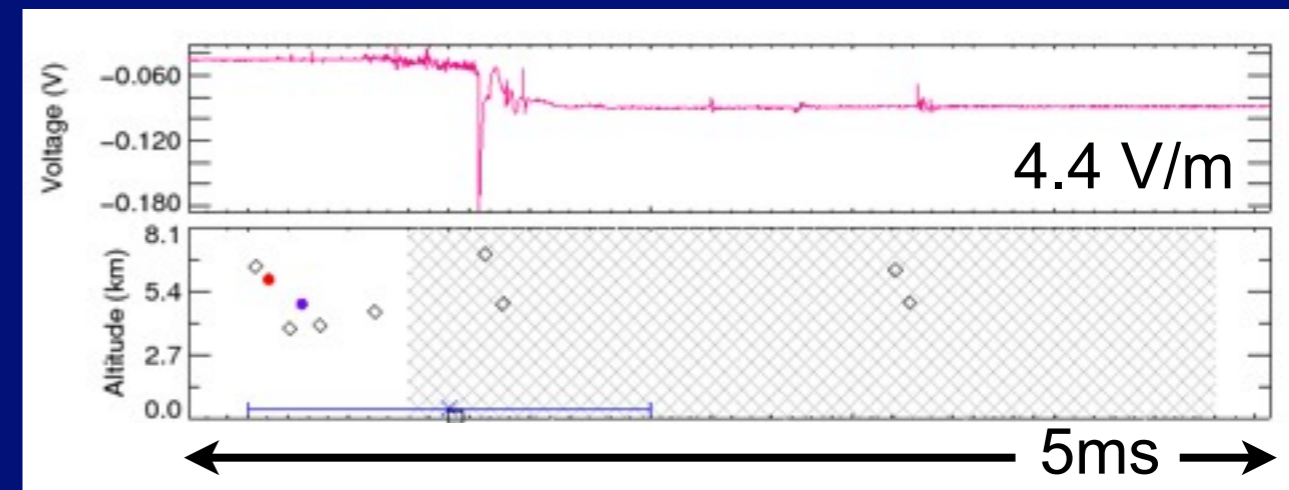
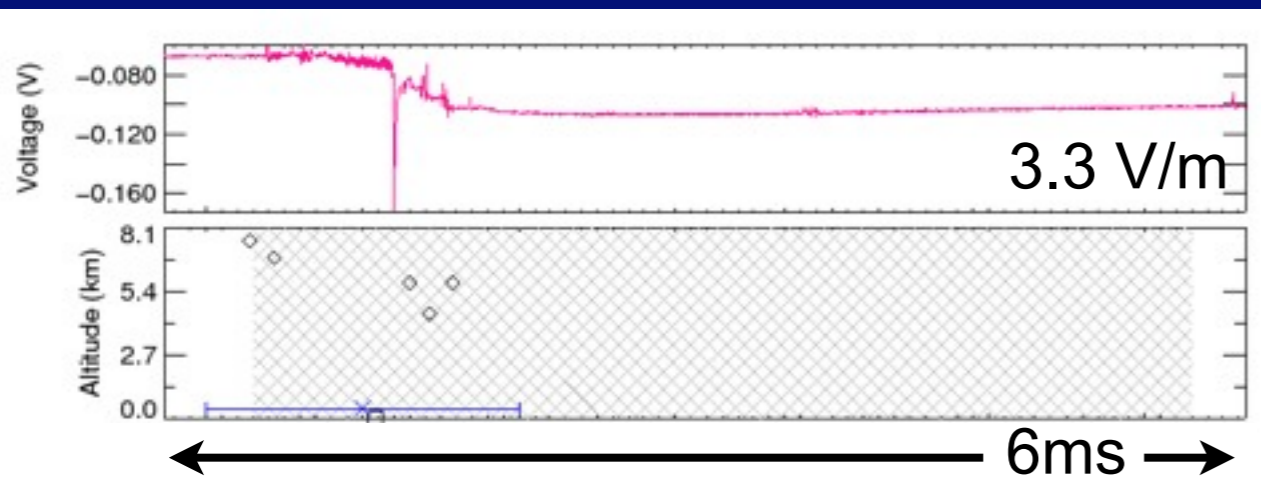
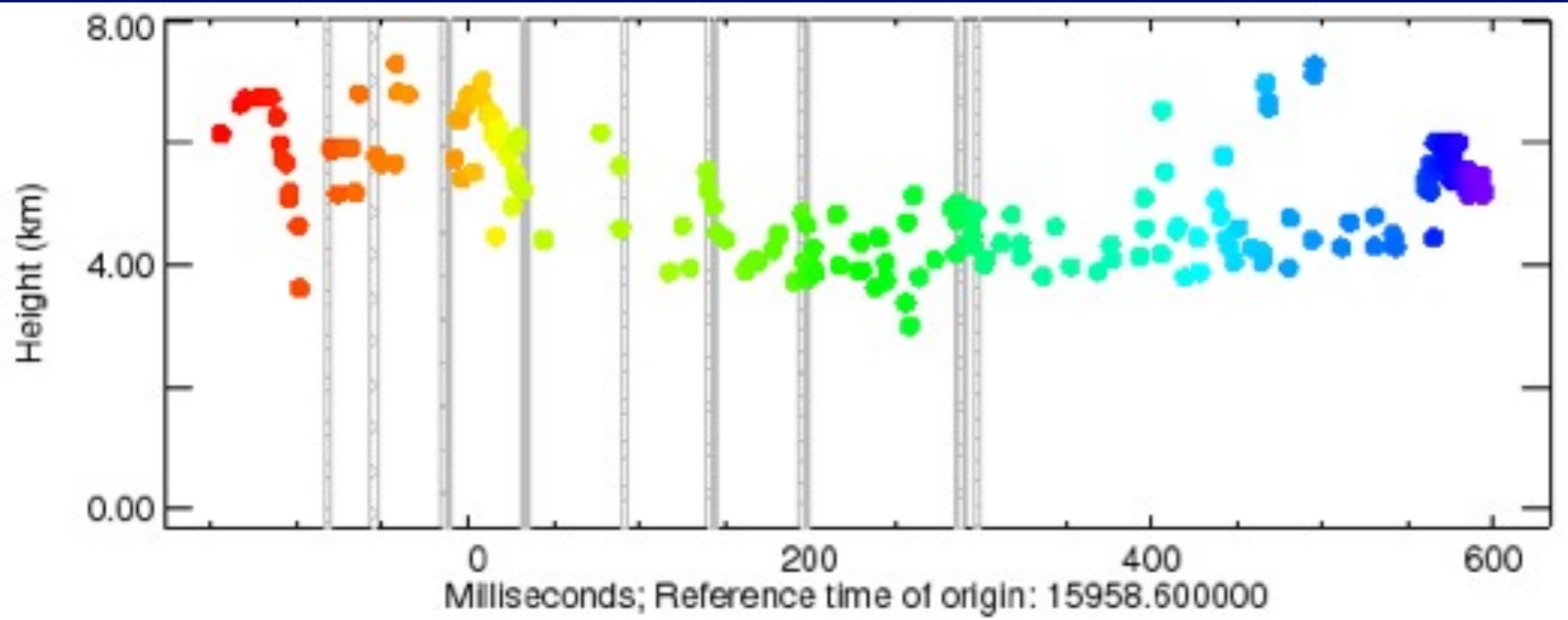
What can VHF tell us?



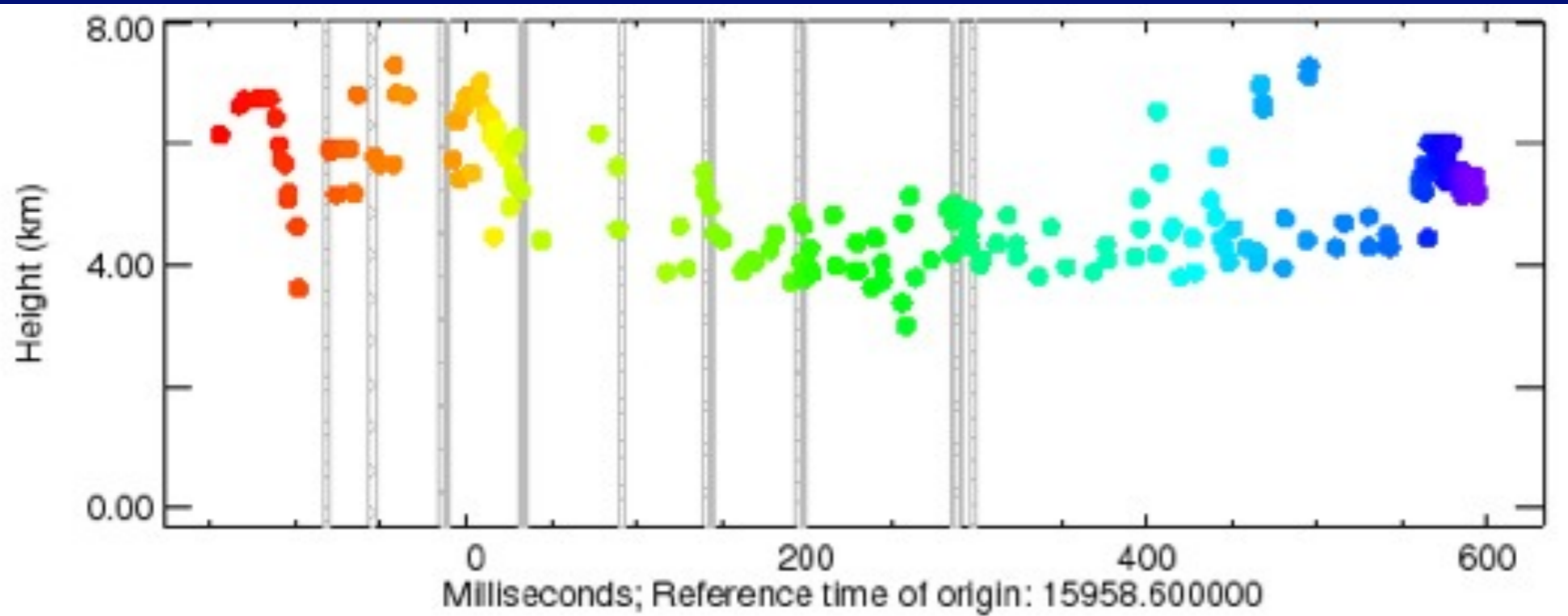
Nothing is really unique about VHF sources...

What can VHF tell us?

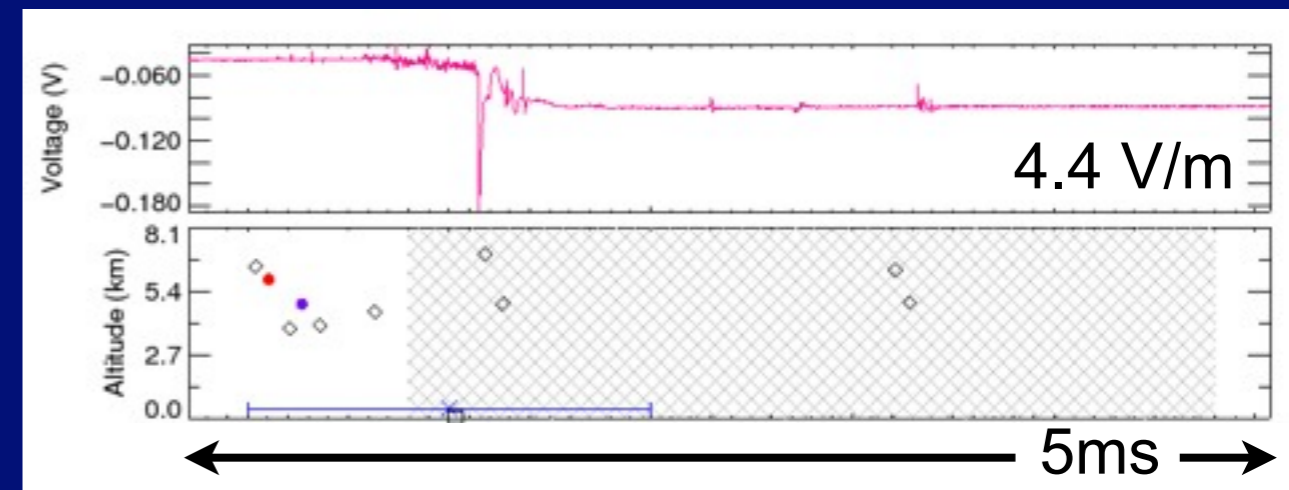
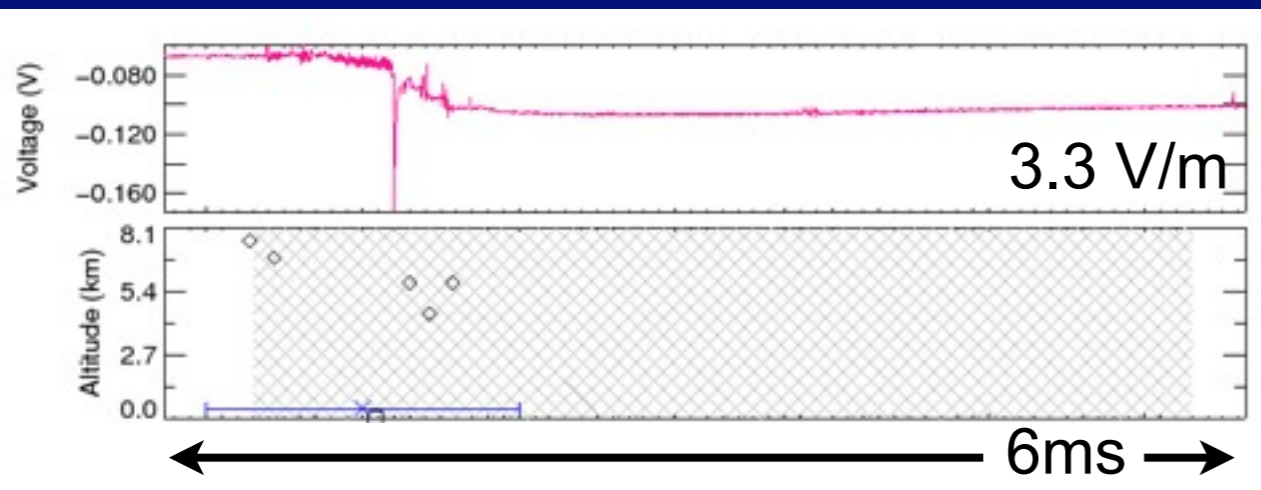
Nothing is really unique about VHF sources...



What can VHF tell us?

