1 Observations of ozone production in a dissipating tropical convective cell dur-2 ing TC4

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17 Abstract. From 13 July – 9 August 2007, the Nittany Atmospheric Trailer and Integrated Vali-

- 18 dation Experiment (NATIVE) team launched 25 ozonesondes as part of the Tropical Composi-
- 19 tion, Cloud, and Climate Coupling (TC4) mission. On 5 August 2007, a strong convective cell
- 20 formed in the Gulf of Panama. World Wide Lightning Location Network (WWLLN) data indi-
- 21 cated 563 flashes between 0900 and 1700 UTC in and around the Gulf. NO₂ data from the
- 22 Ozone Mapping Instrument (OMI) show enhancements near Panama, suggesting lightning pro-
- 23 duction of NO_x. At 1505 UTC an ozonesonde ascended into the southern edge of the now dissi-
- 24 pating convective cell as it moved ashore and west across the Azuero Peninsula. Due to
- 25 condensation and/or down drafts, the balloon ascended from 2.5 to 5.1 km five times between
- 26 1512 and 1700, providing a unique examination of ozone photochemistry on the edge of a con-
- 27 vective cell. Ozone at these altitudes increased 4 12 ppb between the first and last ascent, re-
- sulting cell wide in $\sim 3.3 \times 10^6$ moles of ozone (assuming uniform production), of which only
- 29 ~30% can be explained by descent. This estimate agrees reasonably well with our estimates of
- 30 lightning ozone production from the WWLLN (~ 2.2×10^6 moles), from the radar inferred light-
- 31 ning flash data ($\sim 1.9 \times 10^6$ moles), and from the OMI NO₂ data ($\sim 2.6 \times 10^6$ moles), though all
- 32 estimates have large uncertainties.

33 1. Introduction

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35 The role of lightning in global ozone (O_3) formation has been studied extensively in the latter half of the 20th Century [e.g., Kroening & Nev, 1962; Orville, 1967; Griffing, 1977; Levine, et 36 37 al., 1981] when it was realized that lightning is a significant contributor to the Earth's reactive 38 nitrogen (NO_x) budget [Noxon, 1976; Peyrous & Lepeyre, 1982; Franzblau and Popp, 1989]. 39 With refinements in regional and global models and lightning flash data from satellites and 40 ground-based networks, the global lightning NO_x (LNO_x) budget has been reliably set at 2 - 8 Tg 41 N/year [Pickering et al., 2009; Schuman and Huntrieser, 2007]. Grewe [2007] estimates the 42 global lightning contribution to tropospheric O_3 at > 30% with the fractional source greatest in 43 the tropics. 44 Martin et al. [2007] used SCIAMACHY data to show enhancements of 10 – 15 Dobson units (DU) of O_3 and $(2-6) \times 10^{14}$ molecules/cm² of NO₂ over the tropical Atlantic and Africa result-45 ing from lightning, with the largest enhancements observed downwind of convection. Ozone-46 sonde data taken in the summer of 2006 aboard the Ron Brown off the Atlantic coast of Africa 47 48 show a 7 DU increase in tropospheric column O_3 during June as a result of biomass burning in 49 Central Africa and lightning over west Africa [Jenkins, et al., 2008]. Several sources of O_3 data near convective systems have produced a wide range of results. 50 51 Schlanta and Moore [1972] examined sonde data and found high O₃ aloft (2 – 3 times pre-storm 52 ground level concentrations) associated with thunderstorms. *Clark and Griffing* [1985] observed 53 large spikes in O_3 (> 500 ppb in a background of ~200 ppbv) on their research aircraft downwind 54 of a thunderstorm near Baltimore, MD in August 1980. Dickerson et al. [1987] show PRE-55 STORM project (1985) data with the highest O₃ concentrations near 5 km (see their Table 1) and 56 a local max at 10 km near the anvil of a thunderstorm. Minschwaner et al. [2008] present evi-

-2-

57 dence that coronal discharges associated with convective cells produce $\sim 20\%$ of the total O₃ 58 from a storm (with the remainder from photochemical production due to LNO_x), while Jadhav et 59 al. [1996] suggest direct lightning production of ozone is limited to < 2 - 3% of the total. 60 Salzmann et al. [2008], on the other hand, used data from TOGA COARE/CEPEX to model 61 the photochemistry in regions of deep convection and found O₃ losses, with a maximum loss at 62 ~5 km (see their Figure 10). Aircraft data from the NOAA P3-B taken during the NASA Pacific 63 Exploratory Mission – Tropics B found no evidence for O₃ or CO production by lightning, but found 1.2×10^{22} molecules/m/flash of NO and O₃ decreases of 6 – 8 ppb (26 – 28 ppb in clear air 64 to ~20 ppb in the convective system) [*Ridley et al.*, 2006]. Ott et al. [2007] report modeled 65 66 losses of O₃ for the 21 July 1998 European Lightning Nitrogen Oxides Project storm, with 67 maximum losses of 9 ppb at 5.5 km during the 3 hours of the storm. Downstream of the storm, 68 they found O₃ production of 1.5 ppb/day, with a maximum change of 5 ppb/day at 5.5 km, while 69 *DeCaria et al.* [2005] found a maximum increase in O_3 production at 9 km of 10 ppb/day. For a 70 tropical case sampled by sondes and the DC-8 over Brazil in 1992, air in the 8 – 12 km convec-71 tive outflow layer enriched by 1 ppbv NO from lightning produced 7 - 8 ppbv/day of O₃ [*Picker*-72 ing et al., 1996; Thompson et al., 1997]. 73 Laboratory studies have suggested that direct production of O₃ occurs mainly during the pre-

discharge period of storms [*Peyrous & Lapeyre*, 1982], and *Franzblau* [1991] found (1) little O₃
production except at very low energies, (2) large decreases in O₃ immediately after discharges,

and (3) nearly full recovery to pre-discharge levels after ~10 min. (see his Figure 1).