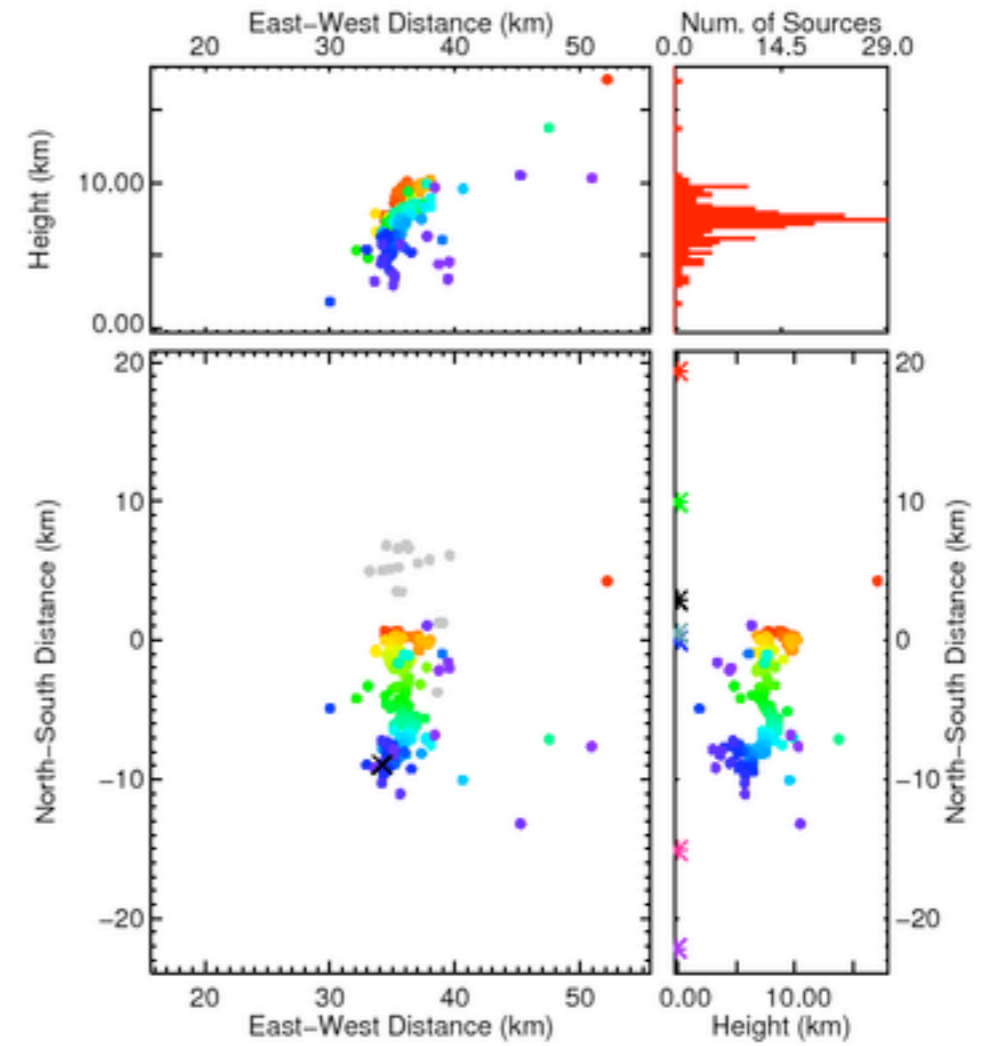
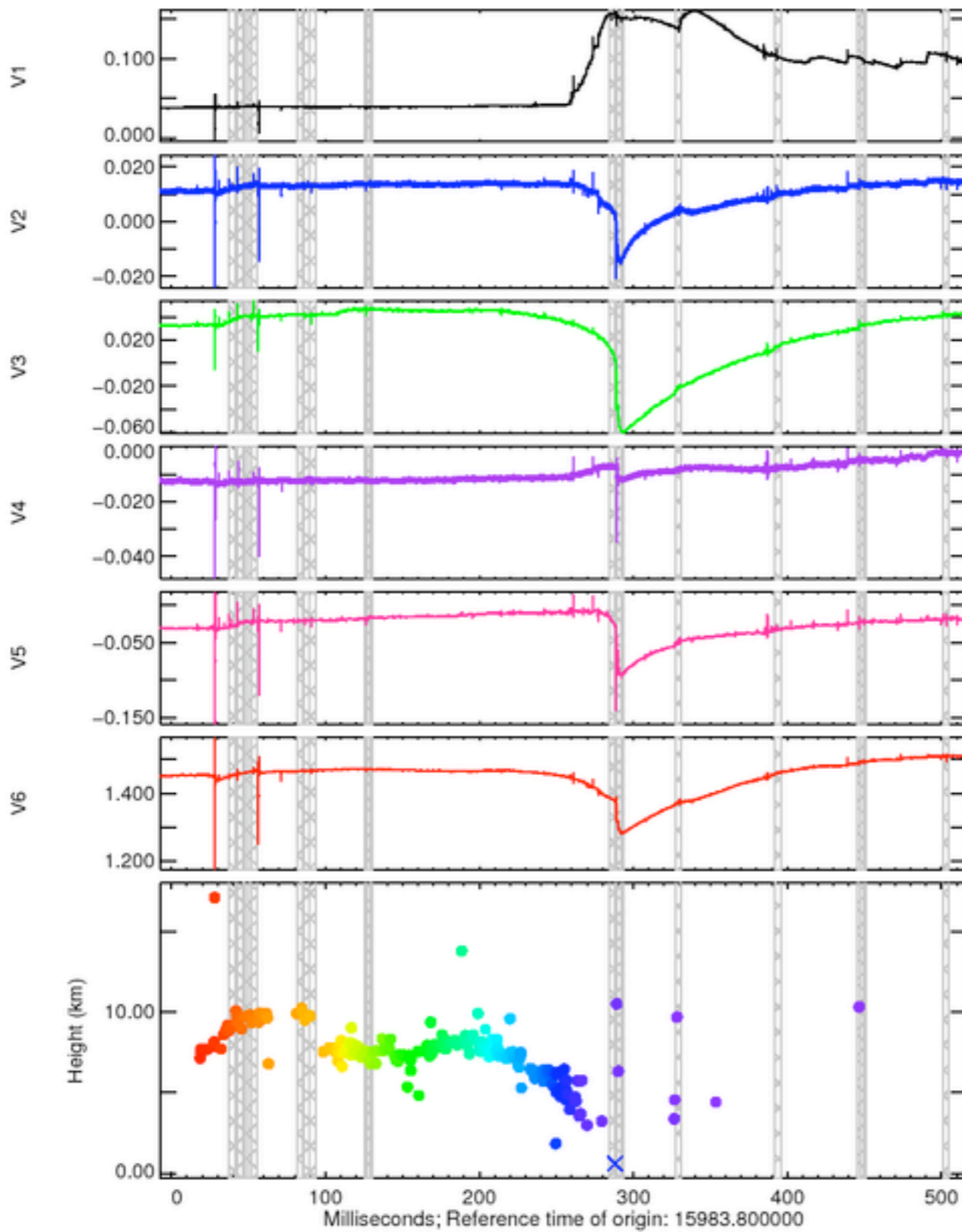


Nothing is really unique about VHF sources...



...in fact, there are LIS groups without any detected VHF sources!

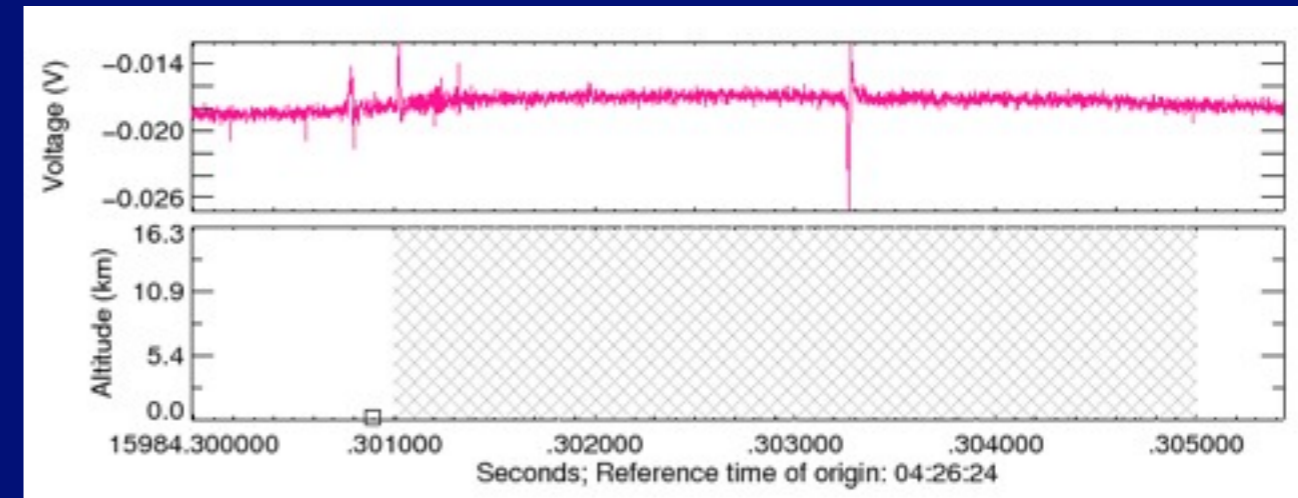
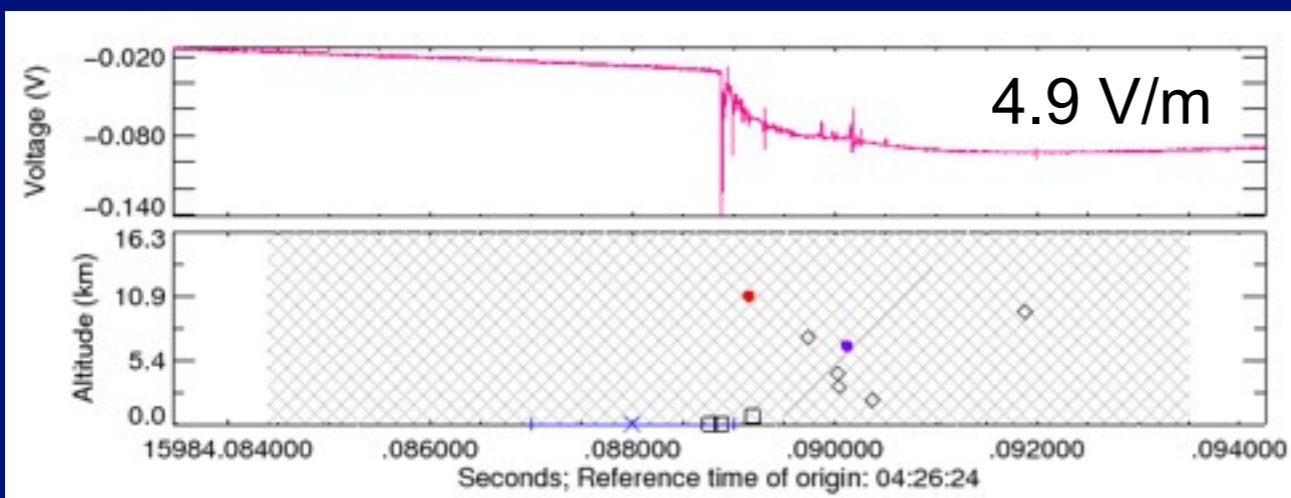
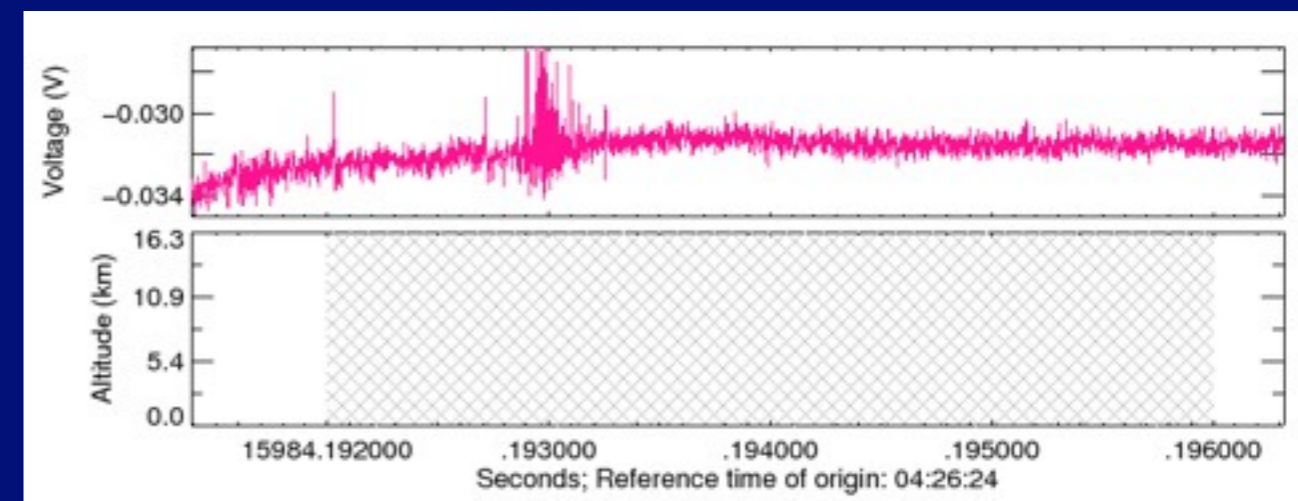
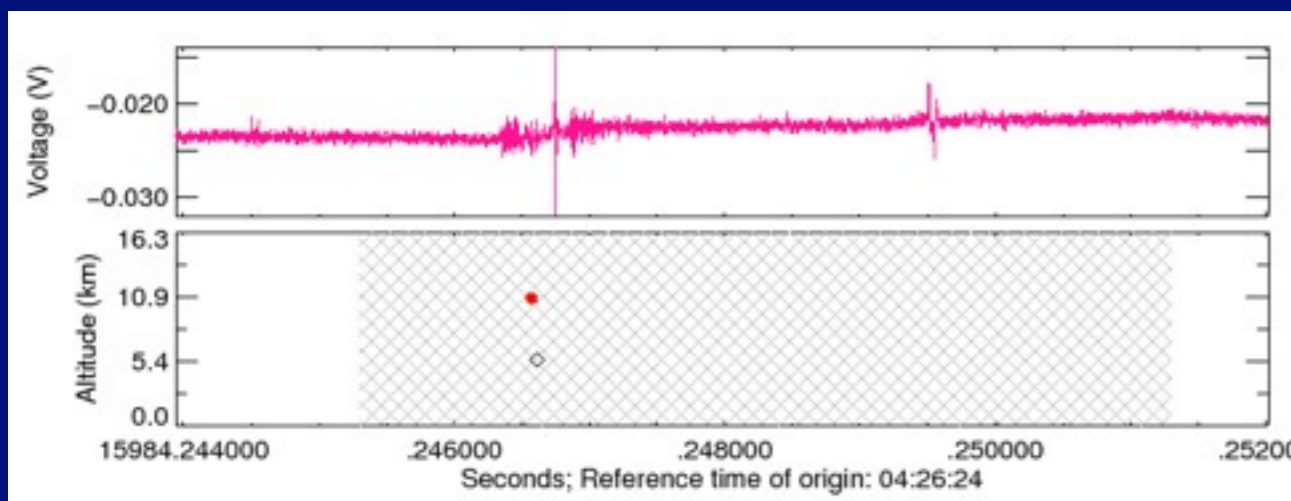
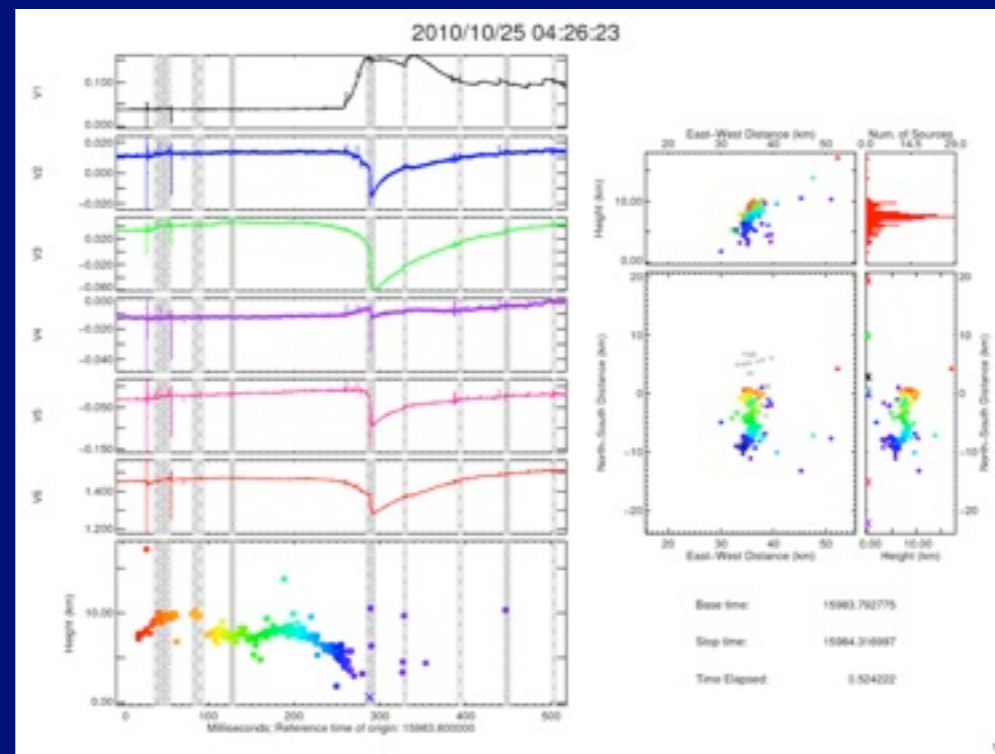
2010/10/25 04:26:23

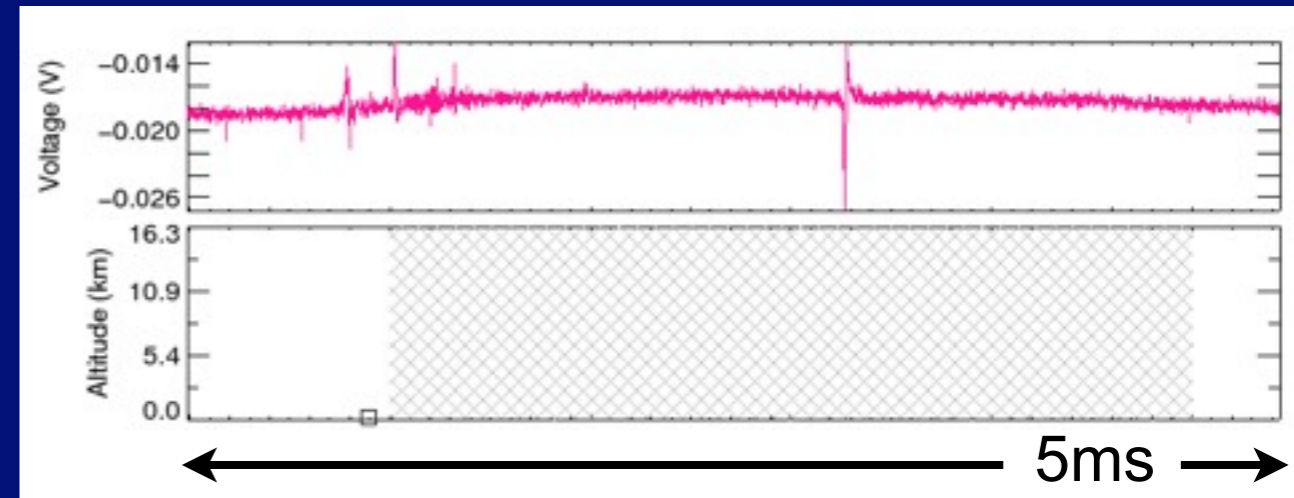
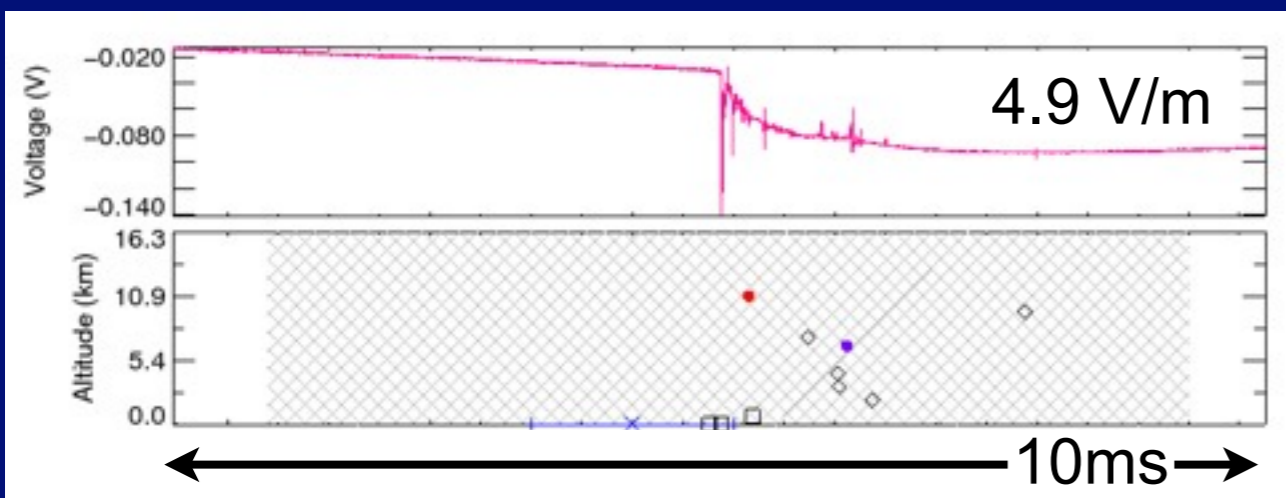
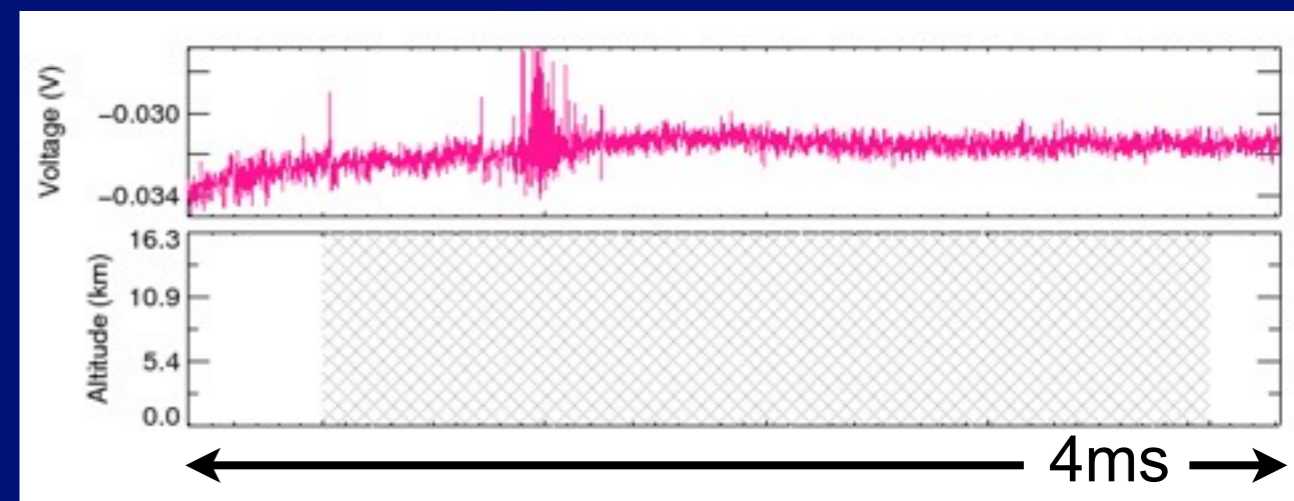
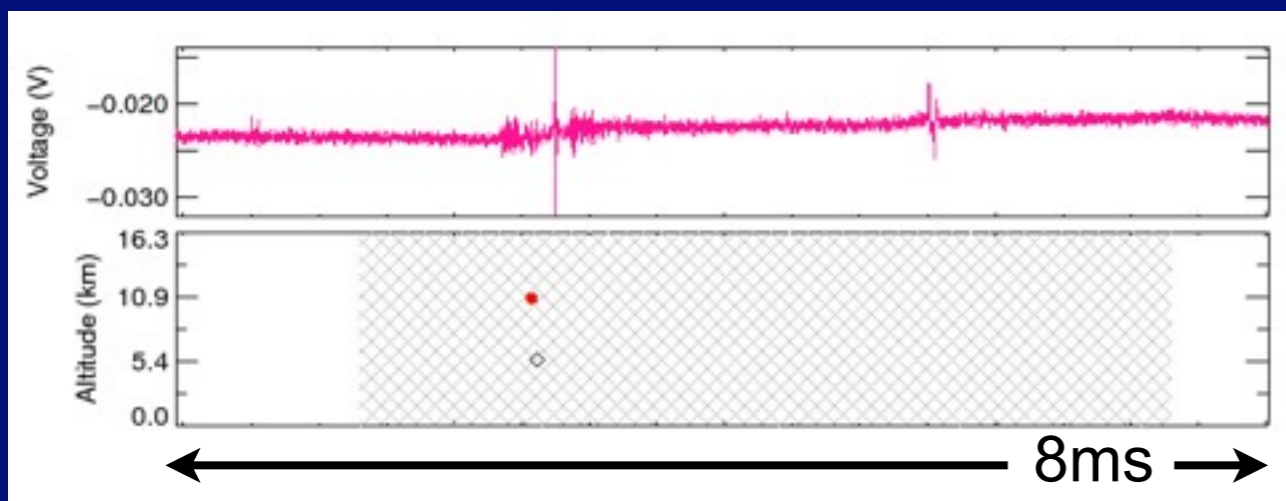
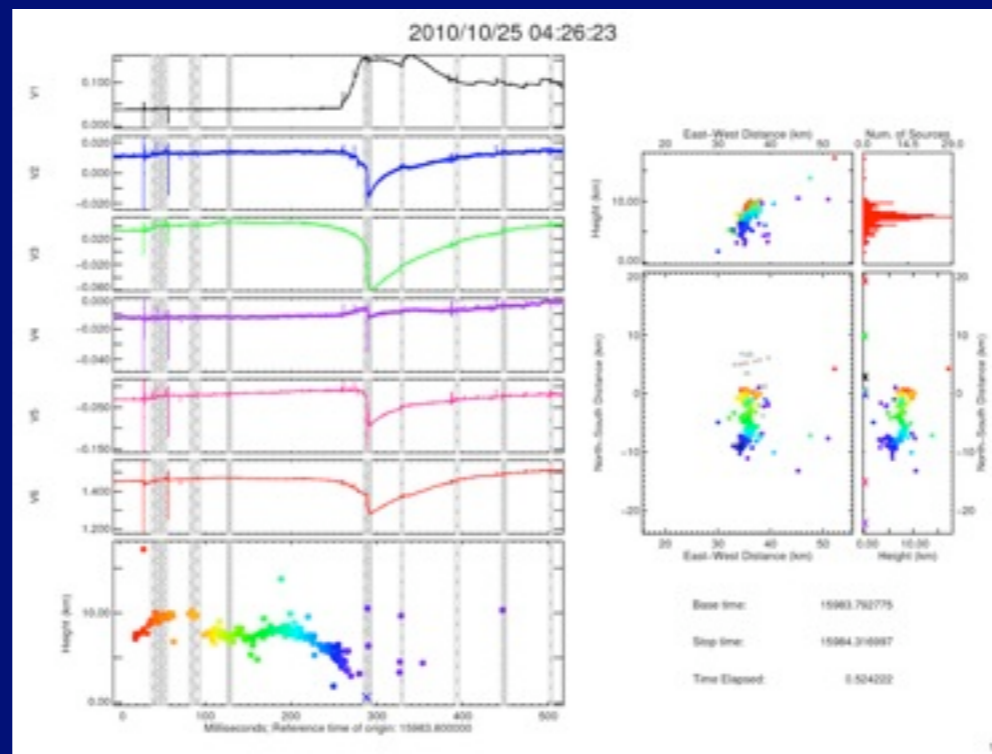