Several recent studies have used strategically-designed ozonesonde networks [*Thompson*,
2009] to evaluate lightning contributions to the tropospheric O₃ budget. Sonde profiles provide a
consistent framework for observing impacts of lightning from the surface to lower stratosphere

-3-

80 and for relating O₃ variability to meteorological changes day-to-day. During INTEX-A (Inter-81 continental Transport Experiment [Singh et al., 2006]) analyses of ~300 profiles collected July-82 August 2004 over North America (the IONS-04 [INTEX Ozonesonde Network Study] series 83 [Thompson et al., 2007a]) provided several insights. By combining FLEX-PART trajectories 84 [Stohl et al., 1998, 2005] and lightning flash maps with the sonde data filtered to eliminate 85 stratospheric air masses, Cooper et al. [2006] deduced that within the middle to upper tropo-86 sphere (9 - 12 km), up to 80% of the O₃ above background originated from LNO. Thompson et 87 al. [2007b] used O₃ laminae from the IONS-04 soundings with tracers (lightning maps, potential 88 vorticity, and satellite images of CO and absorbing aerosols) to compute a four-term O₃ budget 89 for each tropospheric O_3 profile. From these budgets it was determined that 15 - 20% of the tro-90 pospheric O₃ column over northeastern North America was associated with lightning or pollution 91 introduced into the free troposphere by convection. A limitation of the *Thompson et al.* [2007b] 92 analysis is that lightning contributions older than about a week are not distinguished from other 93 sources of aged, background O_3 . *Pfister et al.* [2008] addressed the issue using tagged NO_x 94 sources in a chemical transport model to conclude that an average of $10 \pm 2\%$ of tropospheric O₃ 95 in the IONS-04 series originated from lightning. 96 The summertime IONS-06 soundings [Thompson et al., 2008] included approximately 300

97 profiles collected August – September 2006, from Mexico City to Canada and from the west to 98 east coasts of the United States (US), providing further insights into the lightning – O_3 connec-

99 tion. In *Thompson et al.* [2008] laminar-based budgets implicated lightning in upper tropo-

100 spheric (UT) O₃ enhancements over Houston and Mexico City, although stratospheric O₃ from

101 the extra-tropics was the dominant player over Houston in the latter third of the campaign. The

-4-

dominance of lightning as a UT O₃ source over the south central US emerged from IONS-06 in
the FLEX-PART analysis by *Cooper et al.* [2007], as it had in IONS-04.

104 The more subtropical of the IONS-06 soundings pointed to the complexity of processes af-105 fecting O₃ profile structure during the early stages of the North American monsoon period, when 106 near-daily convection is ascendant over places like Mexico City. Convective activity excites 107 vertically propagating waves that are detected in laminar analysis [Grant et al., 1998; Loucks, 108 2007; Thompson et al., 2007a,b; 2008]. The TC4 campaign [Toon et al., 2009, this issue] was an 109 excellent vehicle to further investigate wave activity, lightning, and O₃ responses in a highly 110 convective environment, closer to the Intertropical Convergence Zone than the August - Sep-111 tember IONS-06 soundings. The concentration of new satellite products and aircraft instrumen-112 tation focused on convection were intended to quantify lightning, to relate flashes to improved 113 representations in models, and to link the latter to validation of NO₂ from the Ozone Measuring 114 Instrument (OMI) aboard the NASA Aura satellite, for example [Bucsela et al., 2009, this issue]. 115 In addition to O_3 measurements on the three TC4 aircraft, O_3 profiles were provided through 116 soundings at the Southern Hemisphere Additional OZonedondes (SHADOZ) [Thompson et al., 2003] Costa Rican station (Heredia, 10^oN, 84^oW) and at the TC4 ground site at Las Tablas, Pan-117 ama (7.75[°]N, 80.25[°]W; see also *Thompson et al.* [2009, this issue]). Free tropospheric (FT) and 118 119 lower stratospheric wave signatures were identified in virtually all the Costa Rican and Panama 120 soundings, with an incidence in the tropical troppause layer (TTL) of > 40% (see Figure 2a in 121 Thompson et al. [2009, this issue]). Case studies of O₃ within segments affected by gravity 122 waves demonstrated a clear link to convective activity, as viewed through aircraft tracers, cloud 123 lidar and radar, satellite NO₂, and lightning network imagery.

-5-

124 The 25 Panama sondes, in a region with a high convective frequency during TC4 [Toon et 125 al., 2009, this issue], displayed prominent wave activity associated with convection near the be-126 ginning of the mission (13 - 22 July 2007) and after 2 August 2007, when the TC4 aircraft coor-127 dinated sampling south of Costa Rica, in the vicinity of the Panama Bight, and as far south as the 128 Galapagos. During the second convective period on 5 August, a day when all three TC4 aircraft 129 flew over the Panama Bight, the ozonesonde launched from Las Tablas was caught in a convec-130 tive system that kept it oscillating between 2.5 and 5.1 km for nearly two hours before it resumed 131 normal ascent. Though many previous studies have provided profiles before and after convec-132 tion, this ozonesonde data set is unique in providing insights into changing O_3 concentrations 133 inside a dissipating tropical convective cell. Ozone increased throughout the oscillatory period, 134 and we trace the cause to lightning photochemistry. The following sections provide background 135 with experimental details (Section 2), observations (Section 3), and a summary (Section 4).

136 **2. Background**

137 2.1 NPOL Radar

The NASA polarimetric Doppler weather radar (NPOL) is a S-Band system operating at a frequency of 2.8 GHz (10 cm wavelength). It has horizontal and vertical beam widths of 1.4°. The antenna is a flat passive array instead of the typical parabolic dish. The design allows the system to be transported easily and allows NPOL to operate in conditions with strong winds (e.g., minimum wind loading). One of the main drawbacks of this design is the signal deteriorates when the antenna is wet [*Theisen et al.*, 2009]. Therefore, periods when there is precipitation at the radar site are excluded from the final quality controlled dataset.