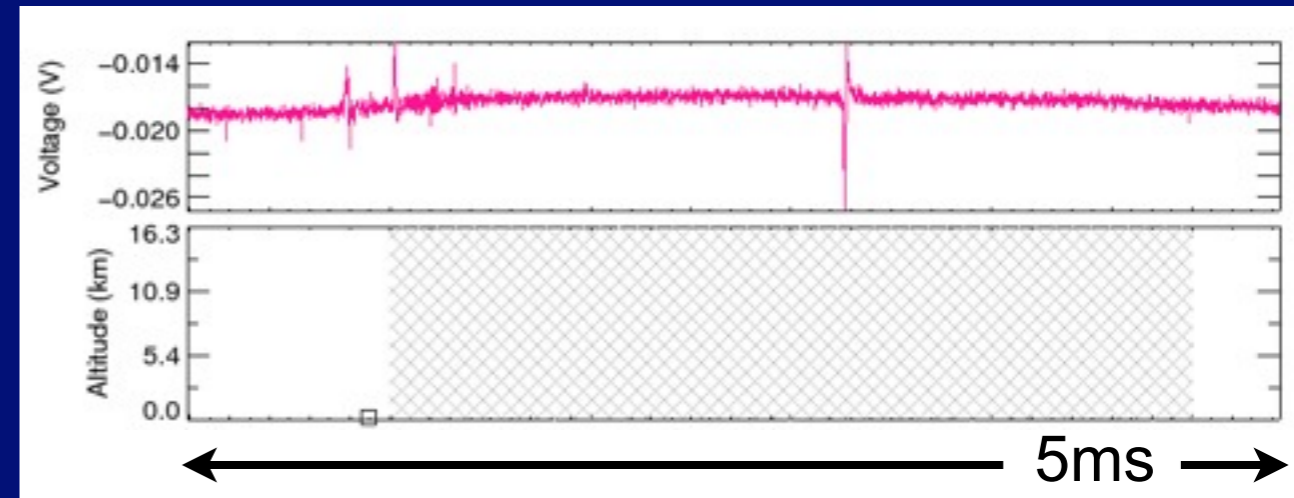
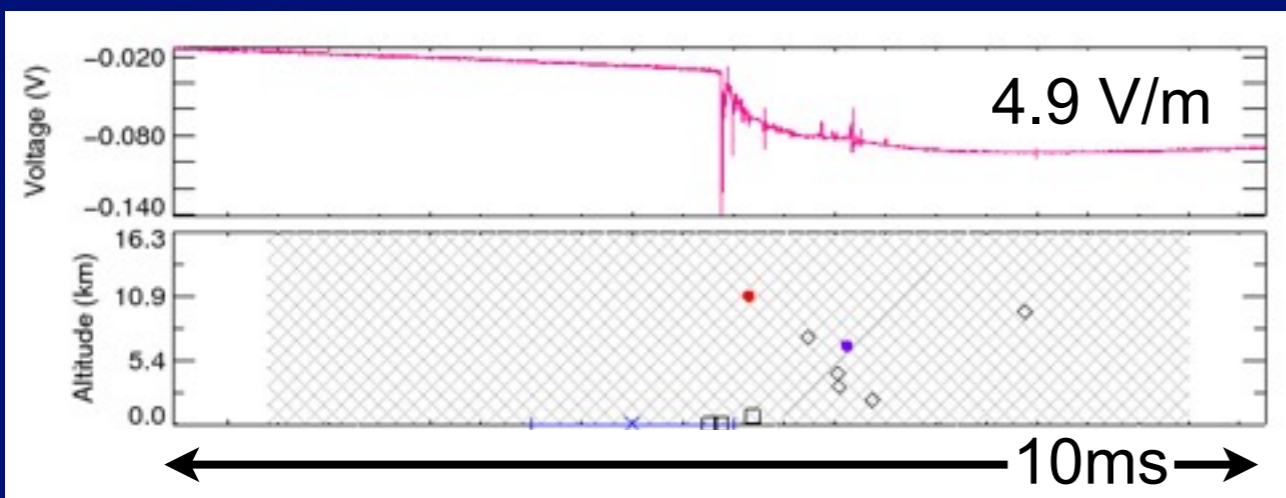
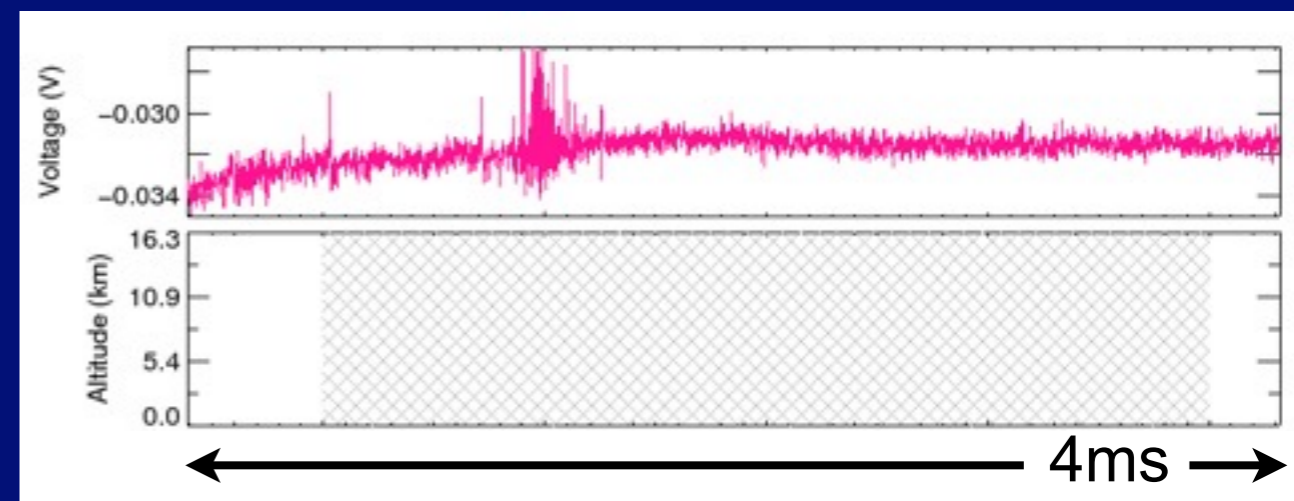
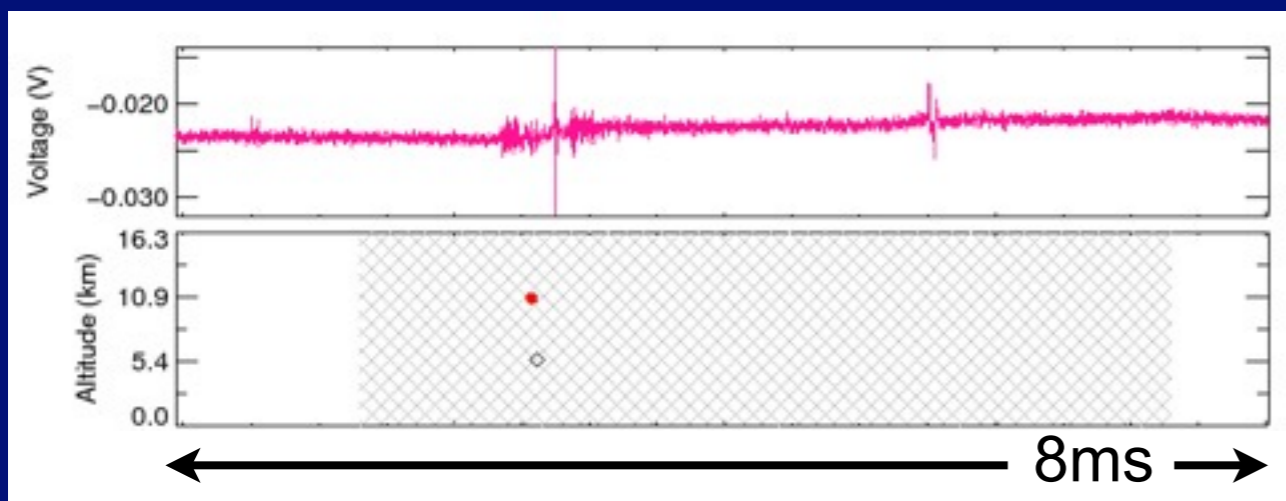
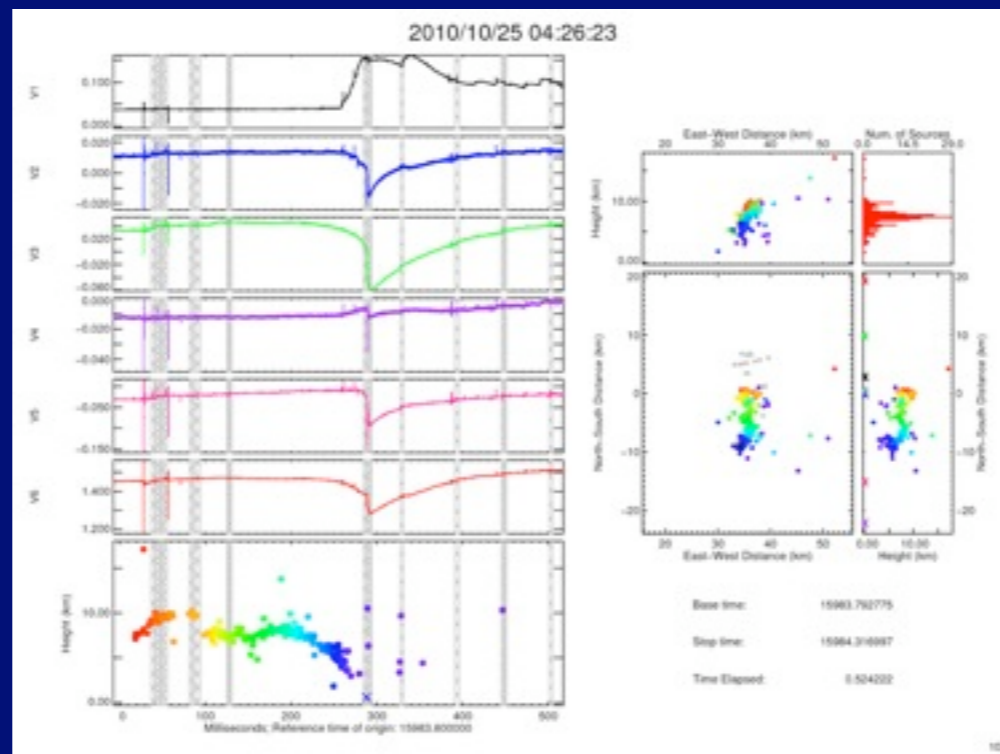
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Stop time: 15984.316997