For TC4, NPOL was deployed with NATIVE near Las Tablas. NPOL operated almost continuously 16 July – 12 August 2007, with the exception of the period 1800 UTC 19 July – 0200

-6-

147 UTC 21 July. NPOL scanned with a temporal resolution of 10 min and spatial resolution of 200 148 m. A 12 tilt scanning strategy was used with elevation angles ranging from 0.7° to 23.3°. A 149 volume scan with a maximum range of 150 km was followed by a long-range (to 275 km) sur-150 veillance scan. NPOL measured or derived quantities included standard radar parameters: radar 151 reflectivity (DZ), Doppler velocity (VR), and spectral width (SW); and polarimetric parameters: 152 differential reflectivity (ZDR), differential phase (PhiDP), specific differential phase (KDP), and 153 cross correlation (corrHV) between horizontal and vertical polarizations. Nearly 3500 volume 154 scans are available for studying the convective properties in and around Panama.

A variety of events were observed during TC4, ranging from short-lived unorganized convection to long-lived mesoscale convective systems. Often systems developed over the Gulf of Panama in the late evening, but weakened or dissipated before reaching land in the mid-morning hours. As a result of strong diurnal heating, however, a second maximum in convection was often observed over the mountainous regions of Panama during midday. A more detailed discussion of the convection observed during TC4 can be found in *Kucera and Newman* [2009, this issue].

162 **2.2 NATIVE**

163 Continuous surface O₃ measurements at the Las Tablas site were made on the Nittany At-164 mospheric Trailer and Integrated Validation Experiment (NATIVE) during the period 17 July – 8 165 August 2007 with a Thermo Electron Corporation (TECO) Model 49C Ozone Analyzer using the 166 United States Environmental Protection Agency standard measurement technique (EQOA-0880-167 047). Carbon monoxide (TECO 48C-TL) and SO₂ (TECO 43C-TLE) were measured at the same 168 time, along with the aerosol size distribution (SMPS). After 29 July, NO and NO₂ (TECO 42CY)

-7-

were also measured continuously. All measurements from the TC4 mission can be found on theweb at http://ozone.met.psu.edu/NATIVE/TC4.html.

171 **2.3 World Wide Lightning Location Network**

172 In 2007 the World-Wide Lightning Location Network (WWLLN) [Rodger et al., 2006] con-173 sisted of approximately 25 sensors detecting lightning flashes at VLF frequencies of 3 – 30 kHz. 174 Flash data (primarily cloud-to-ground or CG flashes) were obtained in near-real-time by NASA 175 GSFC from WWLLN Director, Robert Holzworth, of the University of Washington. Bucsela et 176 al. [2009, this issue] estimated the detection efficiency of the WWLLN network of detectors for 177 total flashes (CG + intracloud or IC flashes) in the TC4 region (over open ocean near Costa Rica 178 and Panama) through comparisons of flash rates from the Costa Rica Lightning Detection Net-179 work (CRLDN, which uses the same sensors as the United States National Lightning Detection 180 Network, NLDN [Cummins et al., 1998]) and the Lightning Imaging Sensor (LIS) on the Tropi-181 cal Rainfall Measuring Mission (TRMM) satellite for six storms. The mean detection efficiency 182 was 0.22 ± 0.08 , in reasonable agreement with, though somewhat higher than the estimate of 183 *Roger et al.* [2006]. There is some indication, however, that the detection efficiency is greater 184 over ocean than over land in this part of the world [Lay et al., 2009].

185 **2.4 OMI**

OMI has been collecting data since shortly after its launch in July 2004 [*Levelt, et al.*, 2006]. The instrument is a nadir-viewing spectrometer with a CCD array having wavelength and spatial dimensions, the latter comprising 60 pixels across the flight track. The pixel area at nadir is 13 x 24 km², although this value increases by approximately an order of magnitude near the edges of the track. Overpass time is ~1345 local time in the tropics, improving the capability of OMI for

-8-

observing afternoon convective events as compared with GOME and SCIAMACHY (whichhave morning overpass times).

193 The retrieval algorithm for NO₂ from OMI has been described by *Bucsela et al.* [2006, 2008], 194 Celarier et al. [2008], and Wenig et al. [2008]. It employs a spectral fitting procedure to obtain 195 NO₂ slant column densities (SCDs) from the OMI spectra. Vertical column densities (VCDs) are 196 obtained by dividing SCDs by air mass factors (AMFs), which are derived through radiative 197 transfer calculations. The tropospheric component of the vertical column, resulting from sources 198 that include pollution and lightning, is obtained by removing an unpolluted (here simply called 199 "stratospheric") component, using a wave-2 analysis in narrow latitude bands. Because of the 200 method used to derive it, the stratospheric VCD is contaminated by small amounts of tropo-201 spheric NO₂, which we remove in the present study using output from the GMI model [Duncan 202 et al., 2007]. The corrected stratospheric VCD differs from the uncorrected value by approxi-203 mately 5%.

204 For the 5 August 2007 analysis, OMI NO₂ data are used to estimate the moles of LNO_x in the 205 region near the Gulf of Panama. Details of the procedure are found in Bucsela et al. [2009, this 206 issue], but we outline the method here. The first step is the calculation of the tropospheric SCD 207 due to LNO₂, which is given by the total SCD minus the sum of the corrected stratospheric SCD 208 and the tropospheric SCD due to sources other than lightning. This non-lightning tropospheric 209 SCD is obtained by running the GMI model with the lightning source turned off. In these calcu-210 lations, all SCDs and VCDs are related through AMFs derived using radiative transfer calcula-211 tions and climatological NO₂ profiles. The LNO₂ slant column is converted to a vertical column 212 of LNOx using a modified AMF that accounts for the vertical distribution of LNO₂ and the pho-

-9-

213	tolyis ratio of $[NO_2]/[NO_x]$. The former is obtained from TC4 aircraft data, and the latter from
214	model calculations. This approach is similar to that described by Beirle et al. [2009].
215 216 217	2.5 Ozonesondes
218	Ozone profiles during TC4 at the Las Tablas site were measured using the electrochemical
219	concentration cell (ECC) type [Komhyr, 1986; Komhyr et al., 1995] En-Sci 2Z ozonesonde in-
220	struments with 0.5% buffered KI cathode solution. The Jülich Ozone Sonde Intercomparison
221	Experiment (JOSIE) found biases $< 5\%$, a precision of 3–5%, and an accuracy of 5–10% up to
222	30 km for such sondes [Smit, et al., 2007]. With a typical rise rate of ~5 m/s and a measurement
223	time constant of ~25 s, the effective vertical resolution of O_3 features is ~125 m [see also <i>Smit et</i>
224	al., 2007]. Most launches occurred from 12 – 3 pm local time (1700 – 2000 UTC) to coincide
225	with the ~1345 local solar time overpass of NASA's Aura satellite [Schoeberl et al., 2006].
226	Pressure, temperature, and relative humidity (RH) measurements were measured by Vaisala
227	RS80-15N radiosondes on each payload, as described in Thompson et al. [2003, 2007a]. Pay-
228	loads also contained global positioning systems (GPS) that provided latitude, longitude, altitude,
229	wind speed, and wind direction data. Pressure readings are validated through comparisons of
230	pressure altitude with GPS altitude. When pressure offsets are observed, they are usually <2 hPa,
231	meaning that tropospheric O_3 mixing ratios are adjusted <~2% (<~0.2% at the surface). Correc-
232	tions to pressure errors are made through post-flight processing so that at burst altitude, pressure
233	and GPS altitudes agree to within 200 m. For 7 of the launches during TC4, RH data above 300
234	– 500 hPa appear unreliable. All the ozonesonde data can be found at
235	http://physics.valpo.edu/ozone/tc4data.html.
236	Although each ozonesonde is internally calibrated before each flight, Figure 1 shows a com-

237 parison of NATIVE surface O₃ readings with pre-launch ozonesonde readings from the 23 good

flights during TC4. The mean bias (sonde – NATIVE) was -0.4 ± 1.2 ppbv, with a root mean square difference of 1.05 ± 0.76 ppbv.

240 As further validation, ozonesonde columns are compared with OMI total column ozone 241 [Bhartia, 2007; McPeters et al., 2008] for nearby overpasses (< 50 km from launch site). 242 Ozonesonde profiles are integrated to burst altitude, then augmented with either a constant mix-243 ing ratio assumption for the upper stratosphere or the Solar Backscatter Ultra-Violet (SBUV) 244 balloon burst climatology of *McPeters et al.* [1997]. For the former case, the difference is $17.5 \pm$ 245 3.8 DU, while for the latter, the difference is 16.7 ± 6.2 DU, with the sondes higher (by ~6%) 246 than OMI. These results are not inconsistent with the finding that the Paramaribo SHADOZ station (5.8°N, 55.2°W) reports columns ~10% higher than OMI [*Thompson et al.*, 2007c]. 247

248 **3. Observations**

249 **3.1 NPOL**

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251 NPOL observed a large system off the coast of Panama on 5 August 2007. The system de-252 veloped during overnight hours in the Gulf of Panama and slowly propagated westward toward 253 the Azuero Peninsula. The precipitating system covered an area that was several hundred kilo-254 meters in the north-south direction and on the order of 100 km in the west-east direction, with the convective core having a mean area of $5,300 \pm 2,400$ km² between 0900 – 1700 UTC, as indi-255 256 cated by the radar data. The peak in convection occurred around 1311 UTC. Figure 2 shows the 257 reflectivity (DZ) field near the peak in convection at 1255 - 1305 UTC. Reflectivity values 258 ranged from about 0 dBZ in the lighter precipitation areas to a maximum of about 55 dBZ in the 259 embedded convection.

Estimated vertical velocities are derived from NPOL Doppler velocity field through a technique called volume velocity processing (VVP) [*Boccippio*, 1995]. For the 5 August cell, the

-11-

derived wind in the lower atmosphere was generally from the east at speeds on the order of 10 m s^{-1} . With each time step during which the winds were calculated, however, moderate directional shear appeared, with directions fluctuating from southeast to northeast (in agreement with the ozonesonde observations, see Section 3.5 below).

This convective system generated a significant number of lightning strikes. Figure 2 indicates the location of flashes between 1255 and 1305 UTC as indicated by the WWLLN (see Section 3.3 below), which are well correlated with areas active convection indicated by the NPOL radar data. Here we use cloud top height data from NPOL to estimate lightning flash rates, as suggested by *Price and Rind* [1997]. The parameterization of *Futyan and DelGenio* [2007] predicts the flash rate *F* (flashes/min/300 km²) to be

$$F = 0.208(H_{17dB} - H_{0^0C})^{1.8}$$

where H_{17dB} is the storm averaged height (km) of the 17dB radar return signal and $H_0^{o}{}_{C}$ is the height of the 0⁰C (freezing) level. From the ozonesonde data, we find $H_0^{o}{}_{C}$ is ~4.5 km. From the radar data, we determine hourly averages of H_{17dB} and the area of the storm. The resulting hourly flash estimates for 5 August are shown in Table 1. During the period from the genesis of the cell to the final ascent of the balloon (0900 – 1700 UTC), this parameterization predicts a total of 2300 ± 300 flashes.

We can use the flash data to estimate lightning production of O_3 (LO₃) from LNO. Esti-

280 mates of LNO vary widely [*Pickering et al.*, 2009; *Huntrieser et al.*, 2008; *Koike et al.*, 2007;

281 Hudman et al. 2007; Drapcho et al. 1983] from a low of 43 moles/flash [Skamarock et al., 2003]

to a high of 1100 moles/flash [Price et al., 1997, Winterath et al., 1999]. Pierce [1970] and

283 Prentice and Mackerras [1977] estimate the ratio of intra-cloud (IC) flashes to cloud-to-ground

284 (CG) flashes for $\sim 8^{0}$ N to be 6.5 – 8.0. Estimates of the NO production efficiency of IC to CG

-12-

Morris et al. – Ozone production in a convective cell Version FINAL – 31 Aug. 2009

285 flashes varies from 0.1 [Price et al., 1997] to 1.4 [Fehr et al., 2004]. Lin et al. [1988] estimate 286 30 moles of O_3 are produced per mole of NO_x , (i.e., an ozone production efficiency, OPE, of 30) 287 while more recent studies have found OPE in the range 4 - 12 [Shon et al., 2008; Wood et al., 288 2009; Zaveri et al., 2003]. (Note: For our calculations of LO₃, we scale the OPE to account for 289 the limited 2-hour period between the first and final ascent and assume that enough sunlight was 290 available on the edge of the cell to drive the photochemistry).