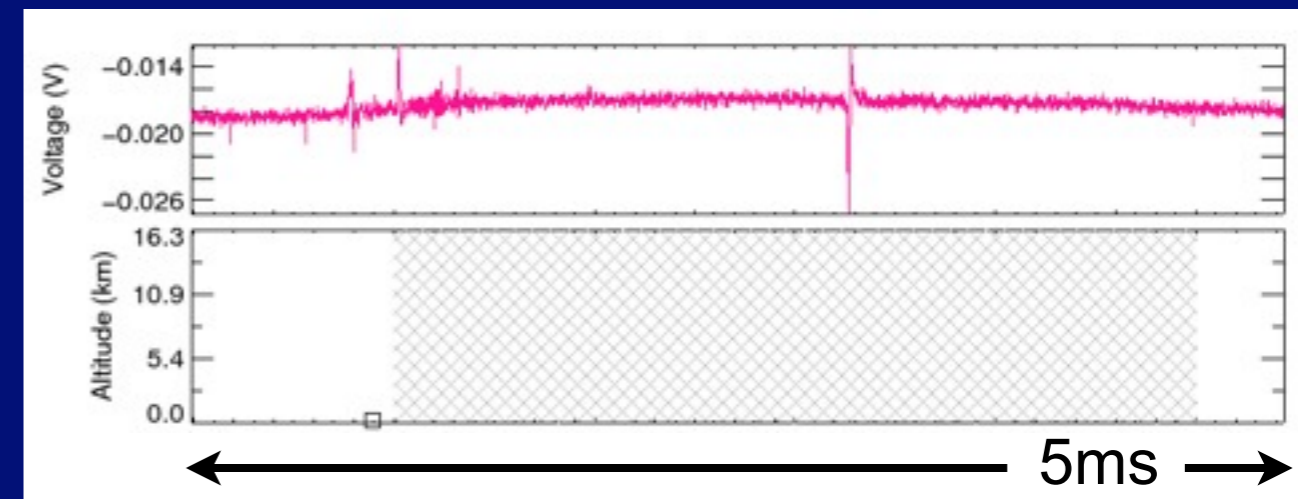
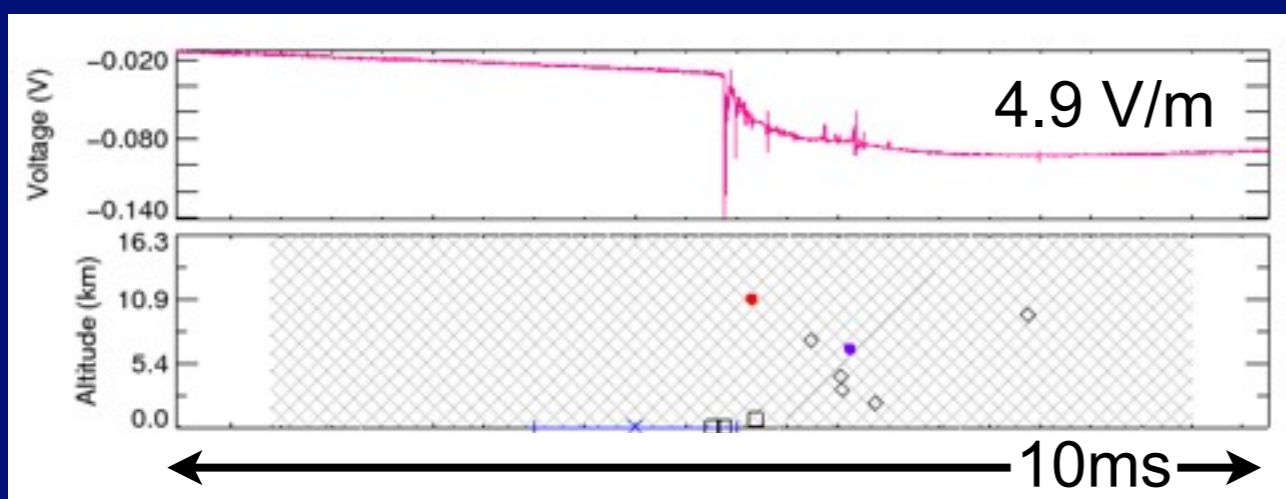
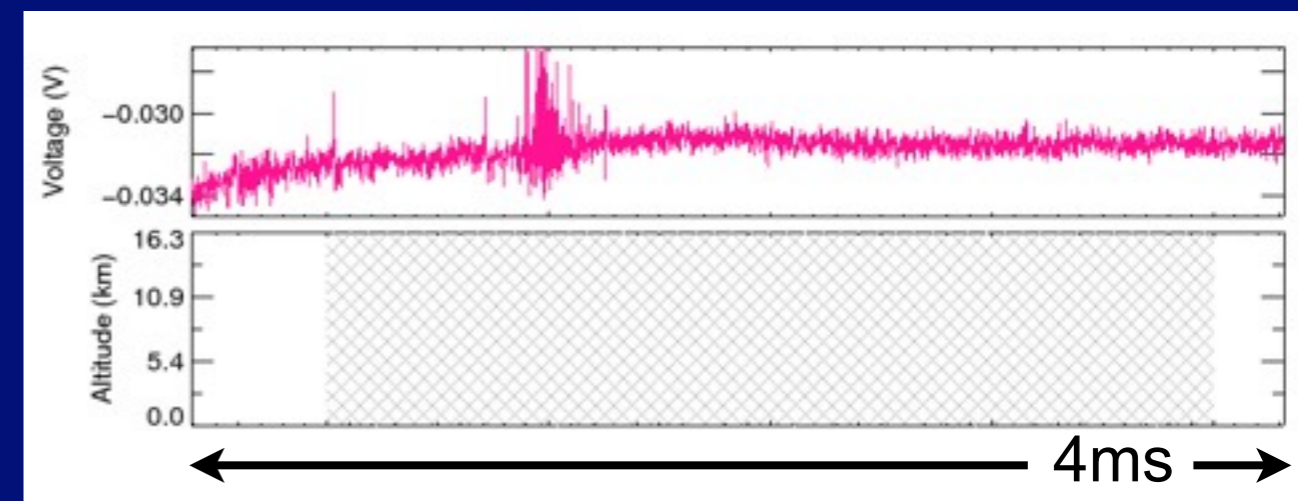
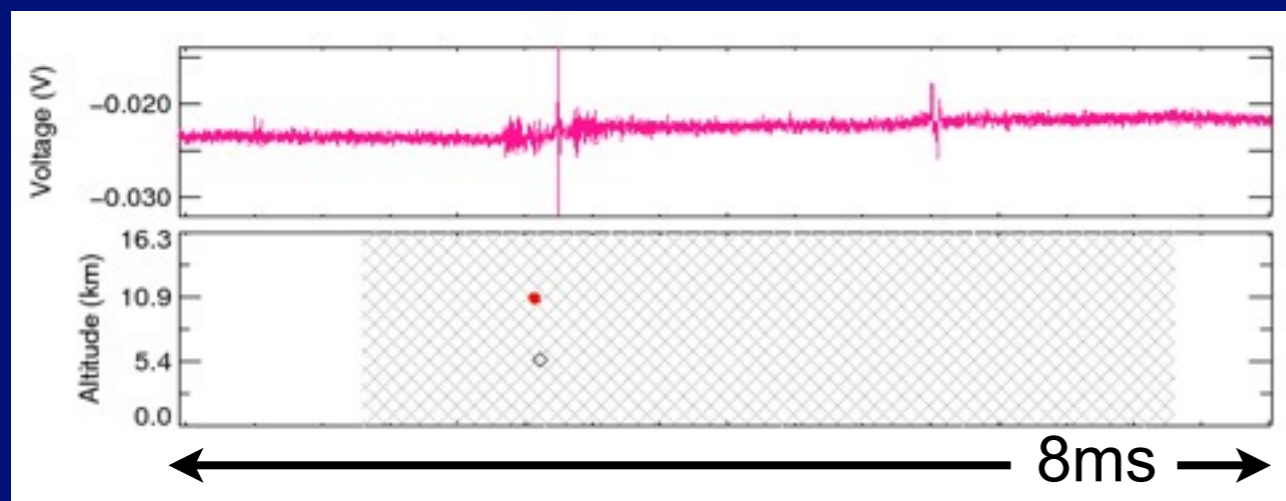
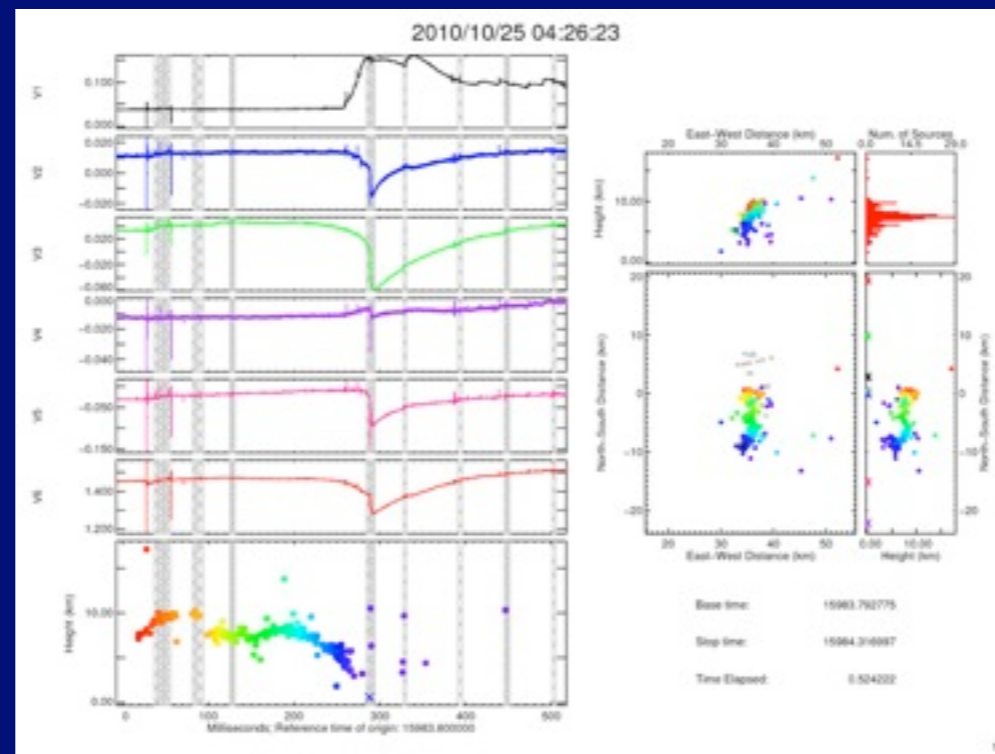
Time Elapsed: 0.524222







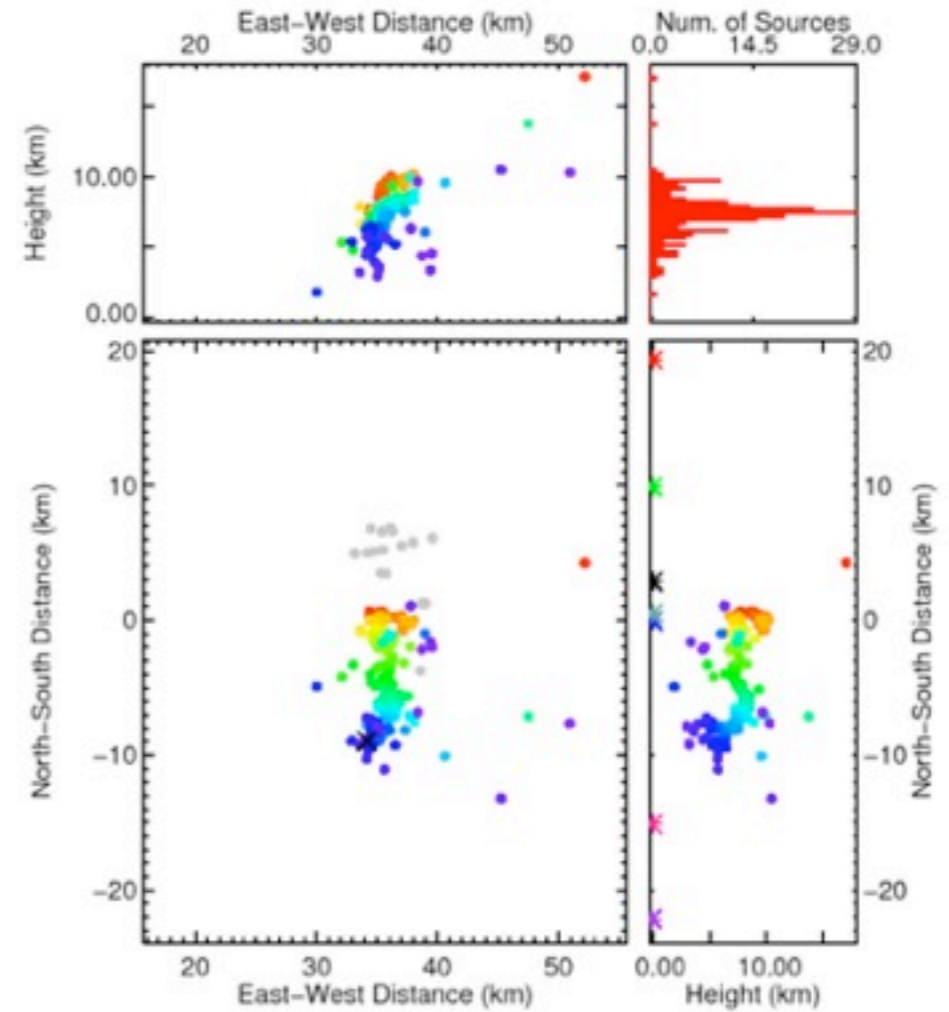
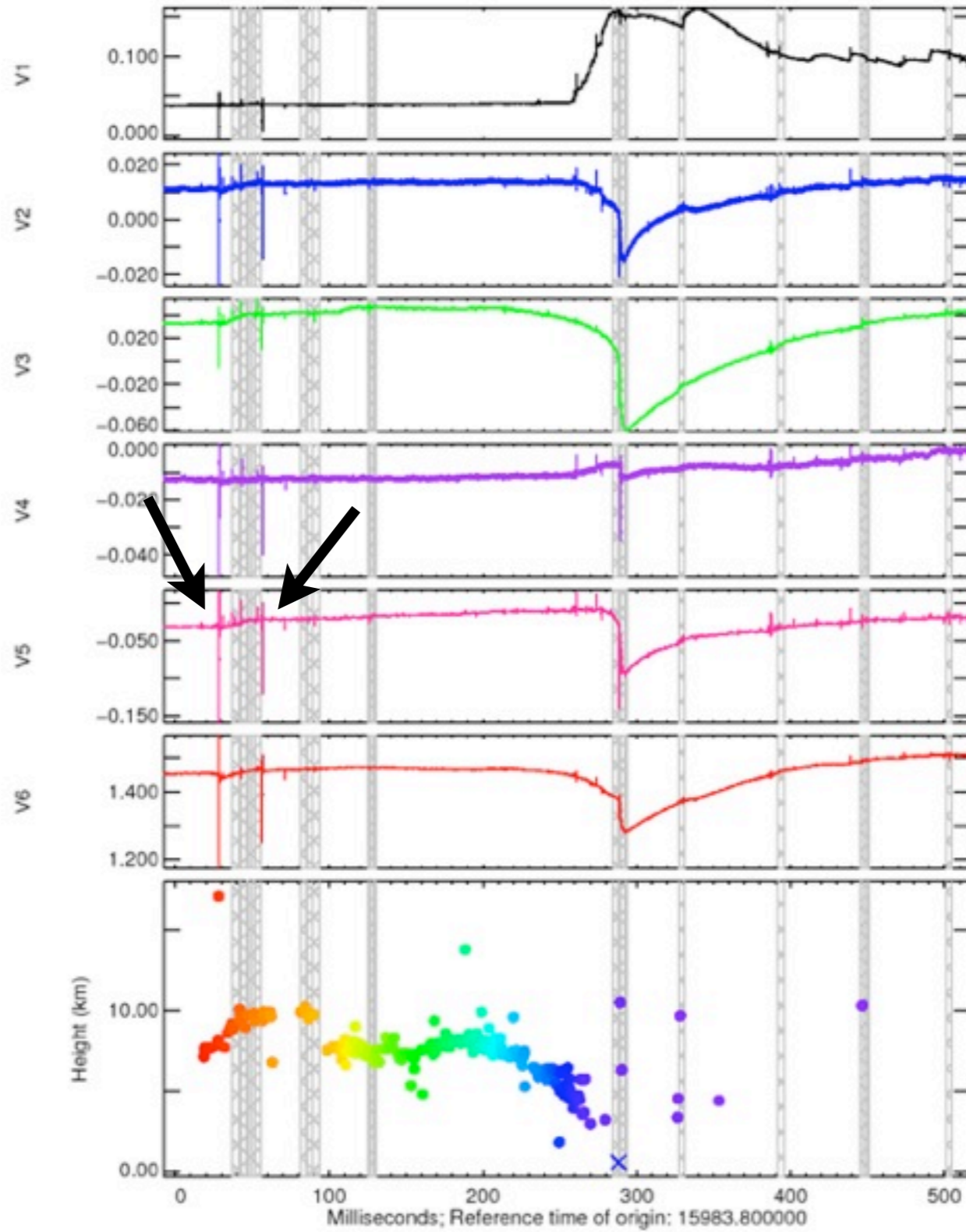
Again, the wideband record is active...