Using the ranges above, we estimate $1.3 \times 10^4 - 2.7 \times 10^7$ moles of O₃ could have been pro-291

292 duced by this cell. The wide range of estimates owes to the remaining high uncertainty in all of

293 the quantities that go into the calculation. Recalculating with values that we feel are most repre-

294 sentative of the conditions for this cell (tropical, marine, etc.: OPE = 2; moles NO/flash = 430 as

295 suggested by *Bucsela et al.* [2009, this issue]; IC:CG NO production efficiency = 1), we deter-

mine our best estimate of LO₃ to be 1.9×10^6 moles. 296

297 If we assume direct production of O_3 from the lightning strikes of 300 - 3000 moles/flash, as

in *Minschwaner et al.* [2008], this translates into $4.0 \times 10^5 - 2.0 \times 10^7$ moles of LO₃ for the 298

299 flashes between 0900 – 1700 UTC. If we restrict our calculation to the flashes during two-hour

period after launch (see Table 1), we find a range of $1.4 \times 10^5 - 7.2 \times 10^6$ moles of LO₃, with our 300 best estimate being 9.2×10^5 moles. 301

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303 3.2 NATIVE - eliminate all but the ozone data

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Figure 3 shows hourly mean O_3 data at NATIVE for the 13 July – 9 August period, with the 305 306 mean over the mission as the black dots and the mean \pm one standard deviation as the dashed 307 lines. The time is UTC, with Panama local time 5 hours behind. A diurnal cycle shows a daily minimum around 1230 (near dawn) of around 13 ppb with a daily maximum around 1830 (early 308

afternoon) of ~22 ppb. The hourly mean O_3 data from 5 August are shown by the gray dots.

310 Overnight (0030 – 1230 UTC), O₃ values are on the low end of the typical range, hovering

around 10 ppb. As the cell moved ashore and rain fell, O_3 decreased, with a minimum of < 5 ppb

around 1430. Ozone recovered to more typical values of 20 - 25 ppb by 1930 as the storm

moved inland.

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- 315 316

3.3 WWLLN and Estimated Ozone Production

317 WWLLN reported frequent lightning in association with the convective cell over the Gulf of 318 Panama on 5 August. Table 1 summarizes the number of flashes per hour detected in the box defined by the latitude range $7.25^{\circ} - 8.75^{\circ}$ N and longitude range $78.75^{\circ} - 81.25^{\circ}$ W. Figure 2 319 320 shows the good correlation between the locations of the lightning flashes and the areas of active 321 convective as seen by the NPOL radar for 1255 - 1305 UTC. Notably (but not shown here), the 322 CRLDN observed few if any lightning flashes over the Gulf of Panama on this day. Given the 323 spatial distribution of lightning flashes observed by the CRLDN, it appears that the Gulf of Pan-324 ama fell in a shadow of the network.

325 Using the WWLLN flash data between 0900 – 1700 UTC, we can estimate the associated to-

tal LO₃, as we did for the NPOL data above with one further modification. A range of detection

327 efficiencies have been reported for lightning detection networks, with *Boccippio et al.* [2001]

reporting a 0.9 efficiency for the NLDN and *Bucsela et al.* [2009, this issue] estimating $0.22 \pm$

329 0.08 for the WWLLN. Accounting for the efficiency of the WWLLN, we find 2560 ± 930

flashes between 0900 – 1700 UTC, in reasonable agreement with the NPOL estimate of $2300 \pm$

331 300 flashes.

Combining the factors in Section 3.1 above with the WWLLN flash estimate, we find a range

333 of $1.2 \times 10^4 - 4.5 \times 10^7$ moles of LO₃, with our best estimate of 2.2×10^6 moles (using values

-14-

associated with conditions more likely to be found in the present case, as we did for the NPOL calculation above), consistent with the NPOL estimate. If we assume direct production of O_3 from the lightning strikes of 300 - 3000 moles/flash, as in *Minschwaner et al.* [2008], this translates into $4.7 \times 10^5 - 1.9 \times 10^7$ moles of LO₃ for the flashes between 0900 – 1700 UTC. Unlike the NPOL data, the WWLLN suggests little to no lightning during the period of the sonde oscillation (1500 – 1700 UTC), so this mechanism is not indicated by the WWLLN data for post-launch LO₃.

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344 It is difficult to discern the LNO_2 signal from an examination of Level 2 OMI tropospheric NO_2 345 and cloud fraction data products that result from the standard retrieval conducted at NASA God-346 dard Space Flight Center [Bucsela et al., 2006], in part because the Level 2 data have not been 347 cloud screened. With reprocessing that includes removing an estimate of background NO_2 and 348 applying an air mass factor more appropriate for convective outflow [Bucsela et al., 2009, this 349 issue], the LNO₂ becomes more evident. Figure 4 shows a map of the LNO_x field near the Gulf 350 of Panama on 5 August 2007 after reprocessing (considering the NO_x to NO₂ ratio at cloud-351 outflow levels). The boxed area (\sim 54,000 km²) contains 1520 +/- 1300 kmol LNO_x, which would result in $(1.5 - 11.4) \times 10^6$ moles of O₃ (depending on the OPE selected). If we scale this 352 353 estimate to the size of the core of the convective cell observed on the NPOL radar (average area of ~5,300 km² from 0900 – 1700 UTC), the estimated LO₃ becomes $(0.15 - 1.1) \times 10^6$ moles of 354 O₃, with a best estimate of $(3.0 \pm 2.6) \times 10^5$ moles, whereas if we scale it to match the area of 355 flashes detected by the WWLLN (~46,000 km²), we find a range of $(1.3 - 9.7) \times 10^6$ moles of O₃ 356 with a best estimate of $(2.6 \pm 2.2) \times 10^6$ moles. While the uncertainties are large, the values ap-357

pear to be above background. Furthermore, scaling by the larger area of the WWLLN estimate
 results in the best agreement with the NPOL and WWLLN LO₃ estimates detailed above.