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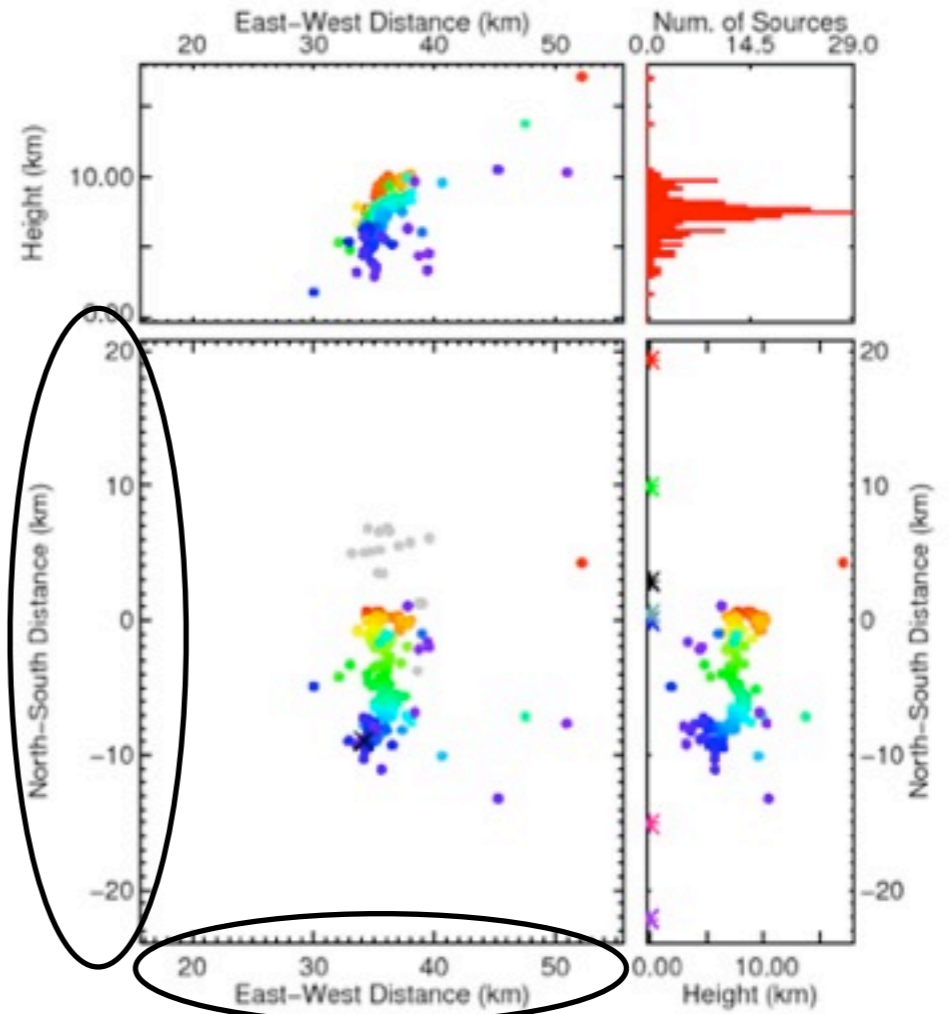
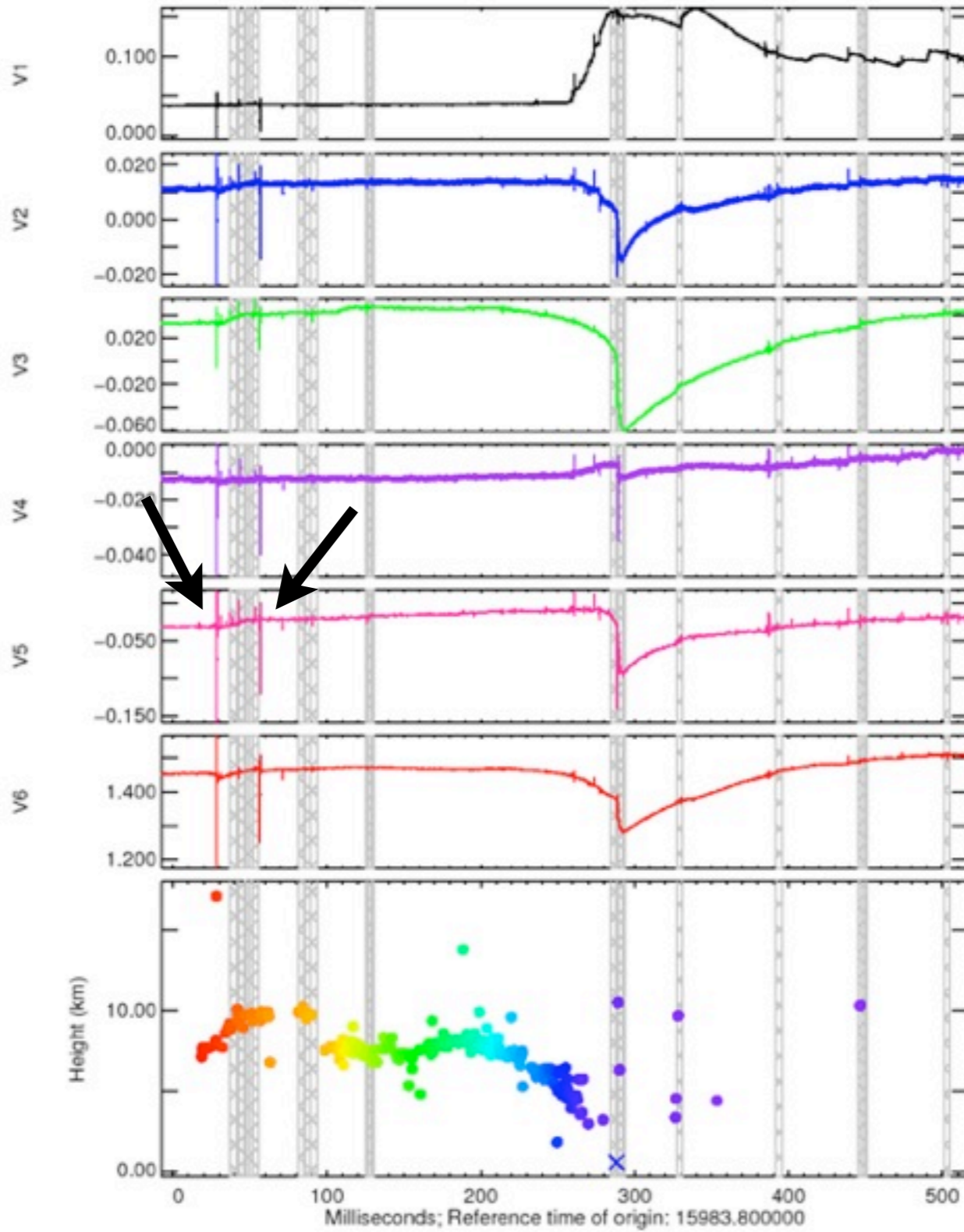
...and there are LIS groups without detected VHF sources.

2010/10/25 04:26:23



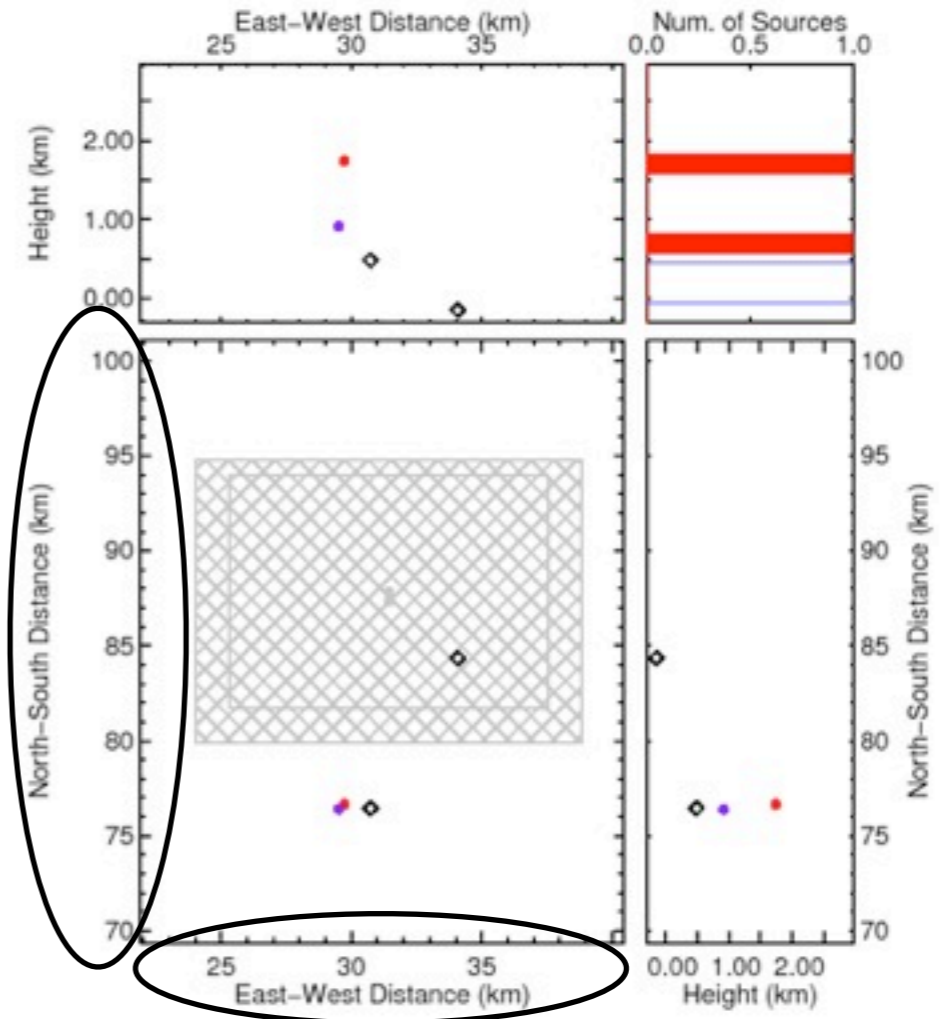
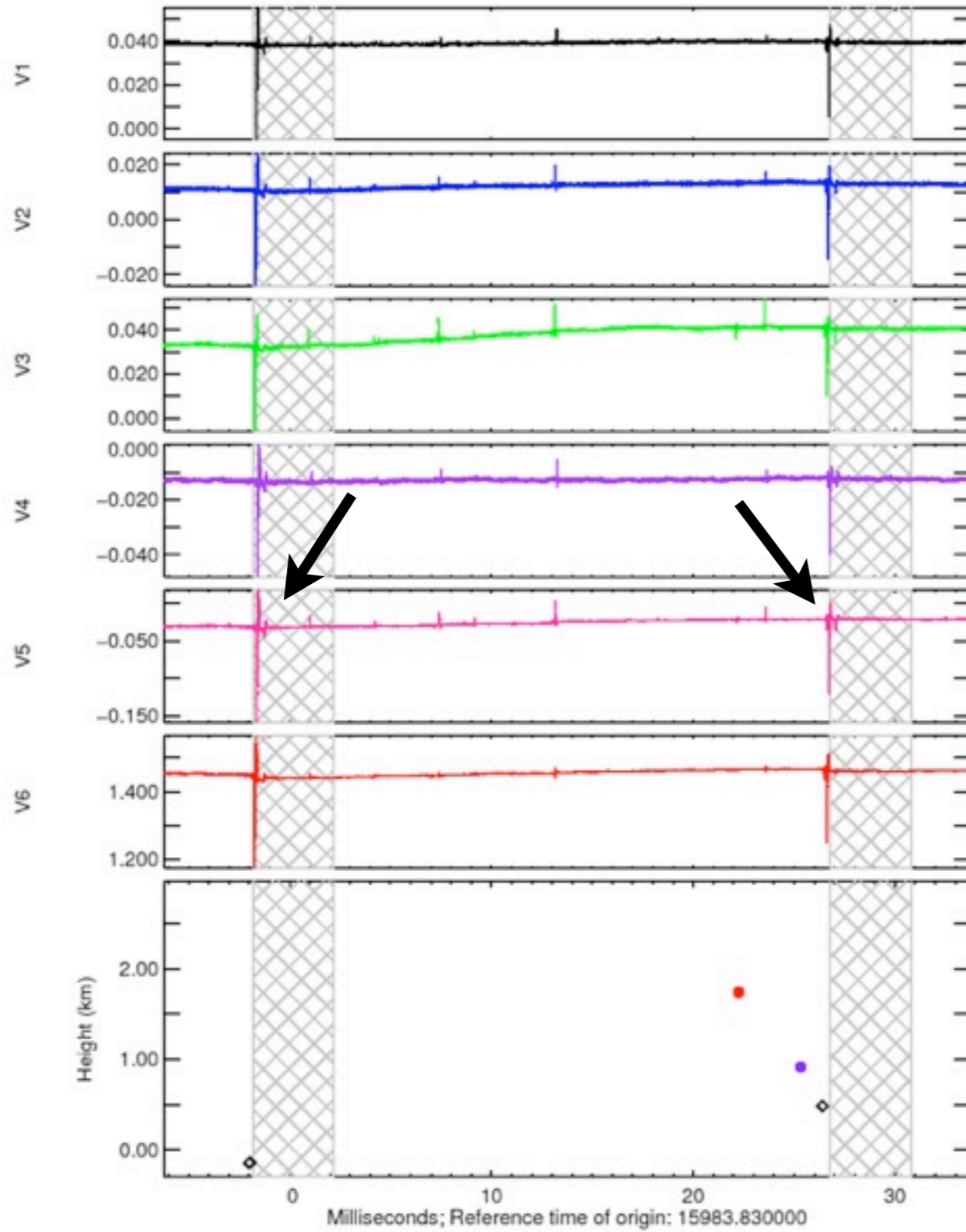
Base time: 15983.792775
Stop time: 15984.316997
Time Elapsed: 0.524222