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3.5 Ozonesonde Profiles

363 The ozonesonde launch on 5 August occurred at 1505 UTC to coincide with the scheduled 364 arrival of the NASA aircraft in the Panama area. At the time of launch, rain was falling as part 365 of the convective cell that had just moved ashore from the east, although no lightning was visible. The surface temperature was $\sim 24^{\circ}$ C with RH of 96% and a surface pressure of ~ 1003 hPa. 366 367 About 20 minutes into the flight, the balloon reached 5.1 km and began to descend. About 368 15 minutes after that, it began to ascend again. Between launch and ~1700 UTC, the balloon os-369 cillated up and down through the air mass between ~2.5 and ~5.1 km five times, as shown in 370 Figure 5. (Note: Each ascent is color coded so that subsequent figures can be analyzed more 371 easily.) Although the detailed explanation for this behavior is beyond the scope of this paper 372 (see *Morris et al.*, to be submitted to the *American Journal of Physics*, 2009), it appears to be a 373 combination of downdrafts on the southern side of the westward moving convective cell (based 374 on NPOL radar data) and increased mass due to repeated condensation/evaporation of water and 375 freezing/melting of ice on the surface of the balloon.

Figure 6 shows the O₃ concentrations measured on each ascent. Over the ~2 hours between the original ascent and the final ascent, O₃ in the layer between ~2.5 and ~5.1 km increased 4 – 12 ppbv, with a mean increase of 7.9 ± 4.7 ppb. Integrating the change in O₃ between the first and last profiles from 2.55 - 5.11 km (the range of the oscillation), and assuming uniform production in the volume of the storm ($5,300 \pm 2,400$ km² as indicated by the NPOL data from 0900 - 1700 UTC, assuming a 2 km depth), we find a potential of ~ 3.3×10^6 moles of O₃ (with ~50%uncertainty) created as part of this cell, a number that agrees reasonably well with the best esti-

-16-

mates from the lightning data (see Sections 3.1 and 3.3) and the OMI data (see Section 3.4 above). We note that the estimates from the lightning data and from OMI represent the total LO₃ throughout the cloud, whereas our sonde only observed the enhancements in the 2.5 - 5.1 km layer. The NPOL radar data, however, indicate that the average height of the 17 dB echo (used as a proxy for cloud height in this study) for the period 0900 – 1700 UTC was 5.24 ± 0.73 km, suggesting that the estimates are indeed comparable.

JIO 1

 $\vec{\nabla}[O_2]$

389 The total change in O_3 as observed by the balloon is given by

$$\frac{d[O_3]}{dt} = \frac{\partial[O_3]}{\partial t} + \vec{v} \cdot \vec{v}$$

where the first term represents in situ photochemical production (loss) and the second term represents changes due to advection. If the balloon had remained at a fixed altitude within the cloud or if the winds were constant with altitude over the vertical range of oscillation, we would assume that the observations were Lagrangian, meaning the advection term would vanish. Figure 7 shows the wind speeds and directions on each ascent of the balloon as determined from on board GPS data. Because of the vertical wind shear within the cell (seen by NPOL and in the balloon data), the advection term may be non-negligible, so we investigate further below.

Table 2 shows the calculated differences between the balloon positions recorded by the GPS and the estimated subsequent positions of the air masses sampled on the first ascent at three levels (2.75, 3.75, and 4.75 ± 0.15 km). To calculate these estimates, trajectories were based upon GPS wind speed and direction data vertically averaged in each of the three layers on successive ascents. The resulting *u* (east-west wind) and *v* (north-south wind) values were multiplied by the time difference between successive ascents to get longitude and latitude displacements. Since the balloon did not oscillate through all three layers each time, the table contains some "No data"

entries. These calculations suggest separations of 15 - 30 km between the originally and finally 405 406 sampled air masses, providing a constraint on the horizontal scale of potential O₃ gradients. 407 Figure 8 shows the change in O₃ with time as a function of altitude, calculated as the differ-408 ence between the O_3 at a given altitude as measured on each ascent with that measured on the 409 first ascent. Changes of 3 - 10 ppbv/hr are observed, with the highest rates at 2.5 - 3.0 km be-410 tween the first and second ascents, and at 3.5 - 4.5 km between the first and third ascents. We 411 note from Table 2 that for the former case, the balloon is located about 5 km upwind from the 412 original air mass, somewhat farther from the center of the storm as it comes ashore. For the lat-413 ter case, the balloon is located about 13 km downwind of the original air mass, closer to the cen-414 ter of the cell. Since both profiles suggest somewhat large O_3 changes with time, with one being 415 upwind and the other being downwind of the original air masses, it is possible that the advection 416 term over these spatial scales is relatively small.

417 One last component of the advection term to investigate is the vertical term. While some of 418 the change in O_3 with time is due to descent, most appears due to other factors. First, the original 419 ascending profile in Figure 6 (purple) can be joined with near perfect continuity to the final as-420 cending profile between 5 and 5.5 km, suggesting the O₃ enhancements are not due to air de-421 scending from above 5.5 km. Second, Figure 9 shows the change in potential temperature (theta) 422 between the first and final ascent as a function of altitude between 2.55 - 5.11 km. Between 2.55423 and \sim 3.40 km, 1.25 – 2.25 K of warming is observed. Between 3.4 – 5.11 km, 0.25 – 1.0 K of 424 warming is observed. If we assume that the change in the lower layer is due solely to 425 descent, the air found in the 2.55 - 3.40 km layer on the last pass was originally between 3.09 -426 3.55 km. The mean O_3 in this layer (defined by theta) on the first pass was 32.1 ± 1.3 ppb while 427 on the last pass, it was 37.9 ± 2.2 ppb, a difference of 5.8 ± 2.6 ppb. If we assume no descent at

428 all, the mean O_3 in this layer (defined by altitude) was 29.5 ± 1.9 ppbv on the first pass, a differ-429 ence of 8.4 ± 2.9 ppbv. Thus, we attribute 2.6 ± 3.9 ppbv (~31%) of the change to descent of air 430 between the first and last ascents, with air descending at an average rate of 6.2 cm/s. *Zahn et al.* 431 [2002] also found that downdrafts were insufficient to explain enhancements in O_3 and NO_x re-432 sulting from thunderstorms as seen by the NOAA WP-3D during the Southern Oxidant Study in 433 1996.