2010/10/25 04:26:23

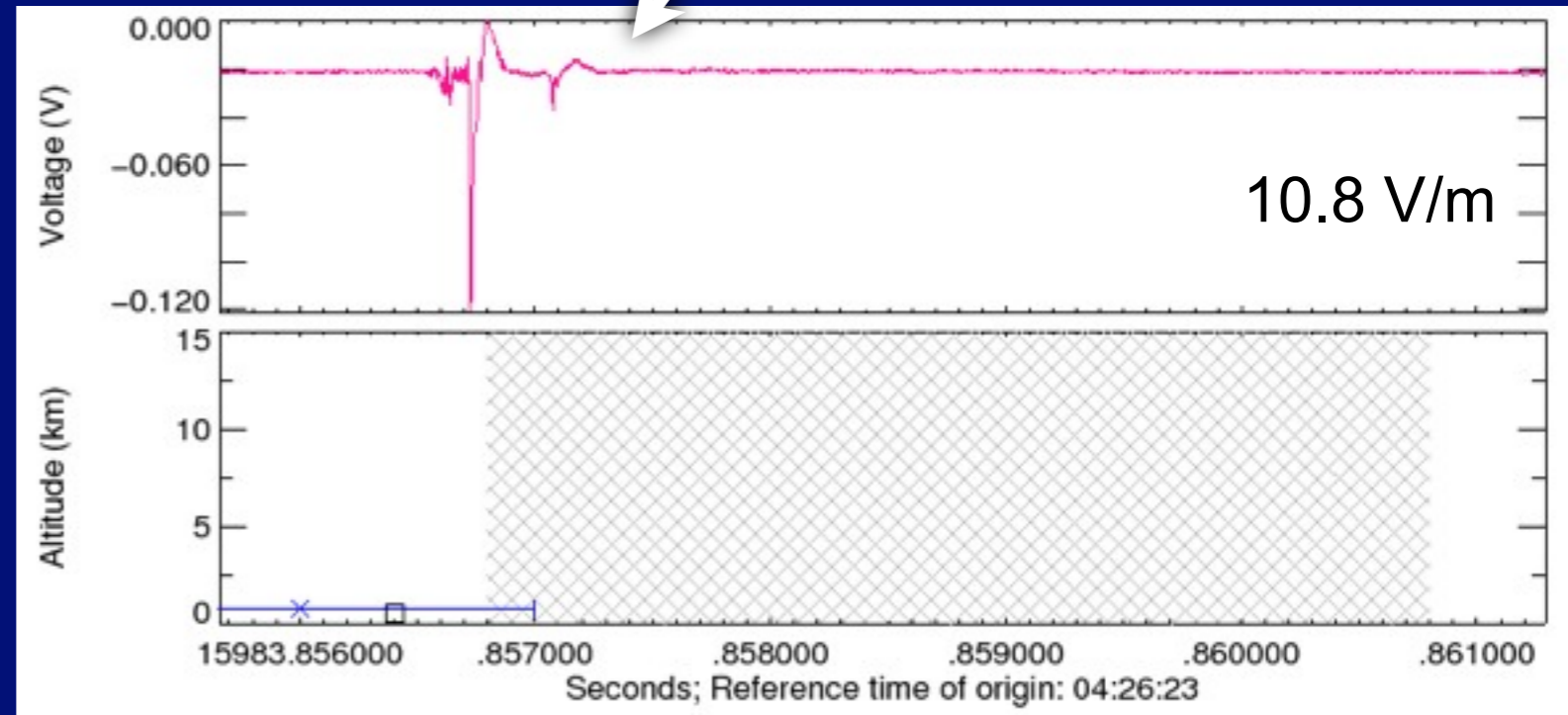
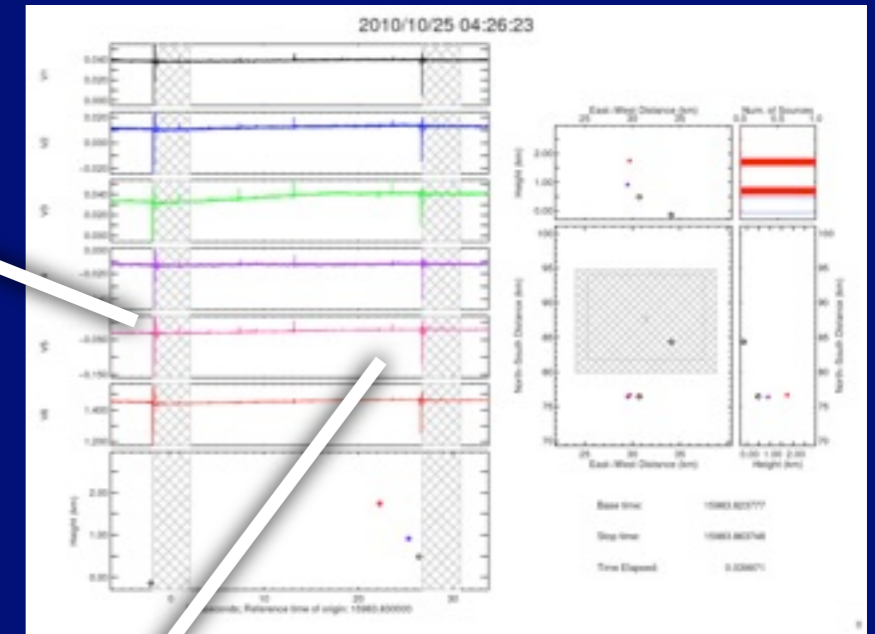
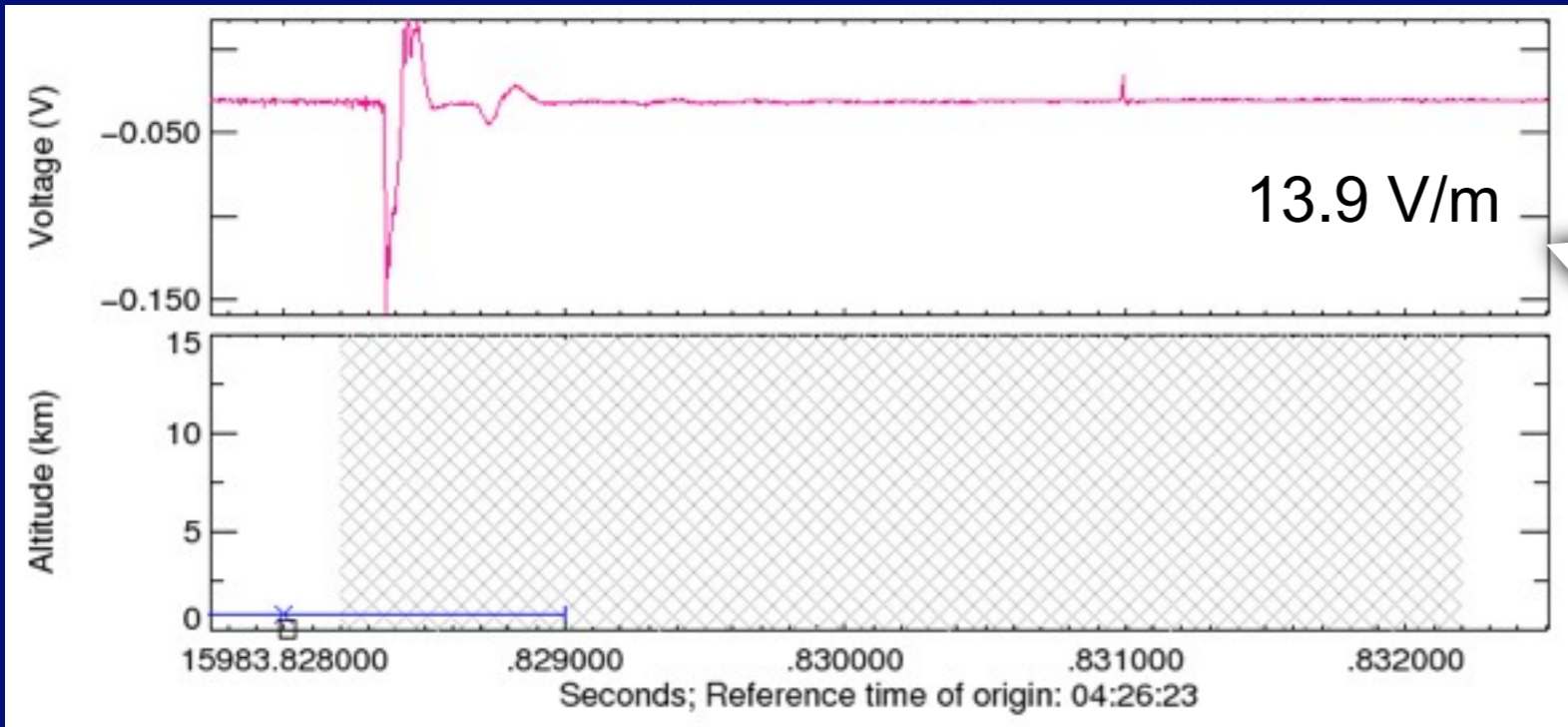


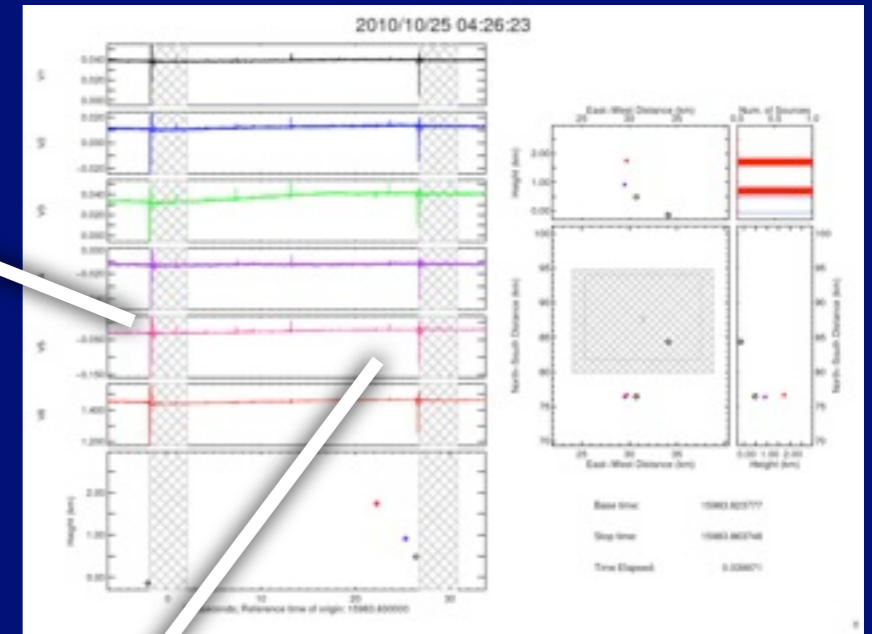
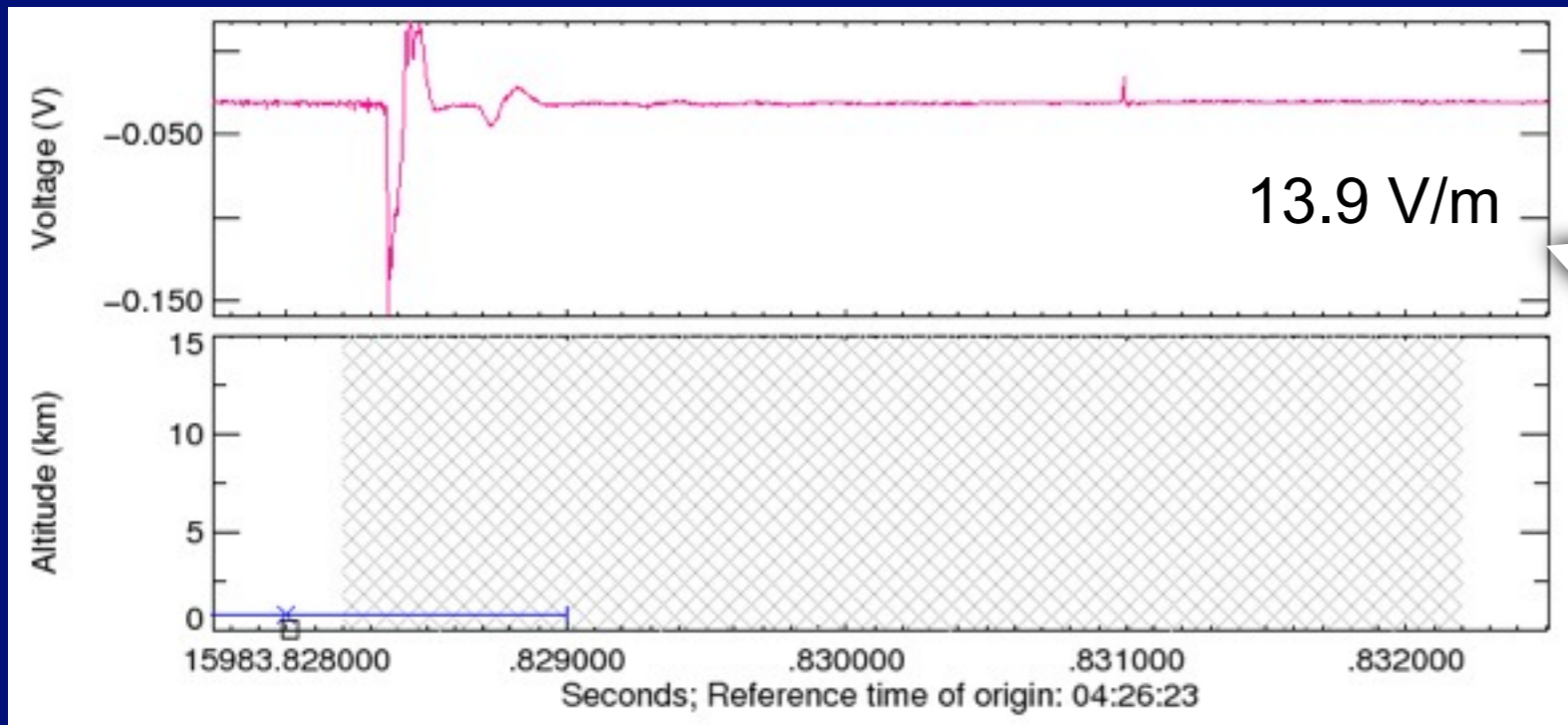
Base time: 15983.792775
Stop time: 15984.316997
Time Elapsed: 0.524222

2010/10/25 04:26:23

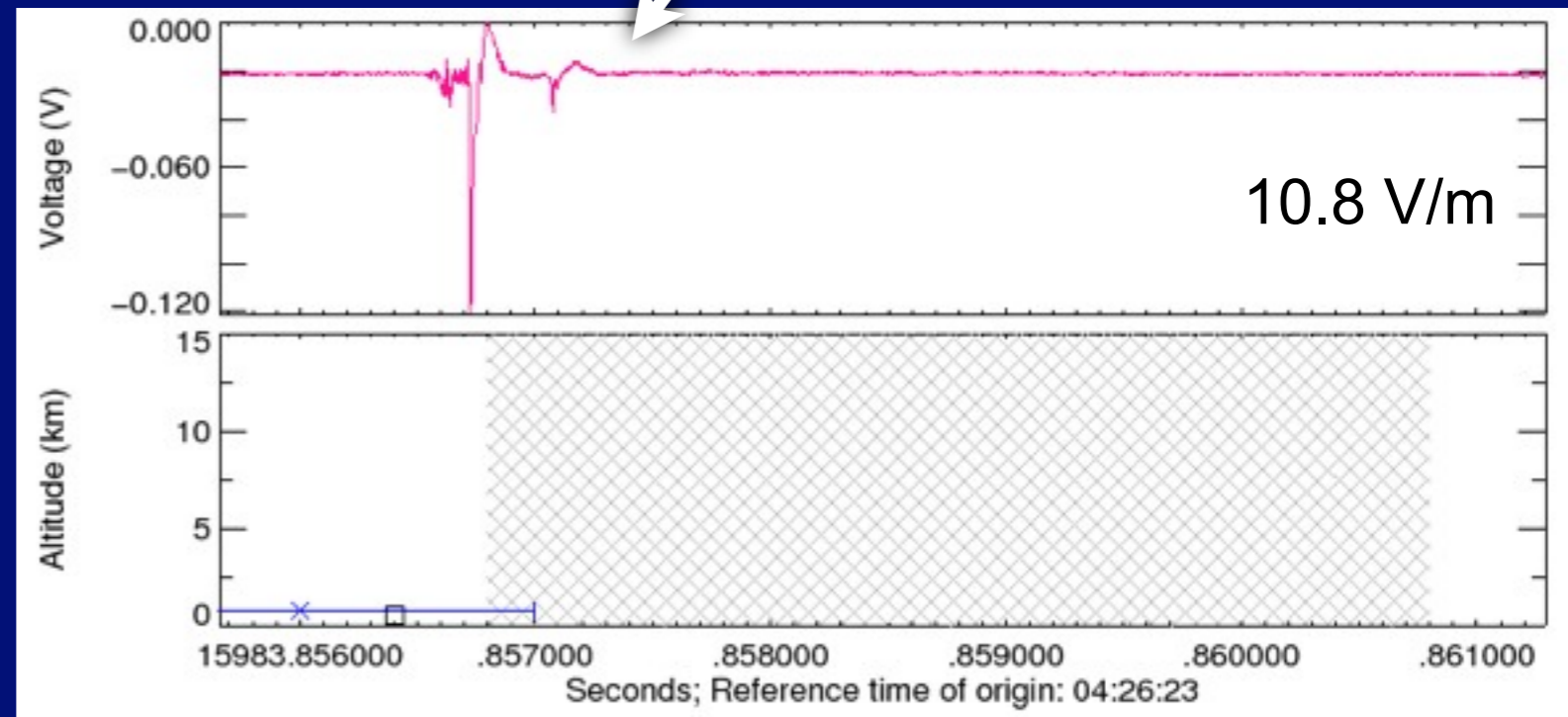


Base time: 15983.823777
Stop time: 15983.863748
Time Elapsed: 0.039971





Without an array of wideband sensors, these might have been misinterpreted as “misses”