Finally, Figure 10 shows O_3 as a function of theta rather than height for the first and last ascents. Integrating the change in O_3 as a function of theta between 311.75 - 320.00 K, and assuming the cell size as before, we find an increase of 2.5×10^6 moles of O_3 , in good agreement with the estimates from the lightning data detailed above.

438 HYSPLIT back trajectories [Draxler and Rolph, 2003] of air parcels in each of the three lay-439 ers described in Table 2 (2750, 3750, and 4750 m) were calculated using the Global Data As-440 similation System (GDAS) meteorological fields. The course resolution of the GDAS meteorological fields (1^0 latitude $\times 1^0$ longitude) does not permit detailed analysis of the air mass 441 442 movement associated with the convective cell, nor do such analyses even approximately capture 443 the vertical motions associated with convection, so such trajectories must always be examined 444 with some caution. Nevertheless, the results indicate that the air masses in these layers of en-445 hancement O₃ remained within the convective cell as it developed over the Gulf of Panama and 446 moved ashore during the previous 6 hours (not shown). Thus, lightning production of NO and 447 the subsequent O₃ photochemistry as the cell dissipated and came ashore appears to be a plausi-448 ble explanation for the changes observed by the sonde.

449 **4. Summary and Discussion**

-19-

450 This work has presented a unique ozonesonde profile over Las Tablas, Panama on 5 August 451 2007. The balloon was launched on the southern side of a dissipating convective cell as it came 452 ashore from the east. Between 0900 and 1700 UTC, WWLLN data indicate 563 flashes (~2600 453 flashes accounting for the lightning detection efficiency of this network) in and around the Gulf 454 of Panama, while estimates of lightning flash rates using NPOL radar cloud height data result in 455 ~2300 flashes associated with this cell. The ozonesonde oscillated between ~2.5 and ~5.1 km 456 for ~108 minutes, during which 5 ascents were made through the air mass. Ozone concentra-457 tions increased by 4 - 12 ppb over a ~2 hour period between the first and last ascending profile, 458 consistent with the 5 ppb/day suggested near 5 km by Ott et al. [2007], the 10 ppb/day near 9 km 459 suggested by *DeCaria et al.* [2005], and the 7 - 8 ppb/day in the 8 - 12 km layer suggested by 460 Pickering et al. [1996] and Thompson et al. [1997]. The more rapid increase observed by the 461 sonde, however, may suggest that much of the O₃ production occurs soon after daylight returns 462 to the air mass affected by the cell rather than over the course of a day, a hypothesis consistent 463 with the laboratory work of Franzblau [1991].

464 Assuming the changes in O_3 observed by the sonde are representative of the entire cell (~5300 km² by 2 km depth), an increase of 3.3×10^6 moles of O₃ may be associated with this 465 storm. Further analysis indicates that descent alone explains only $\sim 30\%$ of the change in O₃ ob-466 467 served. Our ozonesonde LO₃ estimate agrees well with estimates from the WWLLN data ($\sim 2.2 \times$ 10^6 moles), from the NPOL data (~ 1.9×10^6 moles), and from the OMI LNO₂ data (~ 2.6×10^6 468 469 moles), in the last case provided we scale by the larger area considered with the WWLLN data 470 rather than the convective core area indicated by the NPOL data, and in all cases recognizing the 471 very large associated uncertainties.

472 Wind profiles from the sonde and NPOL data indicate some vertical divergence within the 473 layer of observed O₃ changes. Subsequent trajectory calculations suggest the possible separation 474 of the originally sampled air mass from that sampled on the final ascent by 15 - 30 km over the 475 ~108 minutes. The changes in O_3 observed by the sonde thus may be attributed to horizontal O_3 476 gradients with a scale < 30 km and/or in situ photochemistry. In the former case, the original air 477 mass at 2.75 km ends up ~15 km to the southwest of balloon trajectory (see Table 2), so the bal-478 loon ends up falling backward relative to the westward moving center of the cell and the original 479 air mass. As O₃ increases, the dynamical explanation would require higher O₃ concentrations 480 outside of the cell in less dense clouds in which photochemical O₃ production may take place and 481 in which loss of O_3 through reactions with NO from lightning may not have occurred. At 3.75 482 and 4.75 km, however, the trajectories suggest the original air mass ends up 25 - 30 km south-483 east of the balloon trajectory, resulting in the balloon sampling air closer to the center of the cell. 484 After the last oscillation, the balloon position is on the southeastern edge of the storm, whereas 485 for the first oscillation it was just west of the center of the cell. At these higher altitudes, it 486 would seem that the balloon is sampling air more reflective of the cell core, which suggests O_3 487 levels within the cell have actually increased.

If lightning and photochemistry are responsible for the changes in O_3 , our ozonesonde observations may be consistent with the idea that shortly after the lightning strikes, NO reacts with O_3 forming NO₂ and leading to O_3 loss within the clouds. This hypothesis is supported by the modeling studies of *Ott et al.* [2007] and *Salzmann et al.* [2008], the thunderstorm observations in *Ridley et al.* [2006], and the laboratory data of *Franzblau* [1991]. By the time the convective cell reaches the Panama coast, the lightning has subsided (as suggested by the WWLLN data in Table 1) and the clouds have begun to dissipate. Ozone within the clouds, therefore, may be relatively

-21-

depleted compared to its pre-storm values, and our balloon measurements may therefore simply
represent measurements at various stages of recovery as NO_x photochemistry begins to favor
production of O₃.

498 Alternatively or in addition to the loss/recovery process, new O_3 may have been produced 499 within the cloud, as suggested by Winterrath et al. [1999] who report a 62% increase in ozone 500 within the thunderstorm clouds; Shlanta and Moore [1972] who found ozone levels 2.6 times 501 higher at 6 km inside the cloud than pre-storm readings at the surface; and *Clark and Griffing* 502 [1985] who reported 250% increases downwind of thunderstorms near Baltimore in 1980. If the 503 dissipating cell was still producing lightning after launch, as suggested by the NPOL hourly flash 504 estimates (Table 2), it is possible that direct production of LO_3 was occurring within the cloud, as 505 suggested by Minschwaner et al. [2008]. 506 Given the large uncertainties in all of the LO₃ calculations, the agreement we find in the 507 study is reasonable, although we also note that it is possible that the area of the storm sampled by 508 the ozonesonde contained O_3 changes unrepresentative of the larger cell. In a future modeling 509 study, the unique balloon observations reported here will be combined with DC-8 data to further

- 510 investigate O₃ production and loss processes associated with deep tropical convection.
- 511

515 improving our manuscript.

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Version FINAL – 31 Aug. 2009

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807 Tables

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Hours (UTC)	WWLLN	NPOL Flashes
	Flashes	
0000 - 0800	9	0
0800 - 0900	4	0
0900 - 1000	36	9 (10)
1000 - 1100	61	0
1100 - 1200	34	44 (72)
1200 - 1300	174	240 (61)
1300 - 1400	160	450 (240)
1400 - 1500	92	670 (110)
1500 - 1600	6	600 (100)
1600 - 1700	0	240 (78)
1700 - 1800	1	11 (13)
1800 - 1900	2	1.7 (2.1)
1900 - 2000	9	4.3 (6.3)
2000 - 2400	150	108 (63)
Total 0900-1700	563	2300 (300)
TOTAL	738	2400 (310)

810

811 **Table 1.** Flashes detected by the World Wide Lightning Location Network and flashes calcu-

812 lated from cloud top heights estimated by the NPOL radar (uncertainties in parenthesis) near the

813 Gulf of Panama on 5 August 2007. See text for details of each.

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816

Trajectory Altitude	Time from	Balloon	Air mass	Balloon	Air mass	Separation	Direction
(km)	launch (s)	lat (deg)	lat (deg)	lon (deg)	lon (deg)	(km)	(deg)
	614	7.734	7.734	-80.269	-80.269	0.000	0
	2555	7.677	7.661	-80.337	-80.382	5.246	250
2.75 km	no data	no data	no data	no data	no data	no data	no data
	4629	7.660	7.597	-80.418	-80.492	10.816	230
	6208	7.654	7.550	-80.463	-80.551	15.103	220
	859	7.724	7.724	-80.277	-80.277	0.000	0
	2766	7.674	7.683	-80.346	-80.262	9.320	84
3.75 km	3459	7.671	7.657	-80.374	-80.260	12.672	97
	5010	7.654	7.605	-80.431	-80.261	19.494	106
	6535	7.655	7.565	-80.469	-80.235	27.614	111
	1134	7.722	7.722	-80.284	-80.284	0.000	0
	no data	no data	no data	no data	no data	no data	no data
4.75 km	3956	7.667	7.635	-80.393	-80.307	10.116	111
	5335	7.658	7.582	-80.440	-80.300	17.591	118
	6806	7.659	7.544	-80.476	-80.269	26.130	119

817

818 **Table 2.** Balloon trajectory and estimated air mass trajectories at three different altitudes within

819 the 2.5 - 5.1 km layer in which the balloon oscillated. "No data" are reported when the balloon

did not oscillate through the level of the air mass trajectory calculation. The "Direction" column

821 indicates the compass heading from the balloon location to the estimated air mass location.

822 Color coding matches that for the balloon data shown in Figures 12 - 15. See text for details.

823

825 Figure Captions826

- Figure 1. A comparison between the ozonesonde readings and NATIVE surface O_3 measurements at the time of launch. Agreement is good to within ~5% during the TC-4 campaign.
- Figure 2. Low-level (0.7 elevation) PPI images of convection observed east of NPOL at 1311
 UTC 05 Aug 2007. The color scale has units of radar reflectivity (dBZ).
- **Figure 3.** Hourly mean surface O_3 data from 13 July 9 August 2007 (black dots) and mean \pm
- 832 one standard deviation (dashed) recorded on NATIVE. The gray dots are the data from 5
- 833 August. During the period of the arrival of the convective cell (13 18 UTC), O₃ is decreased 834 relative to the mean values.
- Figure 4. Gridded $(1^0 \times 1^0)$ OMI lightning NO_x in the Panama area on 5 August 2007. (a) Shows the larger region used for the first estimate, while (b) shows the smaller region in the Gulf of Panama used for the second estimate (see text).
- **Figure 5.** The altitude vs. time of the ozonesonde flight on 5 August 2007 from Las Tablas,
- Panama shows the balloon oscillating 5 times between ~2.5 and ~5.0 km. The color coding of
- 840 each ascent will be used in successive plots to identify changes with time of other measured pa-841 rameters.
- Figure 6. Ozone vs. altitude on the ascents as the balloon oscillated on the 5 August 2007 flight,
 with color coding to match the ascents identified in Fig. 3.5.1.
- **Figure 7.** Wind speed (thick) and direction (thin) as determined from the ozonesonde GPS data.
- 845 The balloon moved through a region of vertical shear, resulting in the separation of the balloon
- trajectory from the trajectories of the air masses the sonde sampled. See text and Table 2 for fur-ther details.
- Figure 8. Calculated dO₃/dt vs. altitude for the ascents of the flight on the 5 August 2007 flight,
 with color coding to match the ascents identified in Fig. 3.5.1. See text for details.
- **Figure 9.** The change in potential temperature from the first to the last ascent vs. altitude for the flight of 5 August 2007.
- Figure 10. Ozone vs. potential temperature for the first and last ascents of the flight of 5 August2007 shows that the changes cannot be due to descent of the air mass alone.
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- 856 MORRIS ET AL.: OZONE PRODUCTION IN A CONVECTIVE CELL
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858



861 **Figure 2**



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13 July - 9 August 2007

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Figure 4

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873 Figure 5874





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