*Since LIS/GLM does not detect “flashes,”
we must find a way to validate what
LIS/GLM does detect.*

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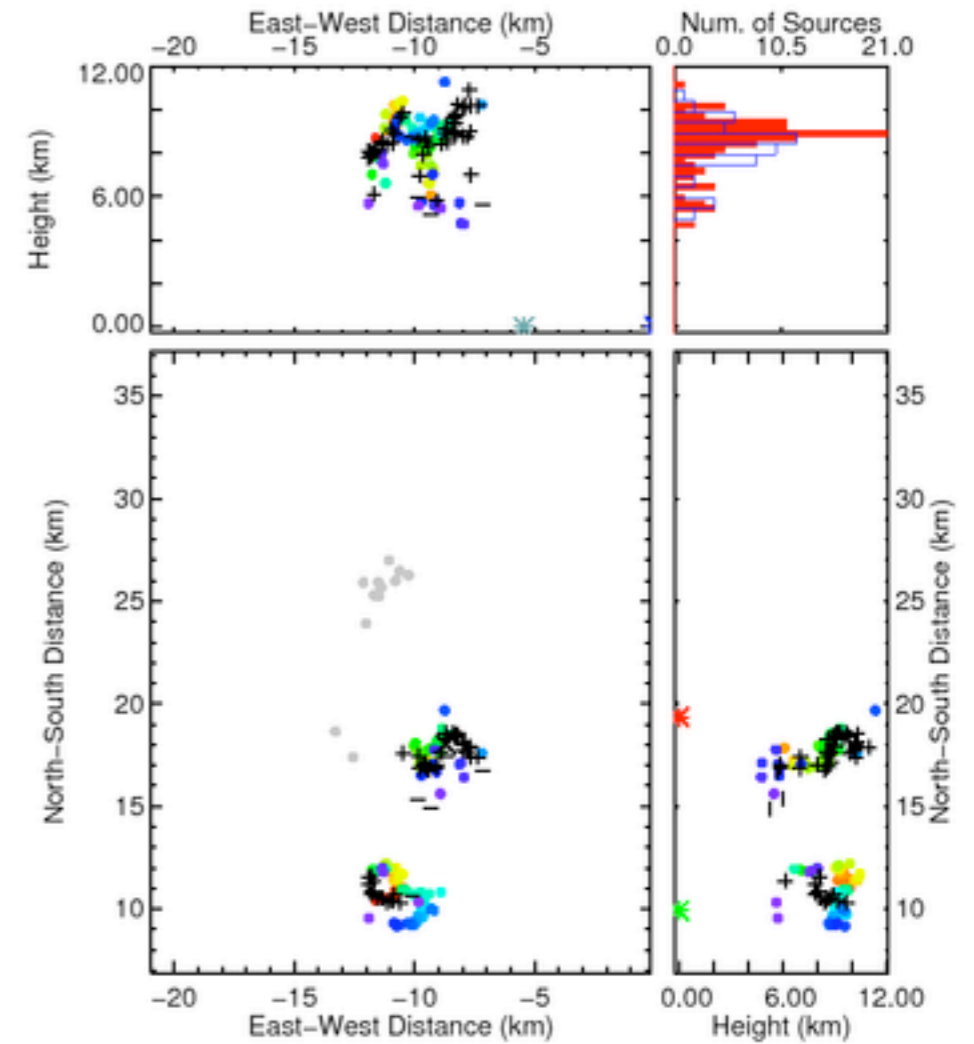
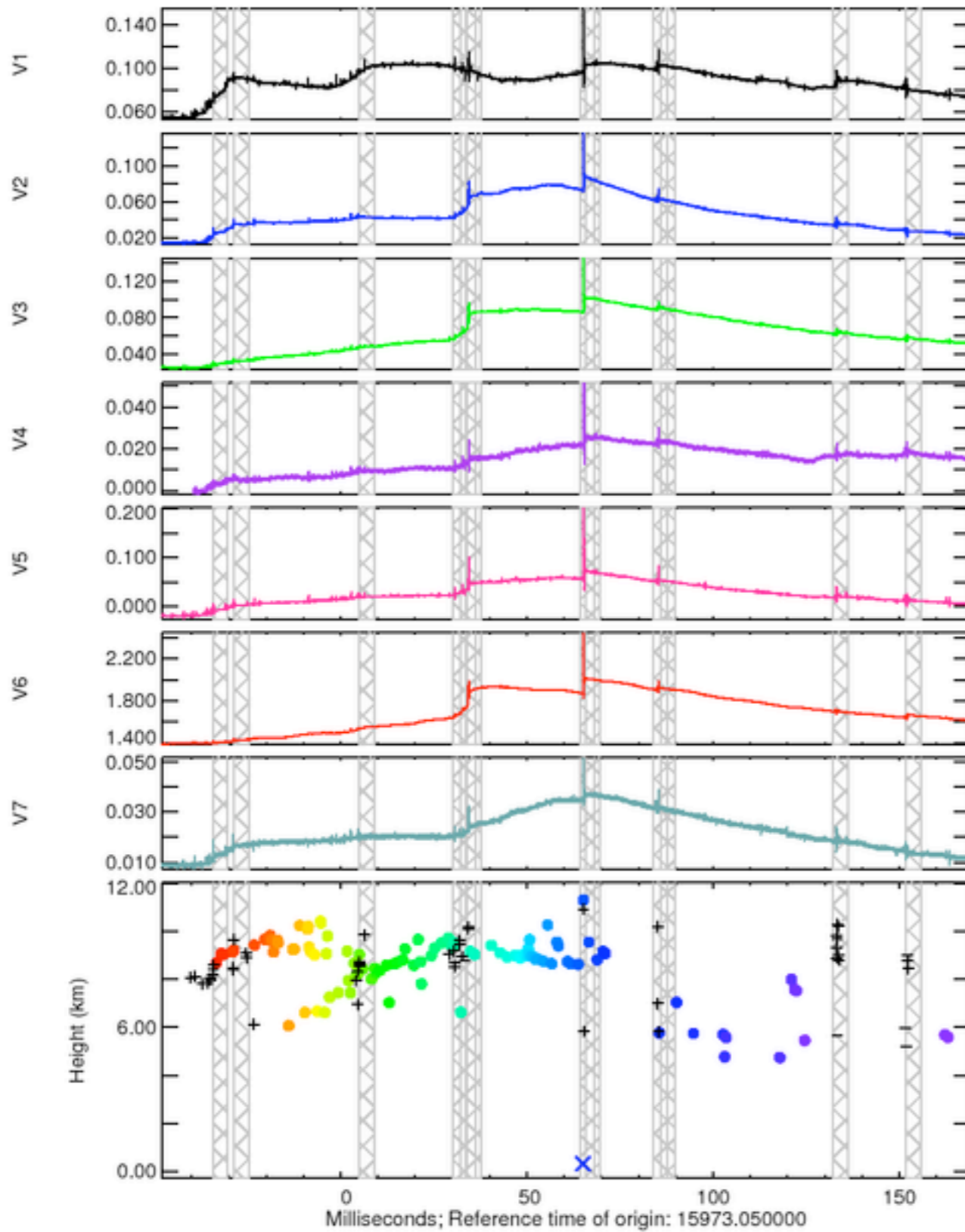
*Clearly, wideband measurements are
better correlated to these optical events...*

*Since LIS/GLM does not detect “flashes,”
we must find a way to validate what
LIS/GLM does detect.*

*Clearly, wideband measurements are
better correlated to these optical events...*

*and even suggest there is a
lower limit to what LIS can detect!*

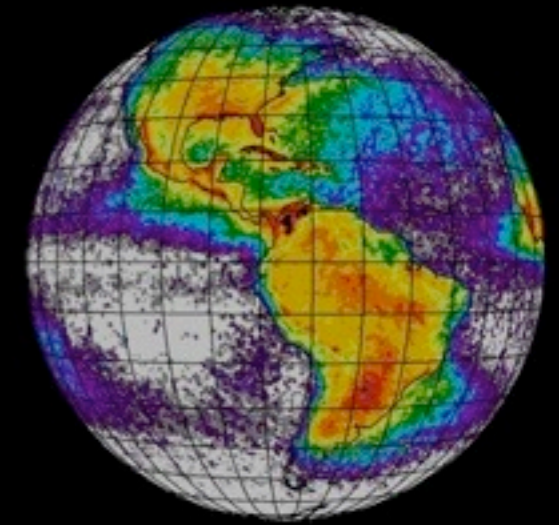
2010/10/25 04:26:13



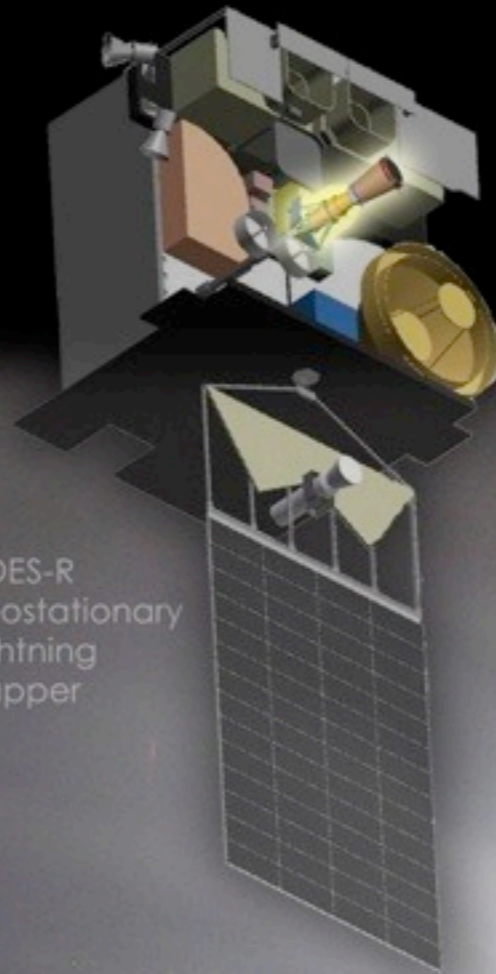
Base time: 15973.002030

Stop time: 15973.219127

Time Elapsed: 0.217097



Fundamental
question for
scientific
validation:



GOES-R
Geostationary
Lightning
Mapper

What is the best
way from the ground
to characterize what
GLM “sees?”

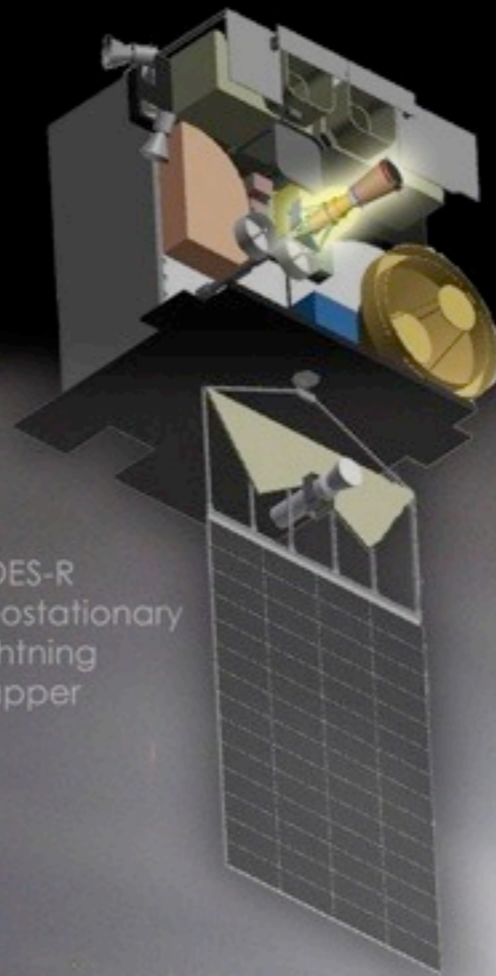
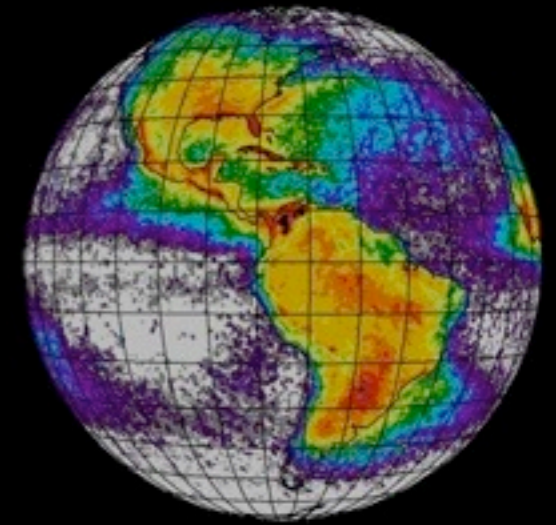


Fundamental question for *scientific* validation:

What is the best way from the ground to characterize what GLM “sees?”

Wideband measurements are ideally suited to the processes in lightning that produce optical emission

UAHuntsville
THE UNIVERSITY OF ALABAMA IN HUNTSVILLE



GOES-R
Geostationary
Lightning
Mapper

